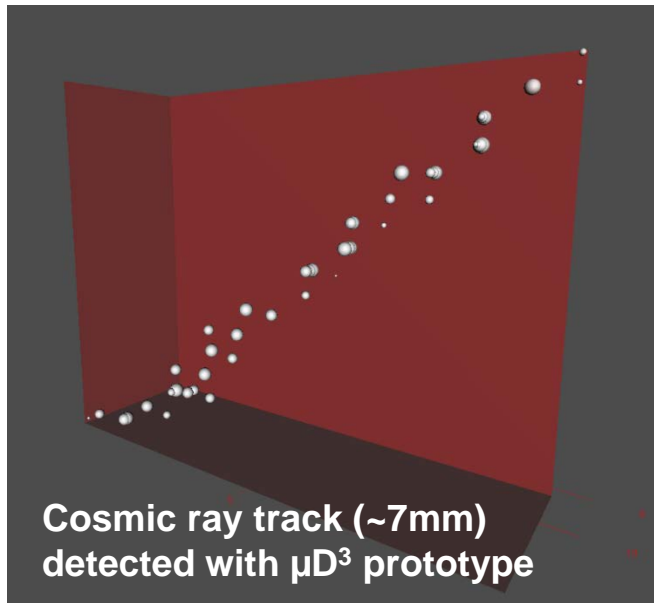


Recent Progress on D³ - The Directional Dark Matter Detector



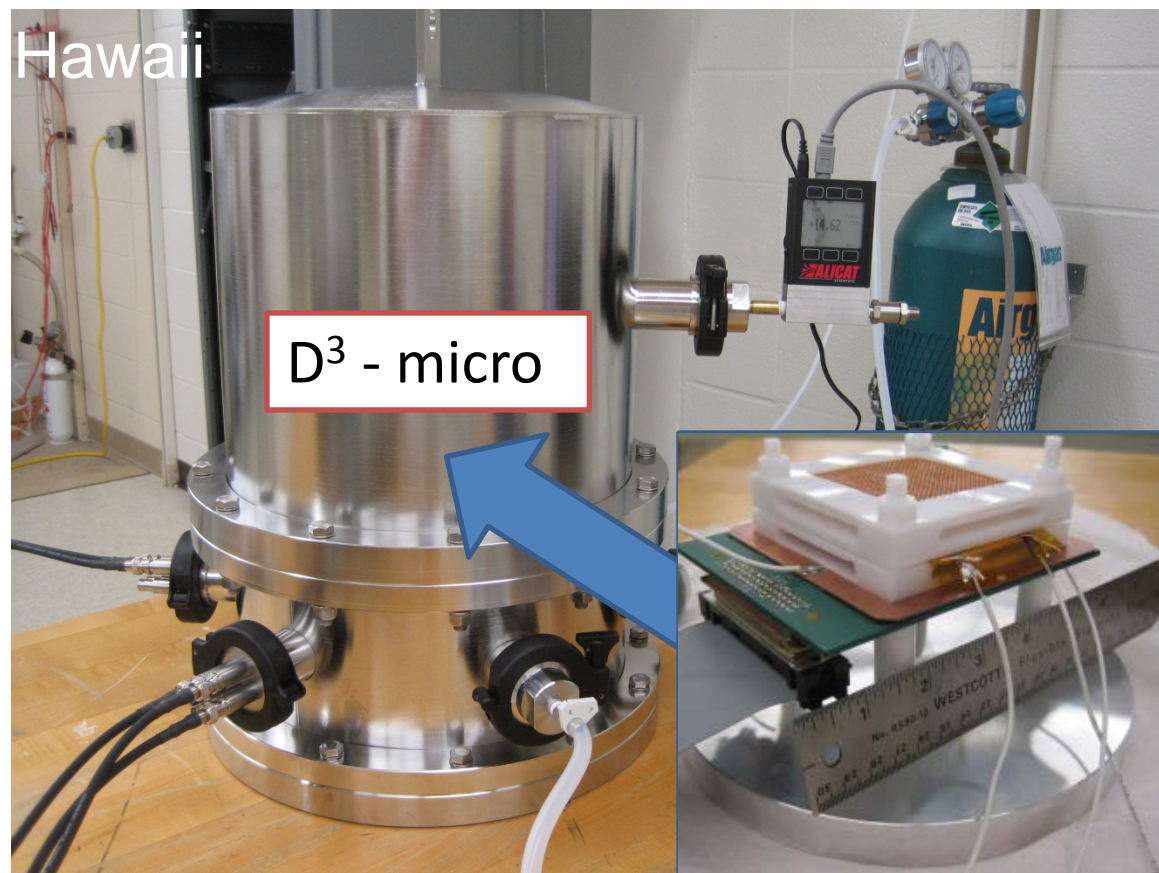
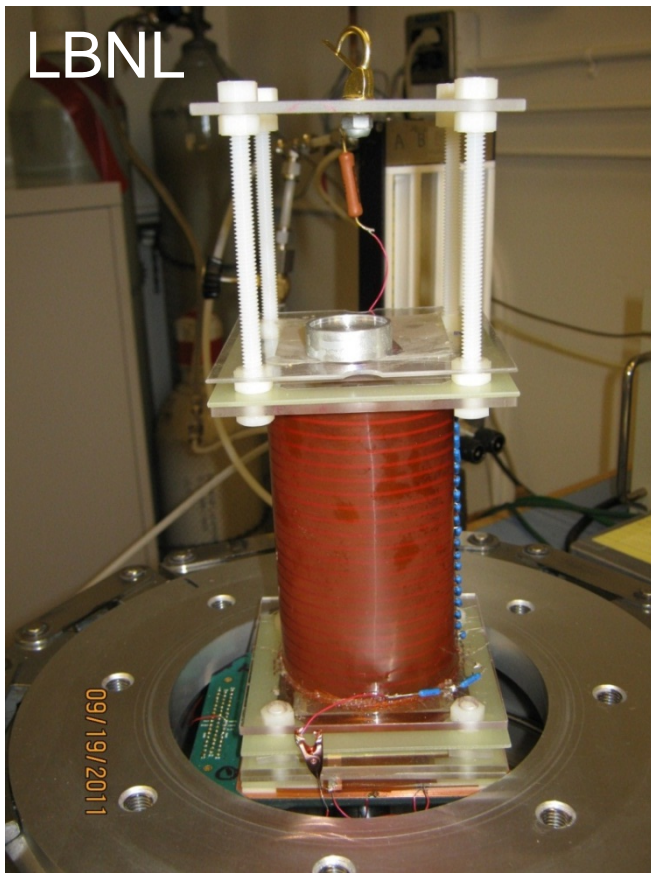
- Motivation
- Detection Principle
- Performance of prototypes
- Plans for future
- Related activities

Sven E. Vahsen, University of Hawaii



D³ - Directional Dark Matter Detector

- Investigating feasibility of directional DM search w/ micro-pattern gas detectors
- Technology also of interest for detecting neutrons and charged particles
- Small (1-10 cm³) prototypes built at LBNL and U. Hawaii
- Ongoing since ~Fall 2010



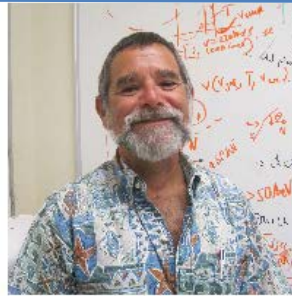
Team at U. Hawaii and Berkeley Lab



Igal Jaegle
Postdoc



Jared Yamaoka
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Marc Rosen
Mechanical Engineer



John Kadyk



Maurice Garcia-Sciveres



Michael Hedges
Graduate Student



Steven Ross
Graduate Student



Thomas Thorpe
Graduate Student



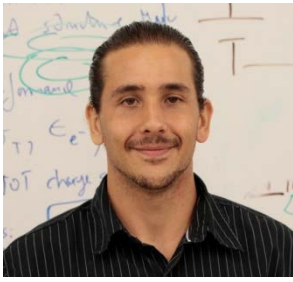
Mayra Lopez-Thibodeaux
(UC Berkeley Student)



Kelsey Oliver-Mallory
(UC Berkeley Student)



Ilsoo Seong
Graduate Student



Kamaluawaiku Beamer
Undergraduate Student



Sven E. Vahsen



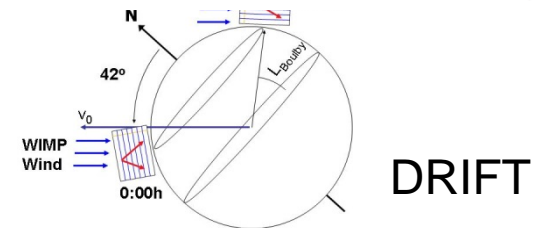
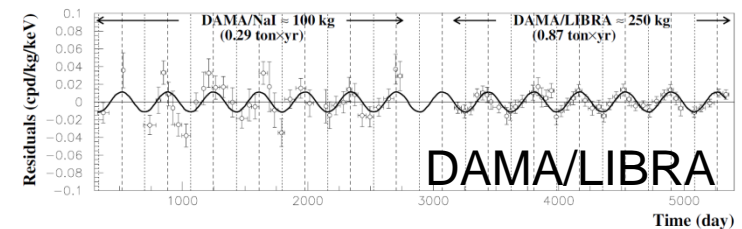
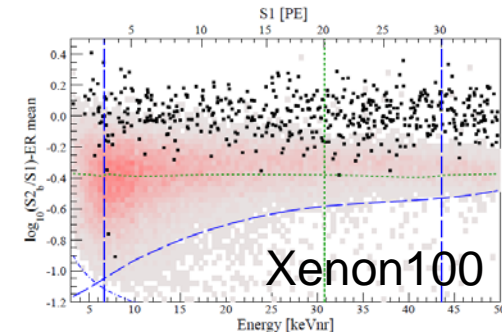
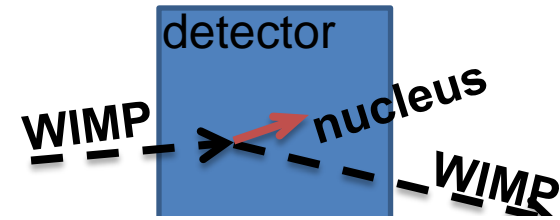
UNIVERSITY
of HAWAII[®]
MĀNOA



Motivation for Directional WIMP Search

Three possible signatures of direct DM detection:

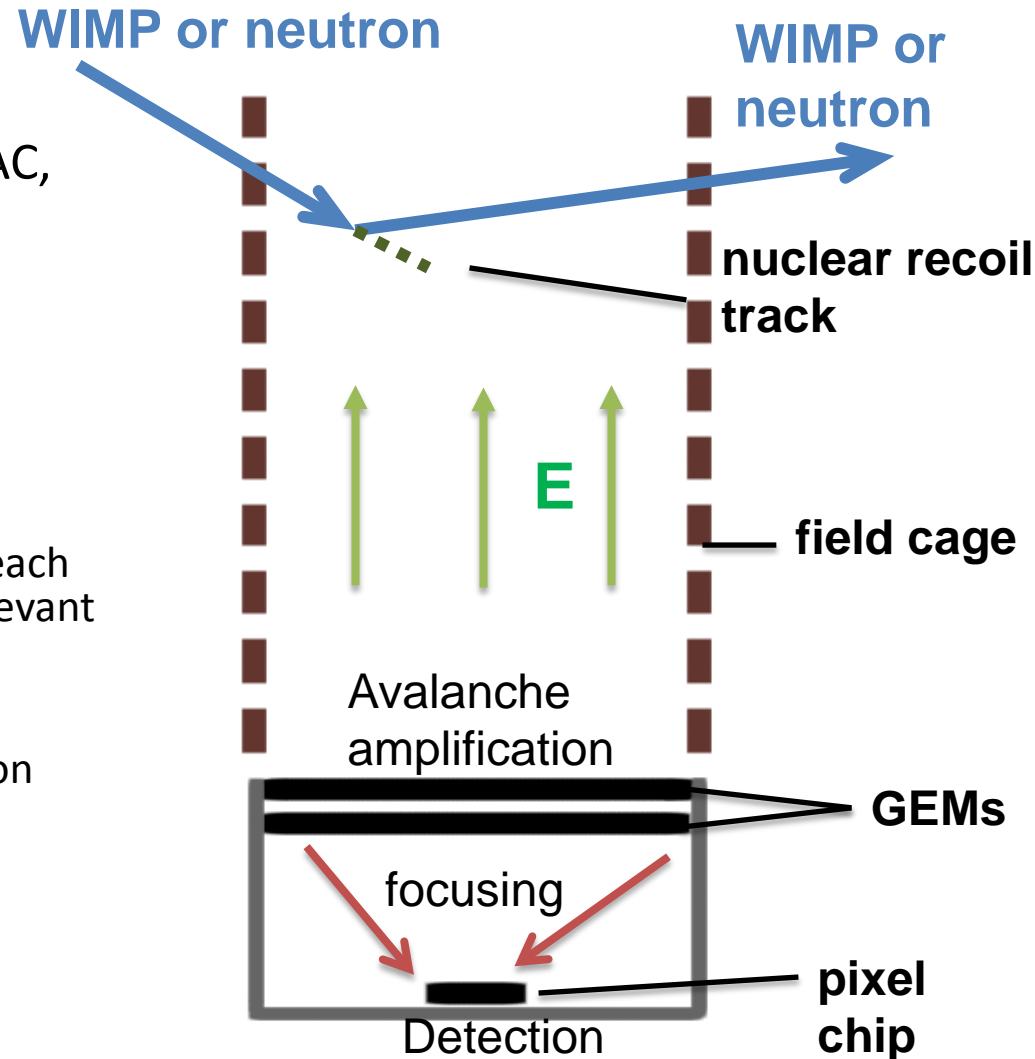
- 1. Excess count rate over predicted BG**
 - Requires ultra-clean detectors & precise understanding of remaining backgrounds
 - neutrons produce identical events
- 2. Annual modulation of count rate**
 - *due to motion of earth around sun*
 - expect %-level effect → requires thousands of signal events & < %-level control of BGs
- 3. Daily oscillation in mean recoil direction**
 - *due to rotation of earth*
 - expect large effect, only ~10 events required
 - no known background with this signature



- Detector with directional sensitivity could provide unambiguous evidence for WIMPs
- Post-WIMP-discovery, a (huge) directional detector could perform WIMP astronomy
- A *small*, m³-scale, directional detector with *low energy* threshold would be relevant *today*
 - could investigate the recent hints for low-mass (~10-GeV) WIMPS

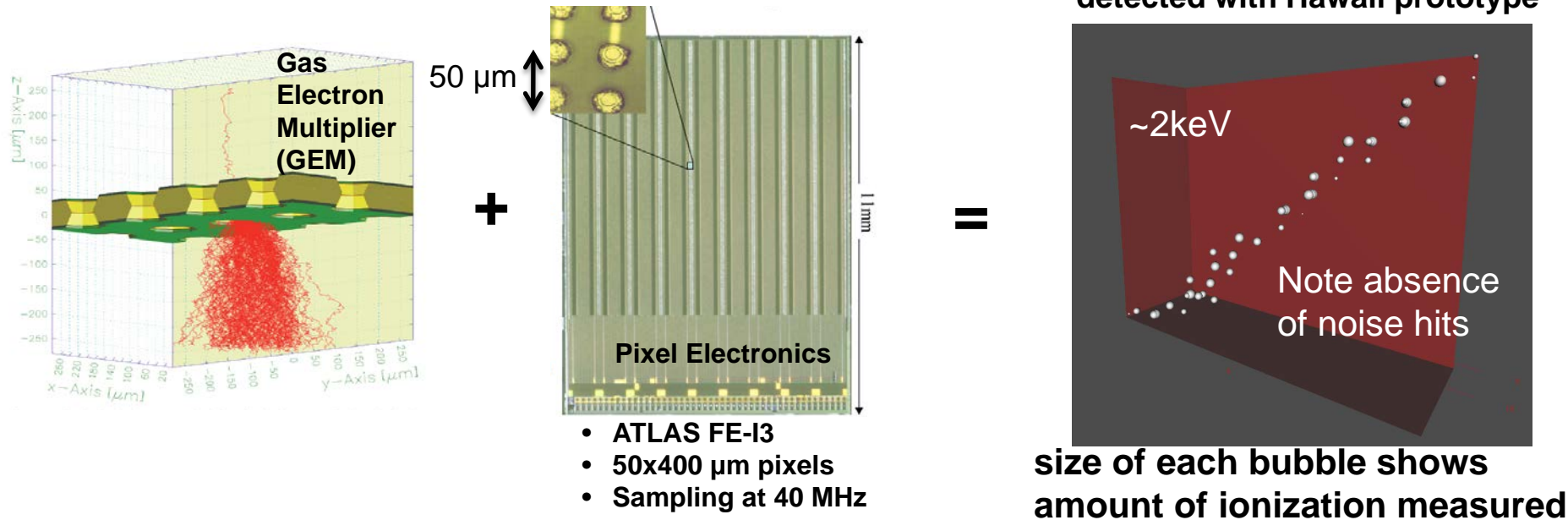
Directional Recoil Detection in Gas TPCs

- Several efforts based on low-pressure gas time projection chambers: DRIFT, DMTPC, MIMAC, NEWAGE, D³
- Benefits
 - Directional sensitivity
 - Can ID recoiling particle
 - Easy to change target nucleus
- Drawback
 - low target density → harder to reach large target mass required for relevant sensitivity
- D³
 - Charge amplified with Gas Electron Multipliers (GEMs)
 - Charge detected with pixel chip
 - Charge focusing - potential for significant cost reduction of large detectors



Charge Amplification and Detection in D³

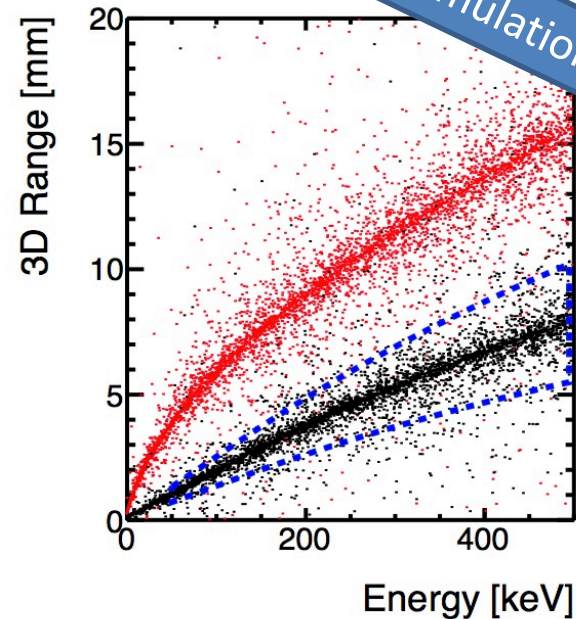
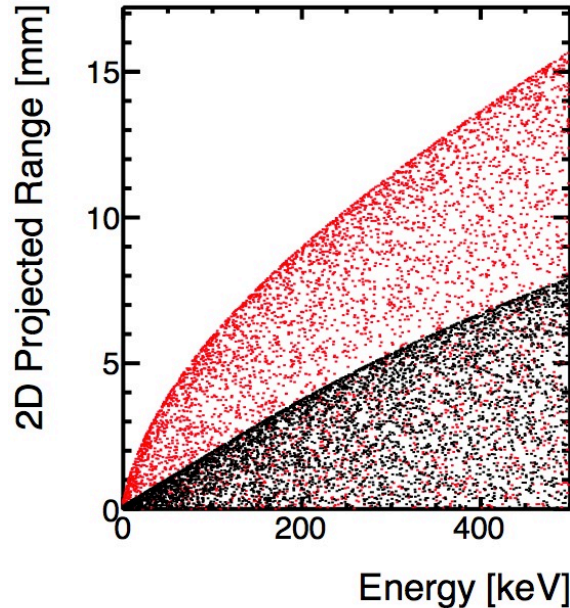
- Drift charge amplified with double layer of GEMs - gain $\sim 20k$ at 1 atm
- Detected with pixel electronics - threshold $\sim 2k e^-$, noise $\sim 100 e^-$



Advantages of this approach

- Full 3D tracking w/ ionization measurement for each space-point (head/tail sensitivity)
→ **improved WIMP sensitivity and rejection of alpha particle backgrounds**
- Pixels ultra-low noise (~ 100 electrons), self-triggering, and zero suppressed
→ **virtually noise free at room temperature** → low demands on DAQ
- High-single electron efficiency → **suitable for low-mass WIMP search**

Advantage of 3D Tracking



(simulation by James Battatt)

Simulation of the range vs. energy profiles for alpha particles (red) and fluorine recoils (black) in 75 Torr CF_4 .

Left. With only 2D range reconstruction, a degeneracy exists between steep-angle alphas and nuclear recoils of the same energy.

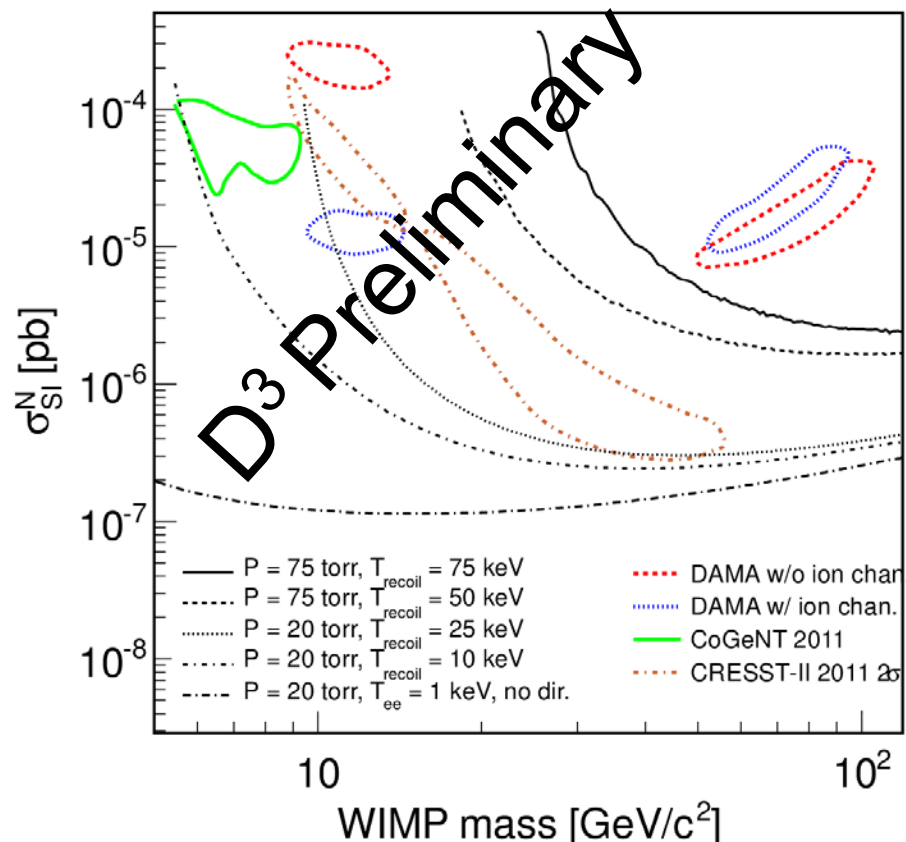
Right. With 3D tracking, the alpha and fluorine recoil bands separate. In this simulation, the angular resolution was 5° . The blue dashed lines represent a cut above 50 keV that achieves a 10^2 alpha rejection with an 86% fluorine recoil acceptance.

3D tracking: much improved rejection of alpha-particle backgrounds

Advantage of Low Track Energy Threshold

- Preliminary evaluation: 3-m³ detector could achieve *directional sensitivity* to (controversial) ~10 GeV WIMPS
- Golden scenario - if DAMA/LIBRA were due to WIMPs:
 - observe ~1000s of non-directional events (can observe yearly oscillation)
 - use subset (~100) of these to search for daily directional oscillation, to determine if BG or WIMPs

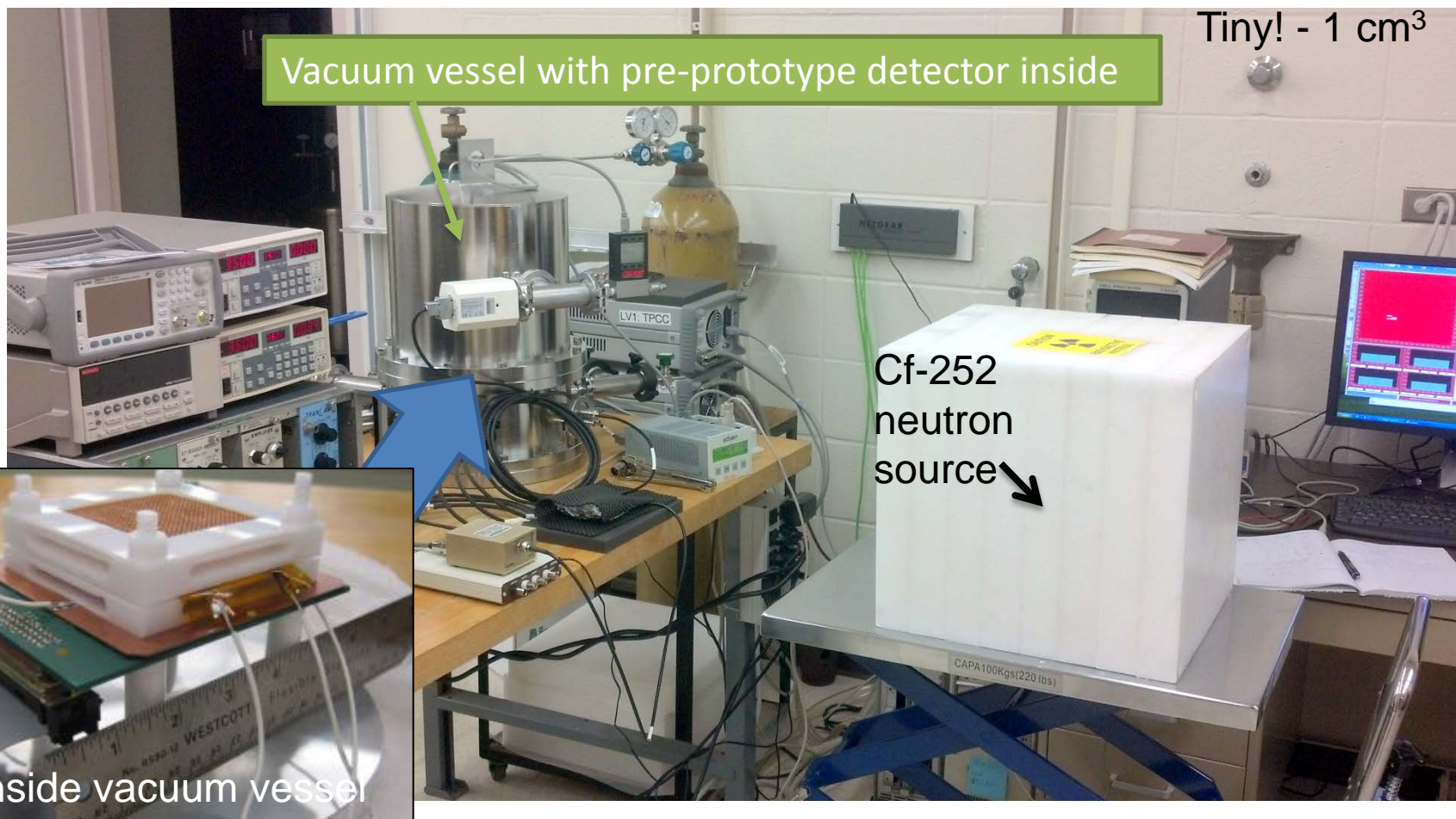
<http://arxiv.org/abs/1110.3401>



Estimated sensitivity to spin-independent WIMP-nucleon scattering, 3-m³ directional dark matter detector, running for 3 years with 33 cm drift length and CF₄ gas, for four different track reconstruction thresholds and for non-directional analysis.

Track energy threshold as low as 10 keV crucial for detecting 10 GeV WIMPS!

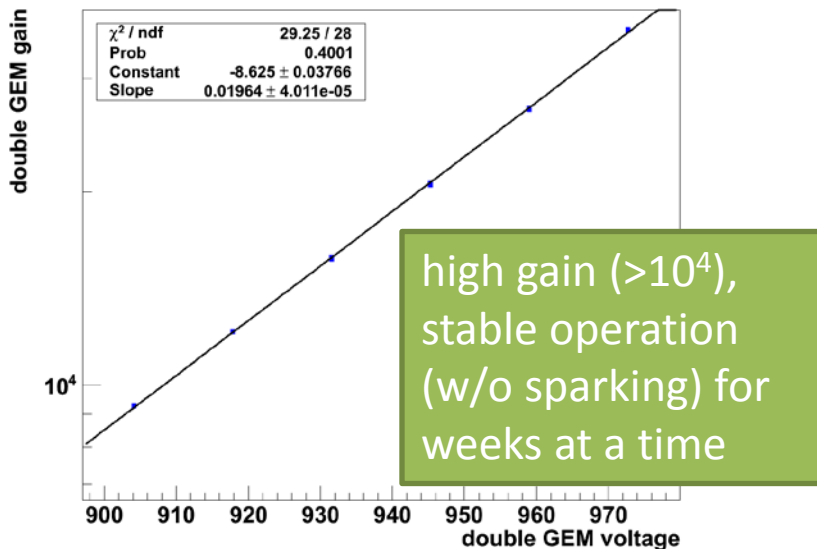
Characterization of Current Prototype - $D^3\mu$



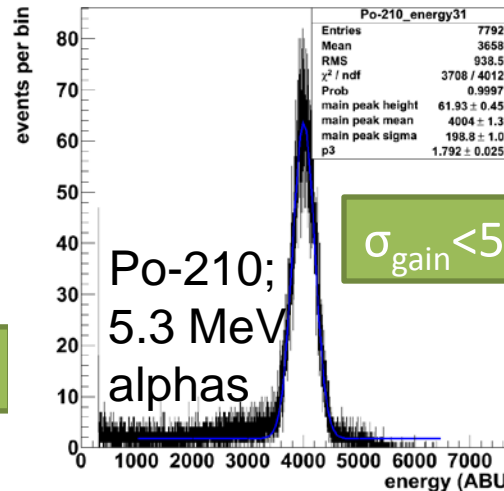
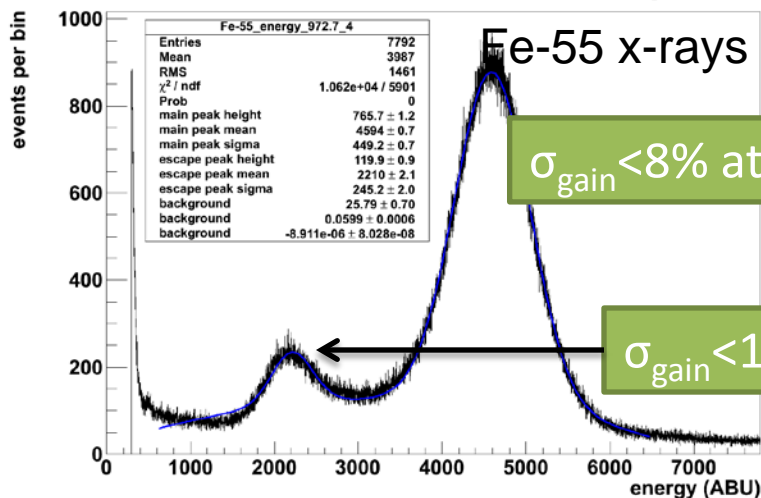
- Stable operation for >1 year, large datasets recorded
 1. commissioning w/ ArCO₂; muons, x-rays, α -particles ('11,'12)
 2. detailed calibration & directional neutron detection w/ HeCO₂ (Fall '12-now)
- Low-pressure operation w/ CF₄, WIMP search surface-run starting this summer

Gain and Gain Resolution of Double GEM

gain vs. GEM voltage

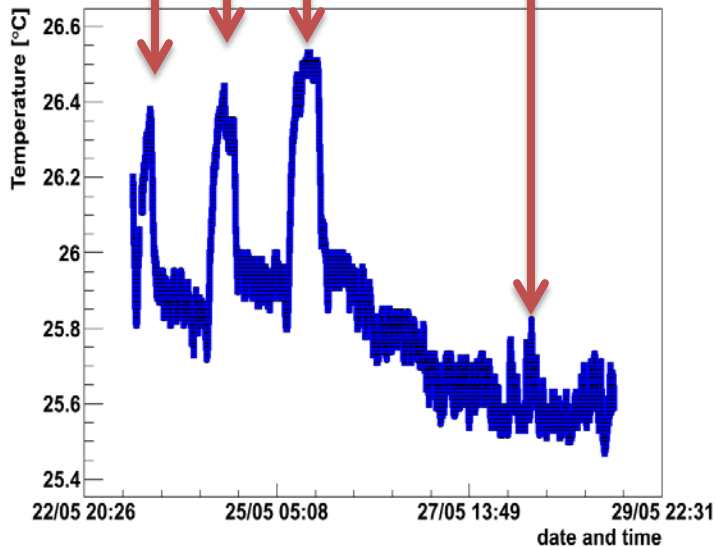
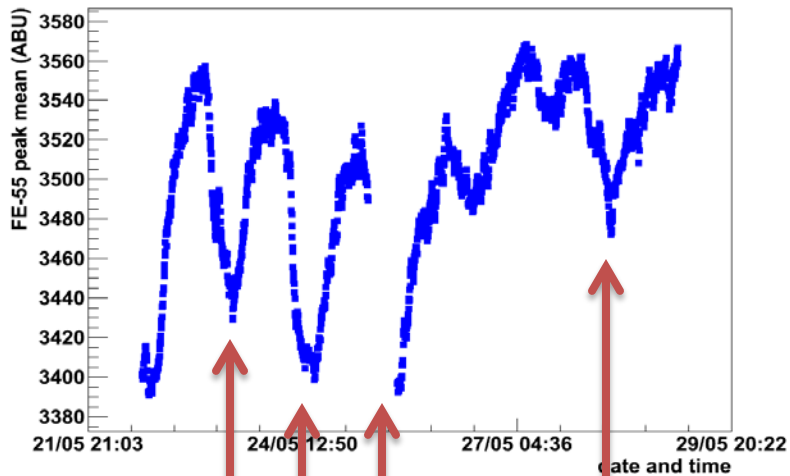


- ArCO₂ at p=1atm
- (HeCO₂ results, not shown, are slight *better*)



- Sufficient gain to achieve single-electron sensitivity if needed
- Good gain resolution for MeV-scale signals, adequate even for few-keV signals!

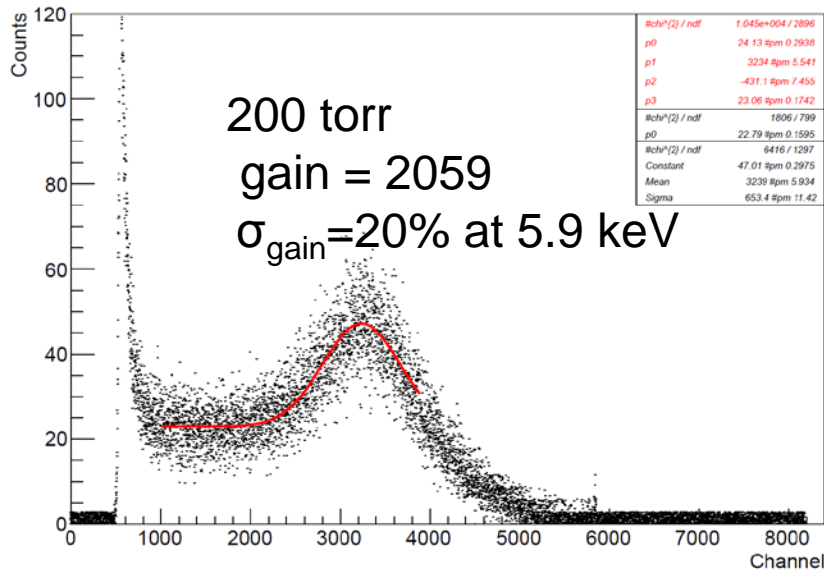
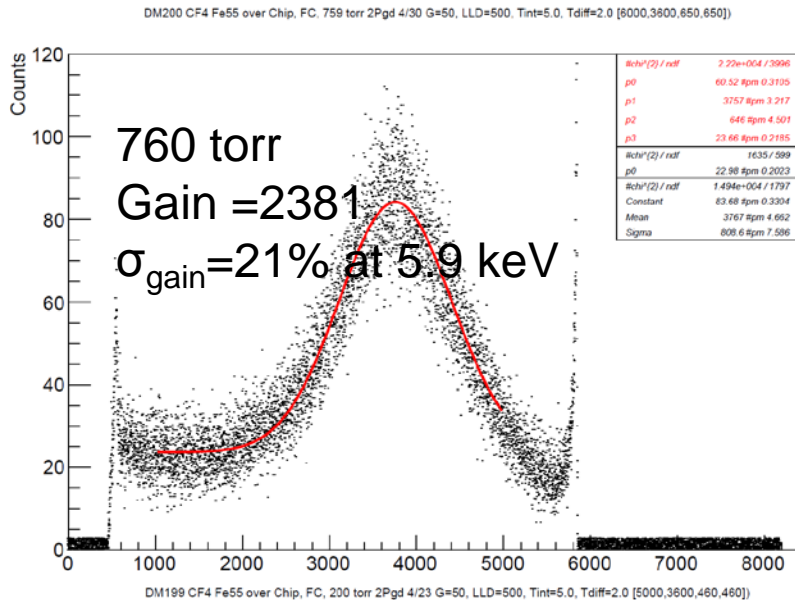
Gain Stability



- Measured gain continuously for 5 days, to test for possible gain degradation due to decreasing gas purity
- Not observed (=good!)
- Instead observed +/- 2% gain variation tightly correlated with lab temperature; guessing this is due to NIM electronics

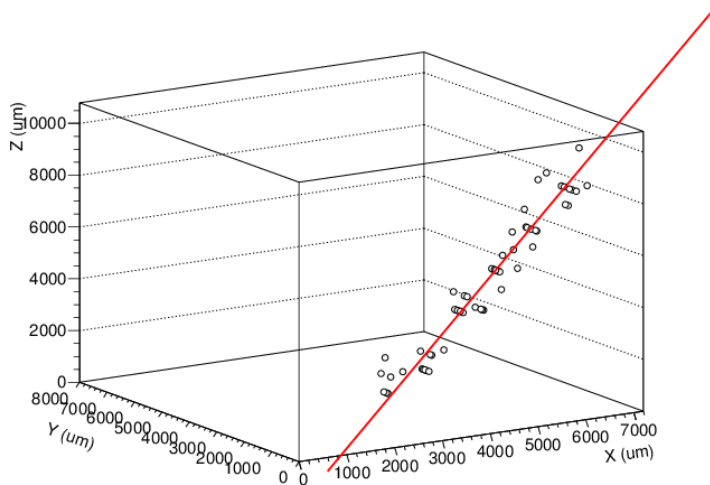
Excellent gain stability without flowing gas

Gain Resolution w/ CF₄ @ Berkeley Lab

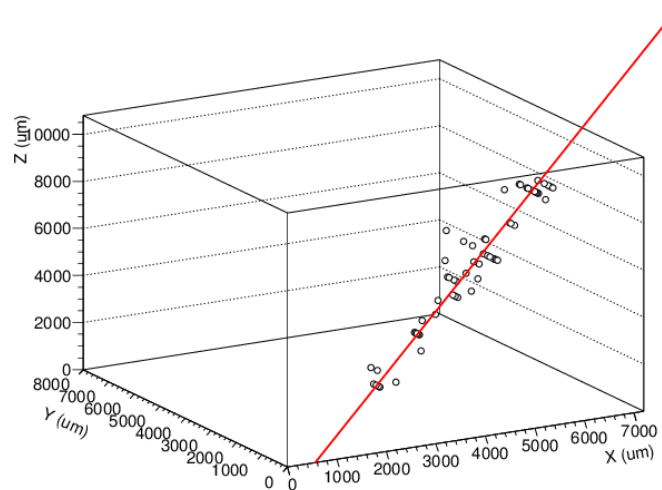
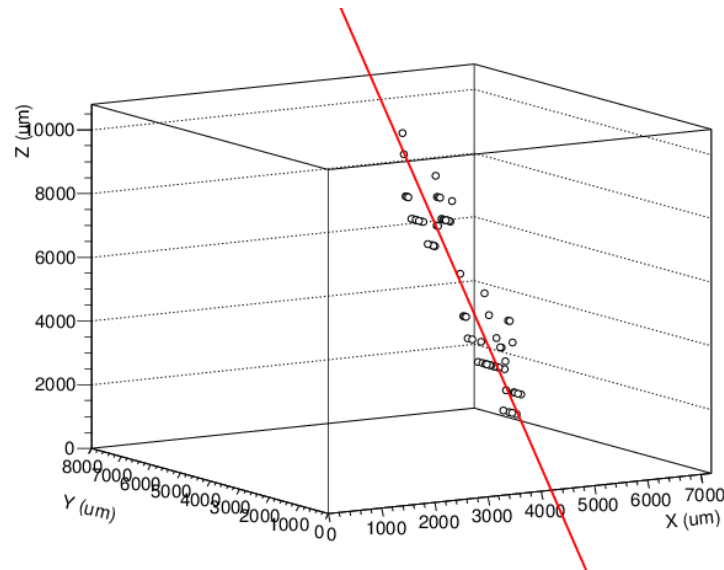


- Detectors work well with CF₄
- May need 3rd GEM to get gain > 1000 at low pressure
- Figures
 - Fe-55 x-rays
 - 12 cm drift in field cage
 - Pulseheight analyzer

3D Point resolution



ArCO₂ at p=1 atm

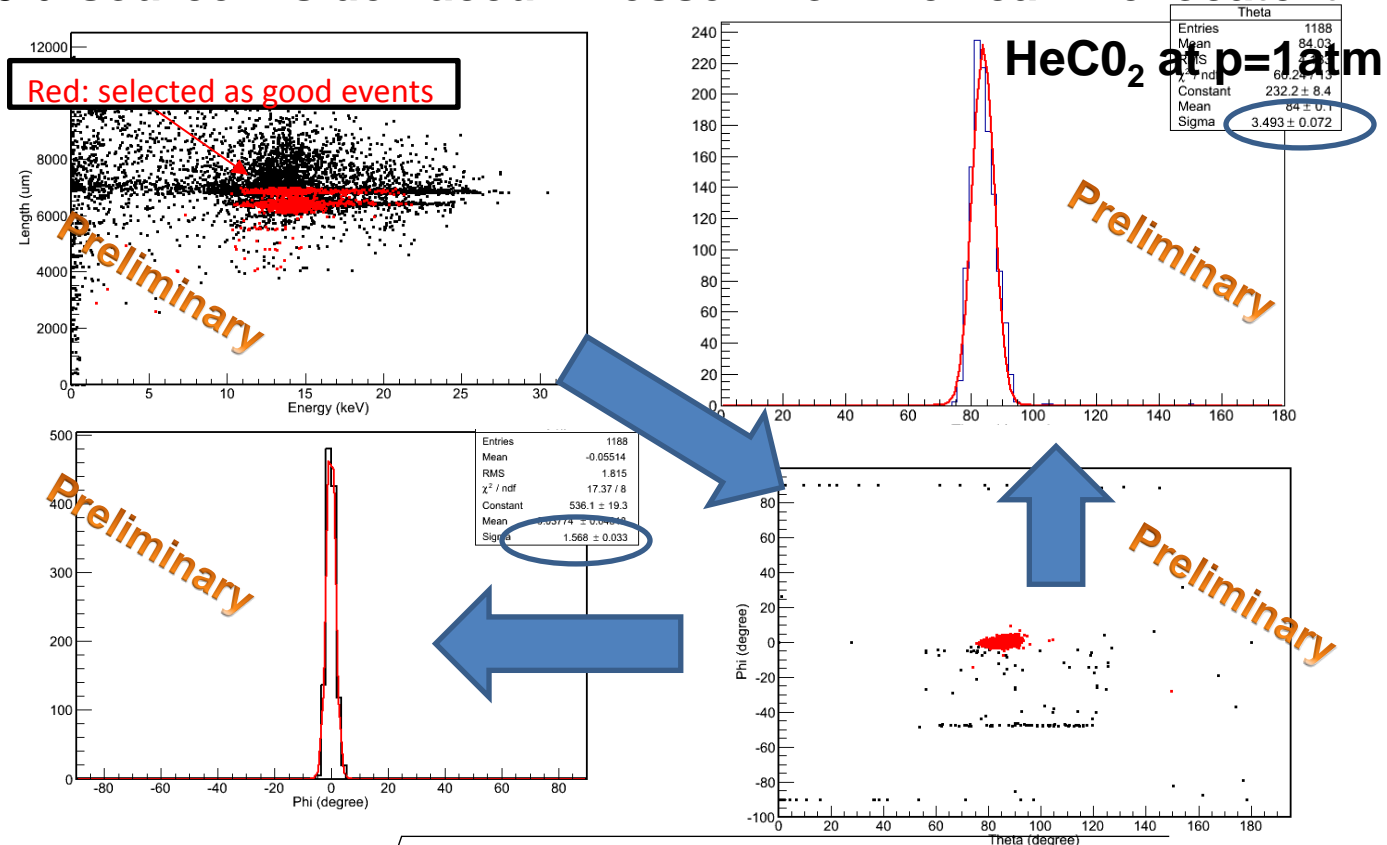


- > 10k cosmic events recorded.
- Use such events to measure detector point resolution ($\leq 200 \mu\text{m}$)

Based on measured point resolution, expect angular resolution on nuclear recoils ~ 1 degree

Angular and Energy Resolution, nuclear recoils

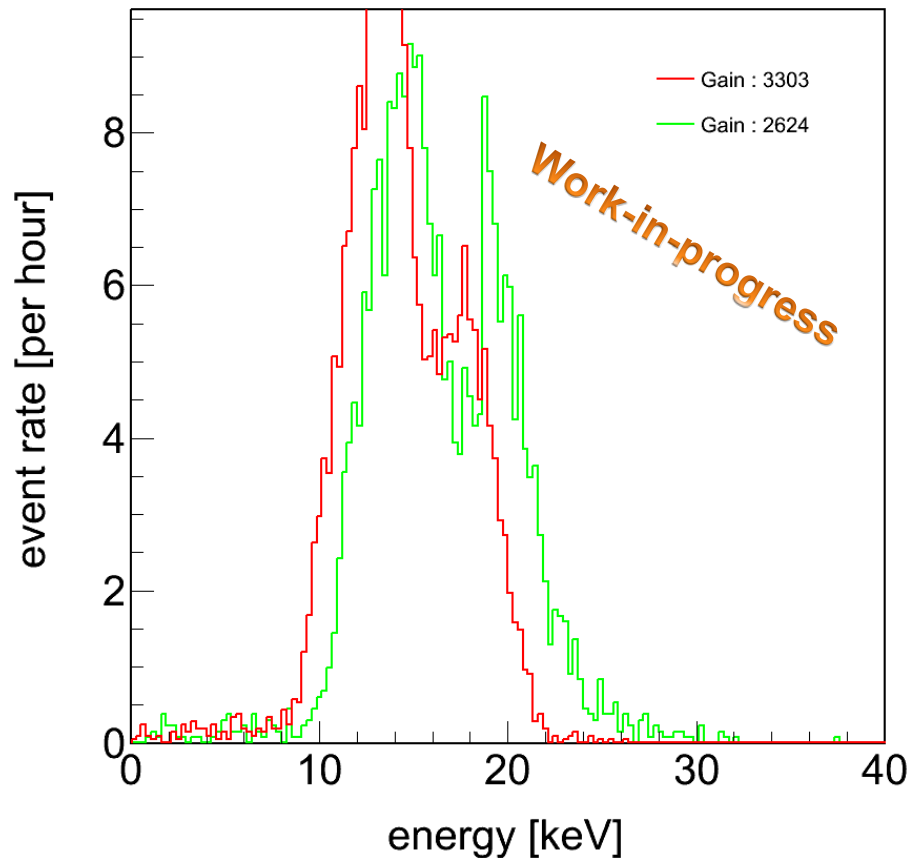
Po-210 α -source inside vacuum vessel. How well can we locate it?



$$\sigma_{angle} = \sqrt{\sigma_{detector}^2 + \sigma_{stragglng}^2 + \sigma_{source\ cone}^2 + \sigma_{source\ size}^2}$$

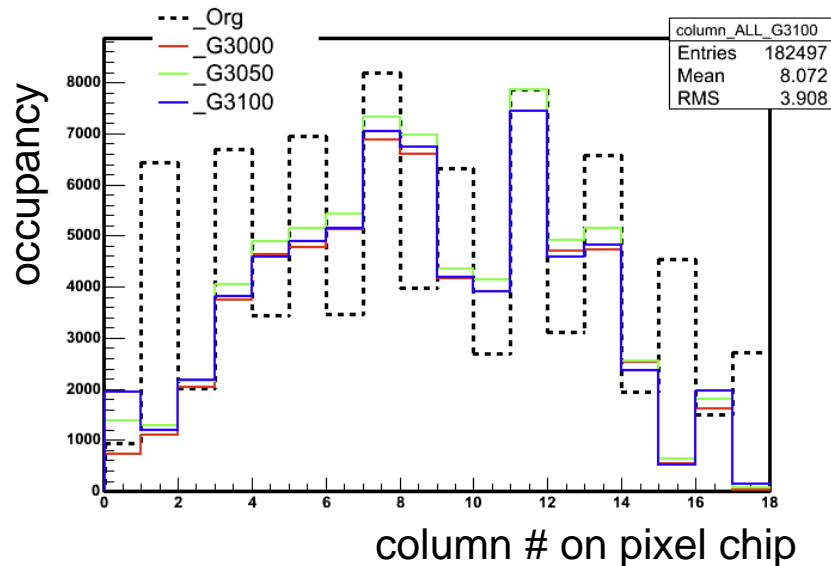
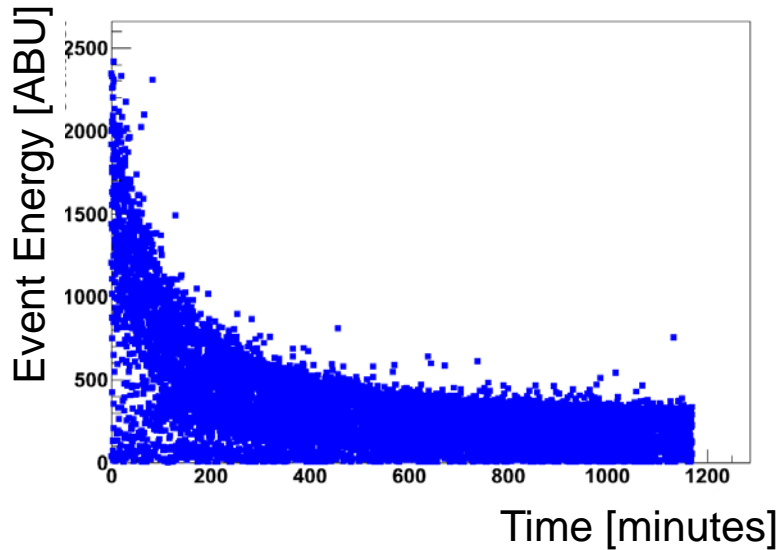
- Selected events clearly point back to a single source
- No BG after good-track selection
- consistent with $\sigma_{\phi, \theta\ detector} \leq 1^\circ$

Energy Resolution - Surprises



- Energy resolution significantly worse than gain resolution when measured over entire pixel chip area
- Surprising, as both GEM gain and pixel chip calibration measured independently to be uniform ($<5\%$) and stable in time ($<2\%$)
- If we restrict only to small region of chip, energy resolution approaches $\sim 10\%$ as expected (not shown)

Energy VS Time and Position



- ...More detailed investigation revealed: even though GEM gain and pixel calibration are stable & uniform, effective gain is time and position dependent
- Hypothesis: charge-up of pixel chip surface distorting E-fields and affecting **charge collection efficiency**
- *Supporting evidence:*
 - Higher gain \rightarrow faster gain reduction
 - Gain recovers when E-field turned off

Studying Time /Position Dependence I

rectangular aluminum pads
deposited on top of chip,
grounded during operation

ATLAS FE-I3

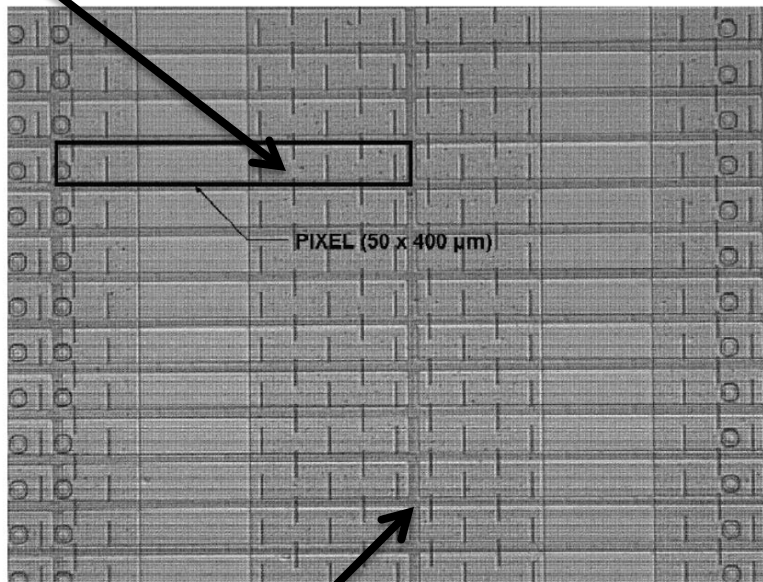
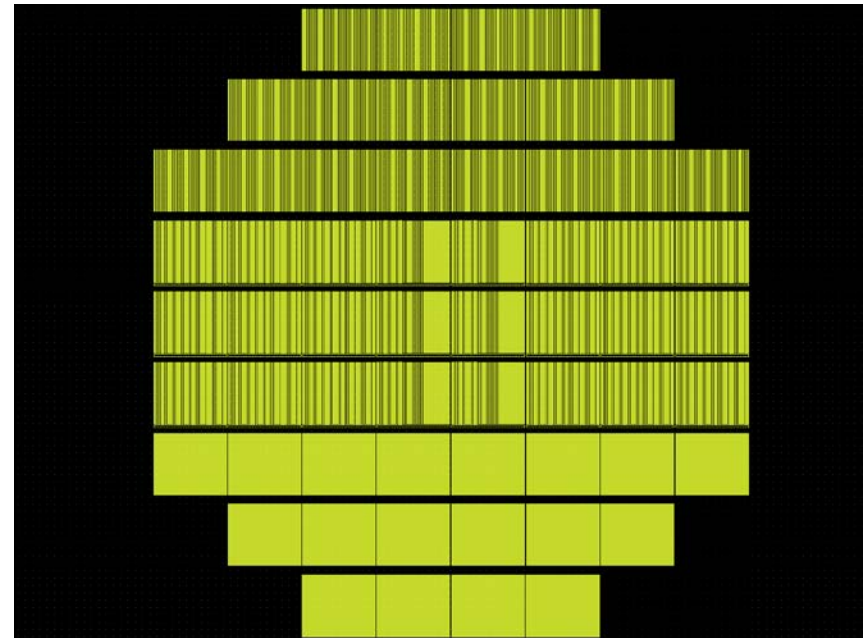


Fig. 6. Microphotograph of the surface of the ATLAS FEI3 chip after deposition of gold. One of the 50 x 400-micron cells is outlined. The entire chip, containing 2880 pixel cells, is 7.2 mm by 10.8 mm in surface dimensions and is 700 microns in thickness.

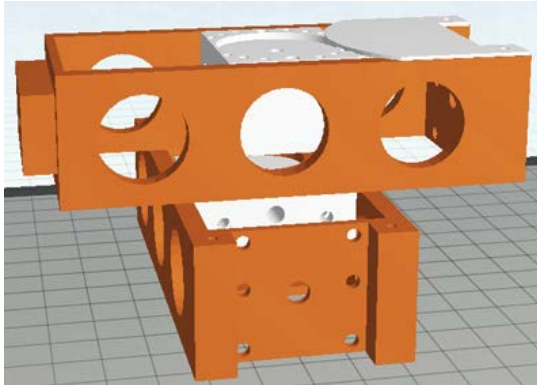
FE-I4: depositing a variety of metal
pad shapes to study effect on effective gain
(see backup slides)

ATLAS FE-I4 Wafer

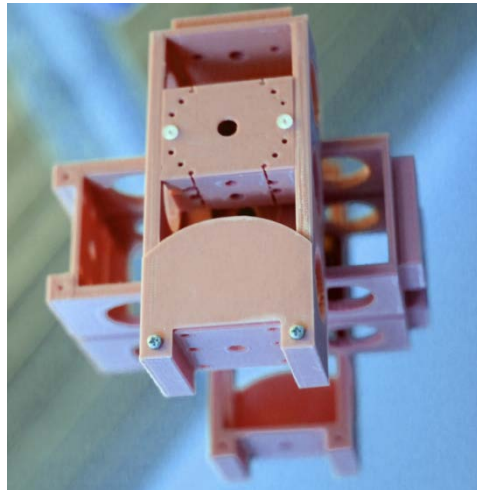


SiOxide between pad is insulating. Charging up at high gains & rates?
→ may explain both position and time-dependence

Studying Time /Position Dependence II

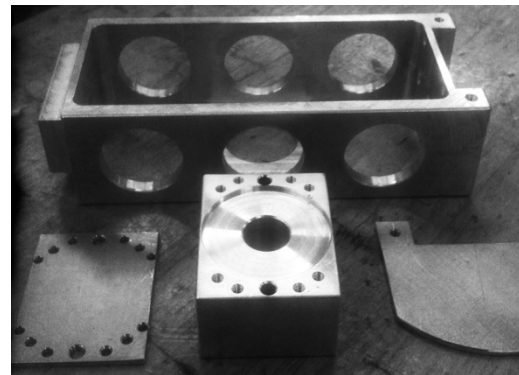


CAD design



3D-printed model

- Undergrad student designed 2D-motion stage for scanning collimated Fe-55 calibration source across chip
- Will allow us to measure position and time dependence of energy scale versus metal pad shape

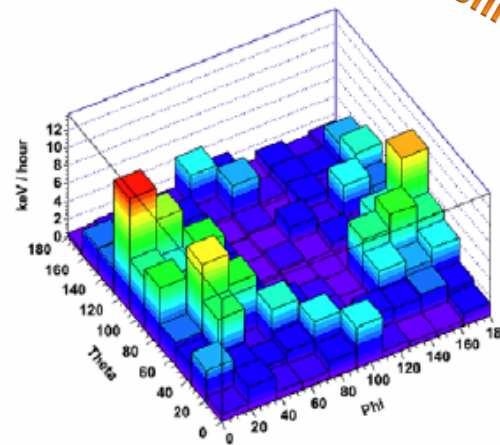
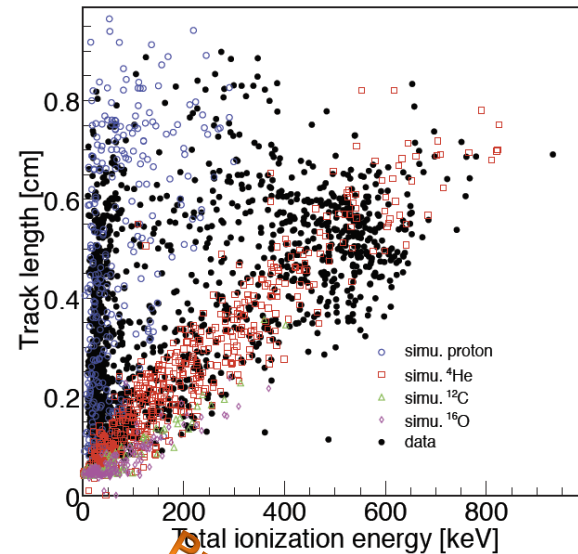


machined, final aluminum parts

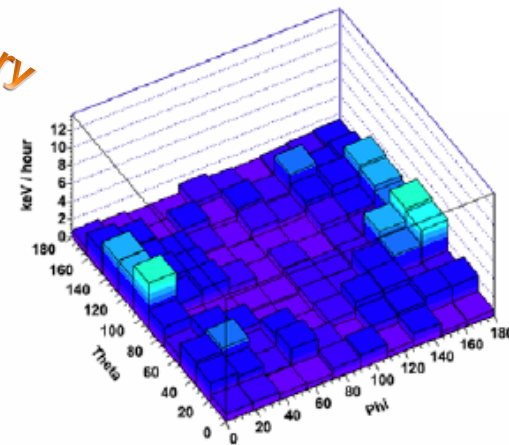
Directional Neutron Detection

HeCO₂ at p=1atm

- **Cf-252 neutron source pointed at vacuum vessel. Can we locate it?**
- Rough agreement with simulation
- Expect broad recoil-angle distribution
- When source present, observe increased energy-flux in expected direction ($\Theta=90$, $\phi=20-30$ degrees)
- Encouraging, but analysis still ongoing
- We have already recorded a number of additional datasets with source at different angles, see if we can track it.



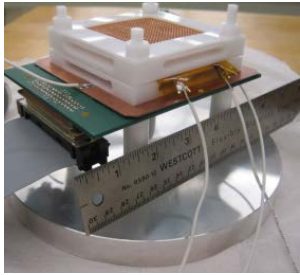
(a) Source ON



(b) Source OFF

Next Generation Detectors

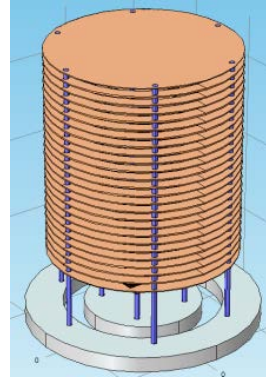
Built, stable
operation for > 1 year



μD^3 (1cm^3)
1 pixel chip



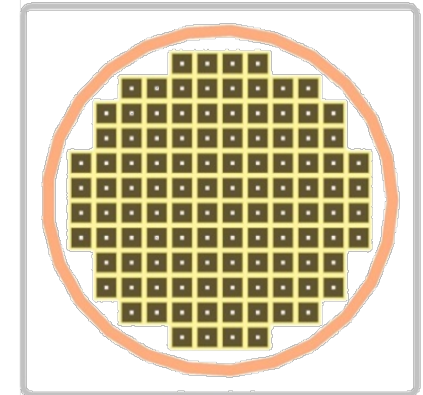
building this year



mD^3 (~ 10 liters)
4 pixel chips



planned



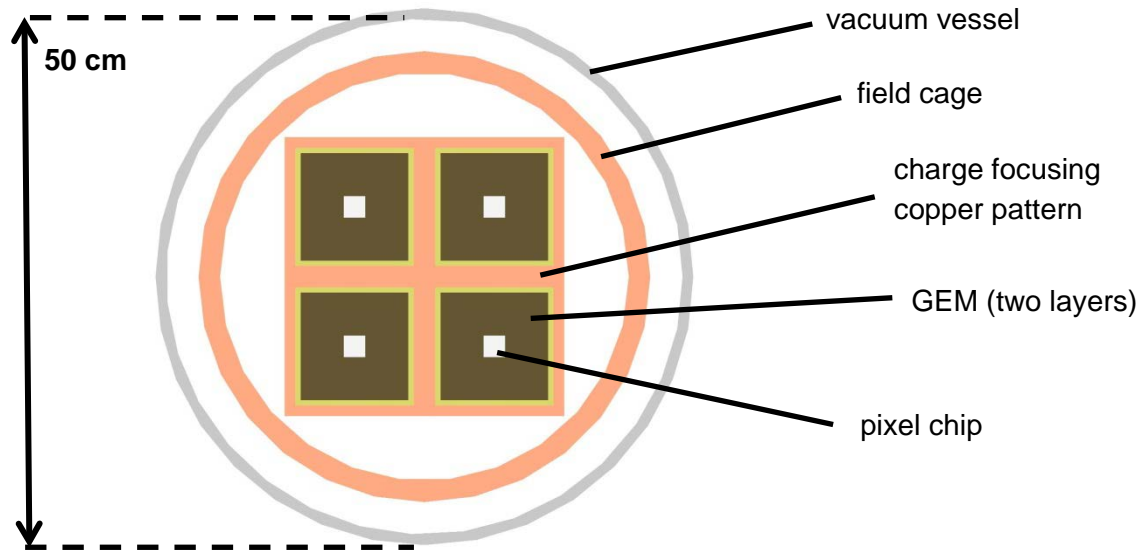
D^3 ($\sim 1\text{m}^3$)
 ~ 400 pixel chips

Ingredient

1. larger pixel chips
2. electrostatic focusing of drift charge
3. existing ATLAS DAQ
4. negative ion drift

Next Prototype, 2013: D³-milli

- Prototype dedicated to studying next generation pixel electronics, trigger, charge focusing
- 10x10 cm GEMs (CERN), 2x2cm Pixel Chip (ATLAS-FE14), SEABAS DAQ System from KEK



Top-view of the 12-liter prototype, which implements four unit cells inside a common field cage. The shown geometry assumes a charge focusing factor of 1.2 before the GEMs, and a charge focusing factor of 5.0 between the GEMs and pixel chips.

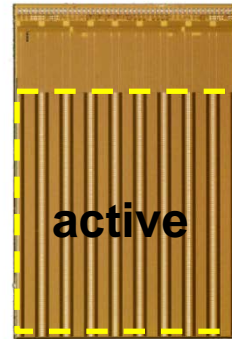
Two possible ways to reduce # chips

Larger pixel chips

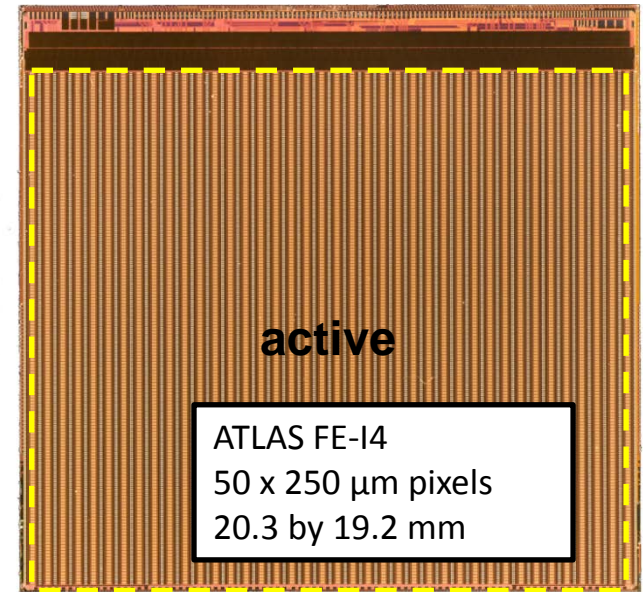
- ATLAS FE-I4: 10 x more pixels per dollar

Focusing of drift charge

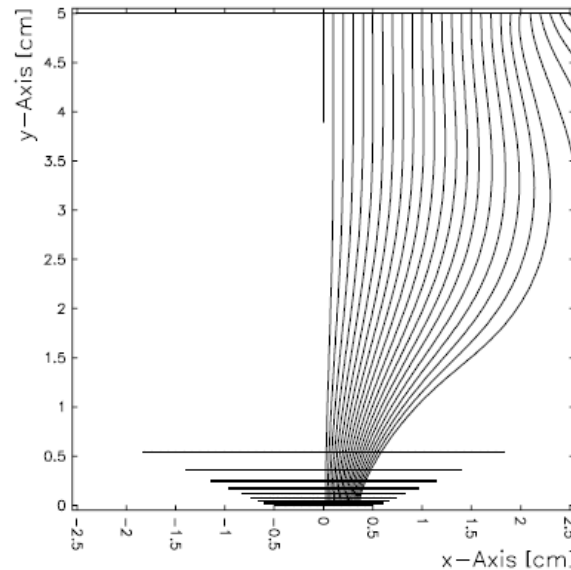
- advantage: read out large volume with small readout plane
- retains key advantage of pixels: small size \rightarrow low capacitance \rightarrow low noise
- status: First experimental test promising, but more detailed analysis needed



ATLAS FE-I3
50 x 400 μm pixels
7.4 by 11.0 mm



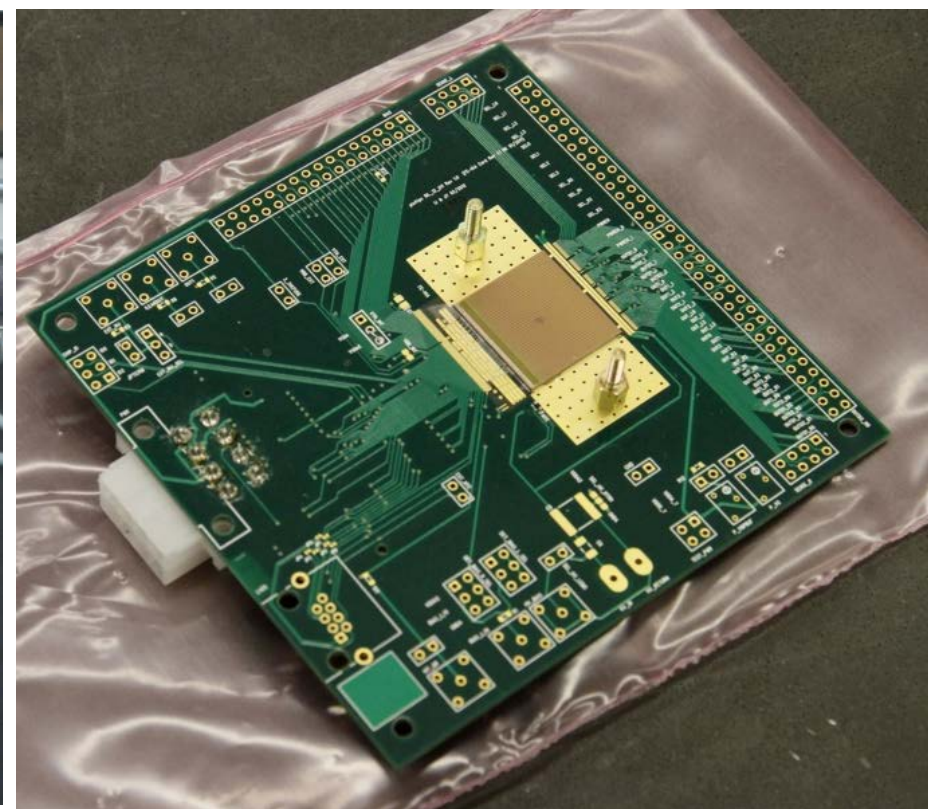
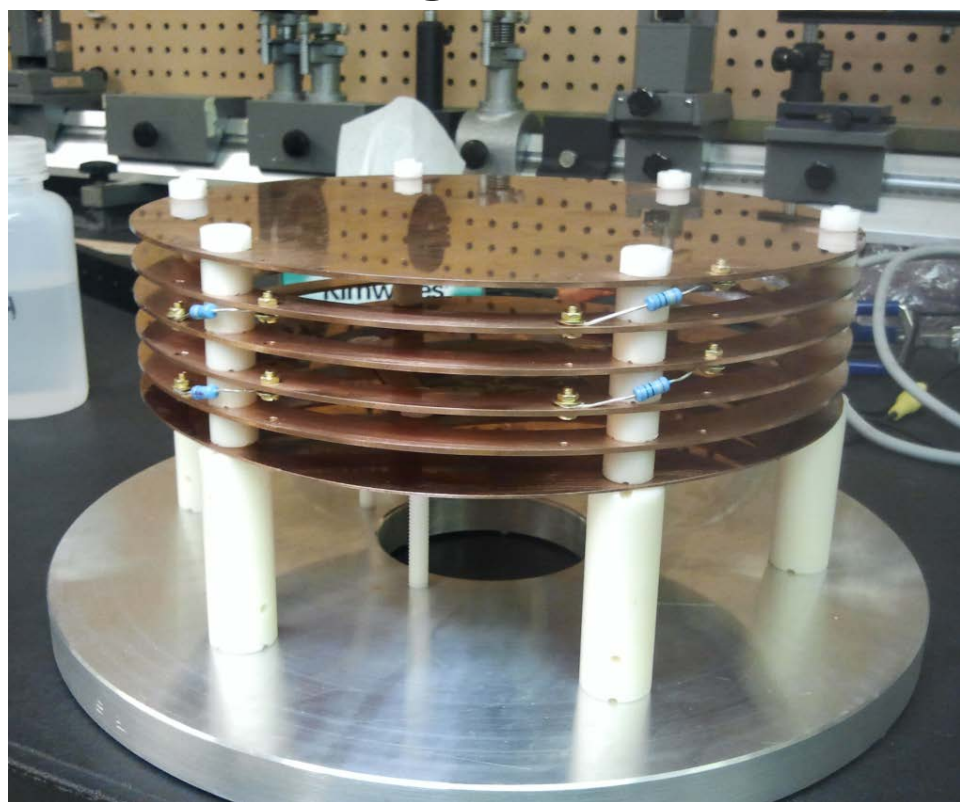
ATLAS FE-I4
50 x 250 μm pixels
20.3 by 19.2 mm



S. Ross et al., “Charge-Focusing Readout of Time Projection Chambers”, proceedings of IEEE NSS 2012

Detectors with FE-I4 Pixel Chip

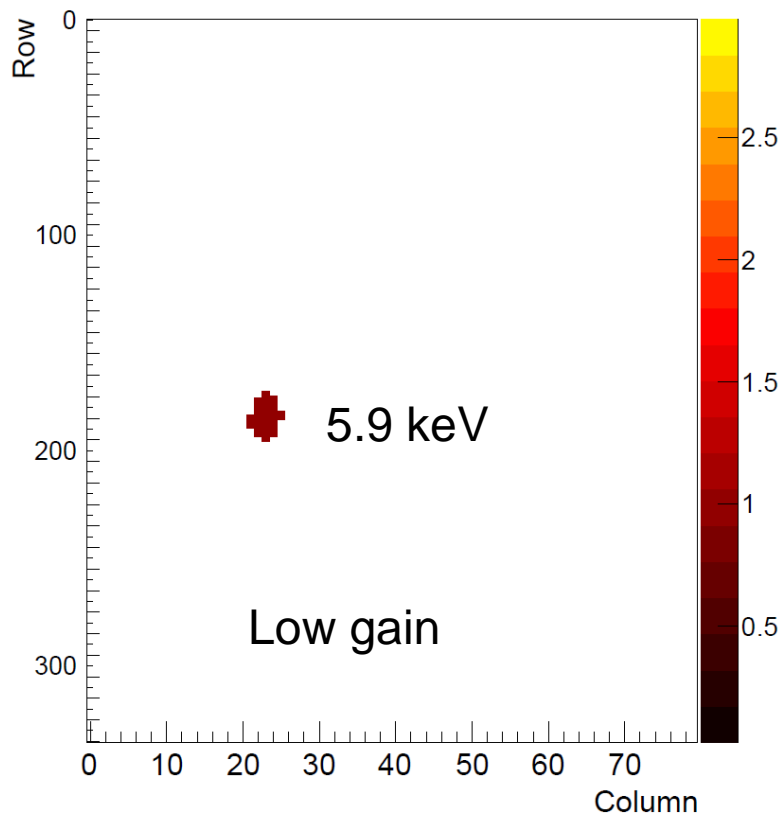
- FE-I4 single chip TPC card developed at Hawaii
- LBNL currently attempting first operation with this chip in their TPC
- Hawaii to operate larger mD³ detector this fall – components under production
 - Field cage
 - FE-I4 TPC card



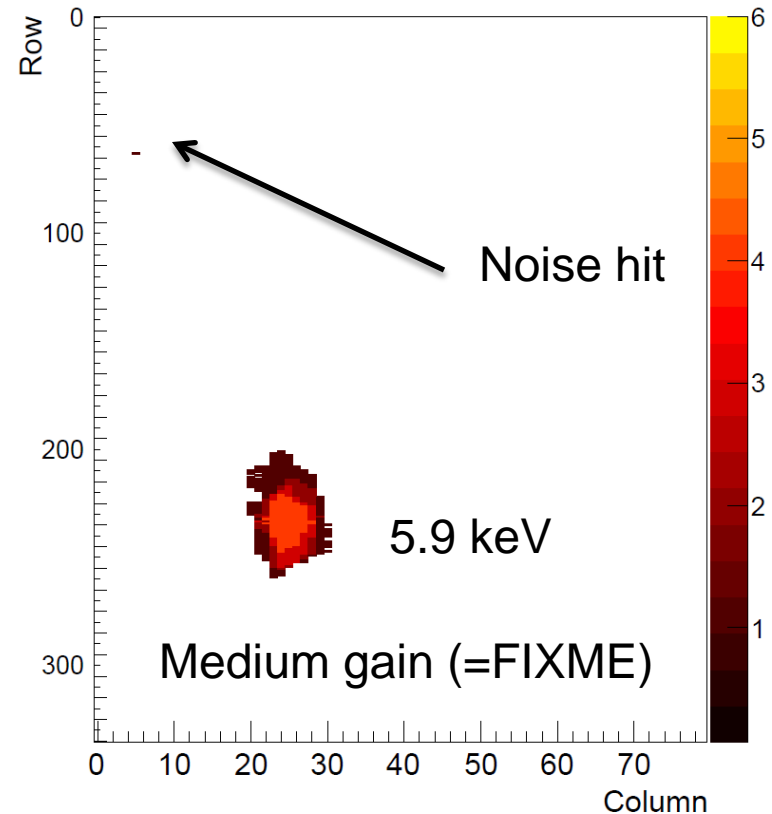
First events with larger pixel chip; FE-14

- Recorded at Berkeley Lab just last Friday - 6/7/2013
- These are self-triggered raw data – no noise suppression!
- Looks better than FE-13 – no column dependence

Occupancy mod 0 bin 0 chip 0

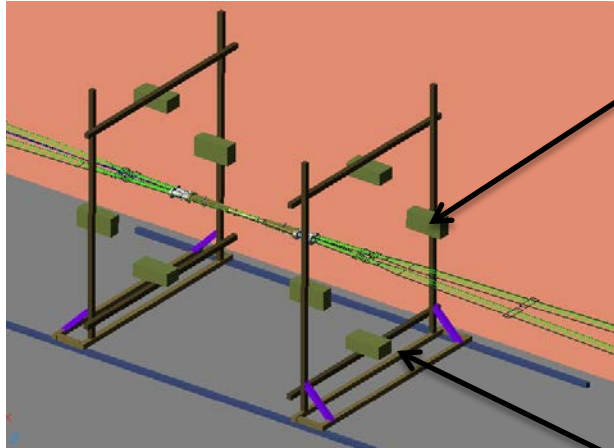


Occupancy mod 0 bin 0 chip 0



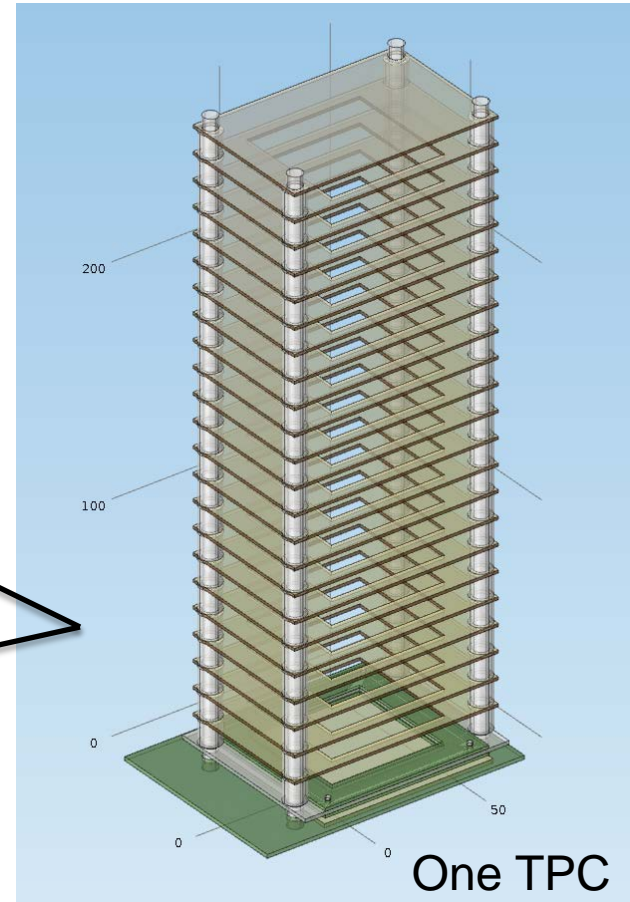
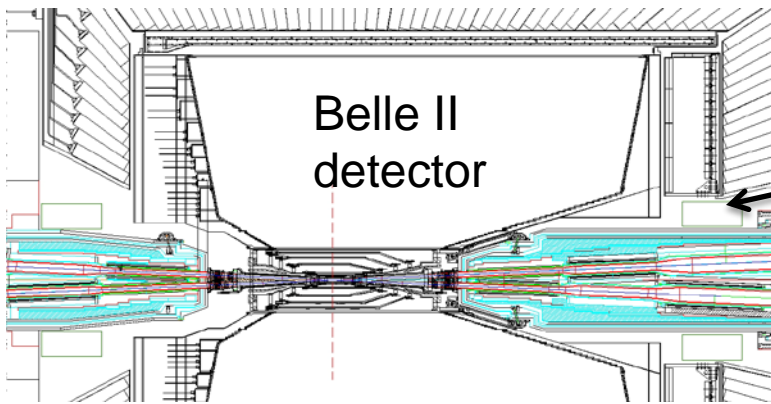
Broader Impacts: Neutrons at SuperKEKB

Commissioning Phase I – Jan 2015



SuperKEKB beam lines

Commissioning Phase II – Summer 2015



Fast neutrons are important beam background component at SuperKEKB e^+e^- collider
→ Will measure with eight neutron-TPCs; 25 cm drift lengths, 4 pixel chips each

Conclusion

- m^3 -scale gas TPC w/ low energy threshold may be sufficient to investigate hints for low-mass WIMPs w/ directionality
 - GEM + pixel readout promising technology for this application
 - 3D tracking, single electron sensitivity
- Characterization of 1-cm^3 prototype “D³-micro” nearly complete
 - Excellent performance at 1 atm
 - Some mysteries related to energy scale still under investigation
- Moving on to low-pressure operation, larger detectors, and next generation pixel chip this year

He-recoil in HeCO₂ at p=1 atm

