

Updates from the Dark Matter Time Projection Chamber Group (DMTPC)

Cygnus 2013 Workshop on Directional Dark Matter Detection
Toyama, Japan
2013 June 10

James Battat (Wellesley College)
for the DMTPC collaboration





DMTPC Collaboration



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Royal Holloway (UK)

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*=PI, postdoc, grad student, *undergrad*²

Gallery of DMTPC Detectors

10L



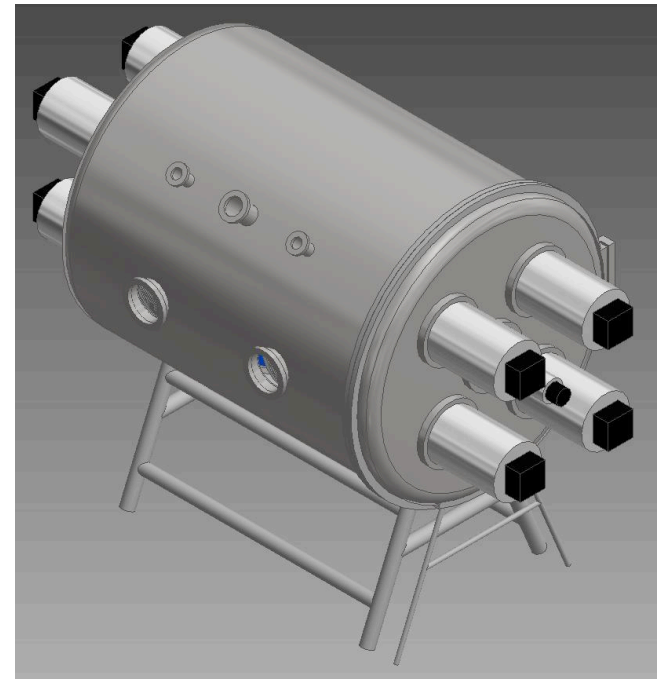
Underground at WIPP

4Shooter (20L)



At MIT

DMTPCino (m^3)



Under development

Gallery of DMTPC Detectors

10L



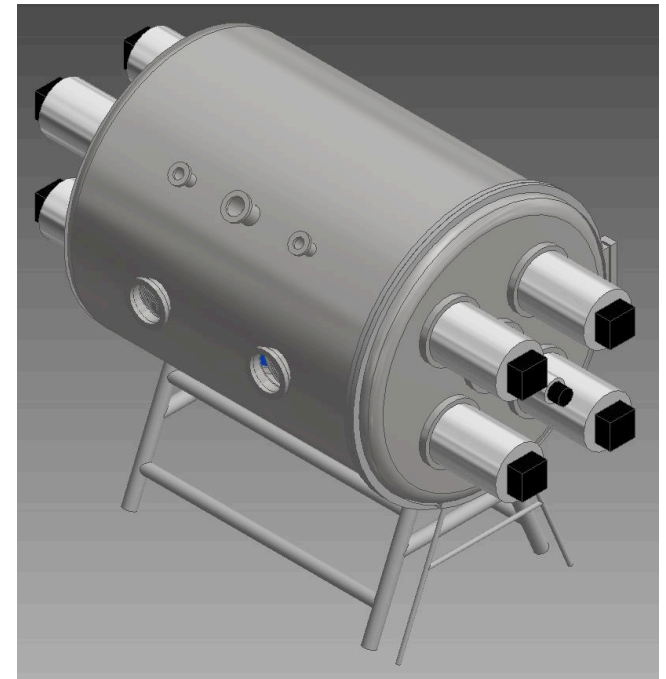
Underground at WIPP

4Shooter (20L)



At MIT

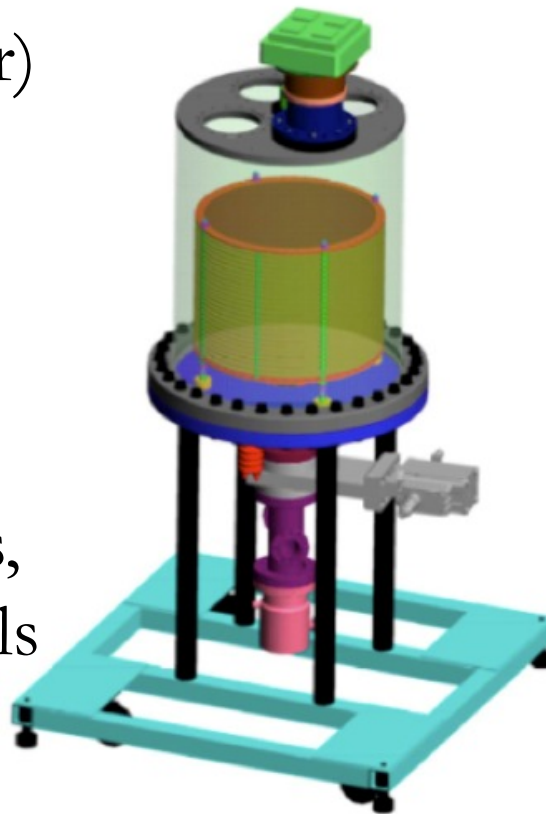
DMTPCino (m^3)



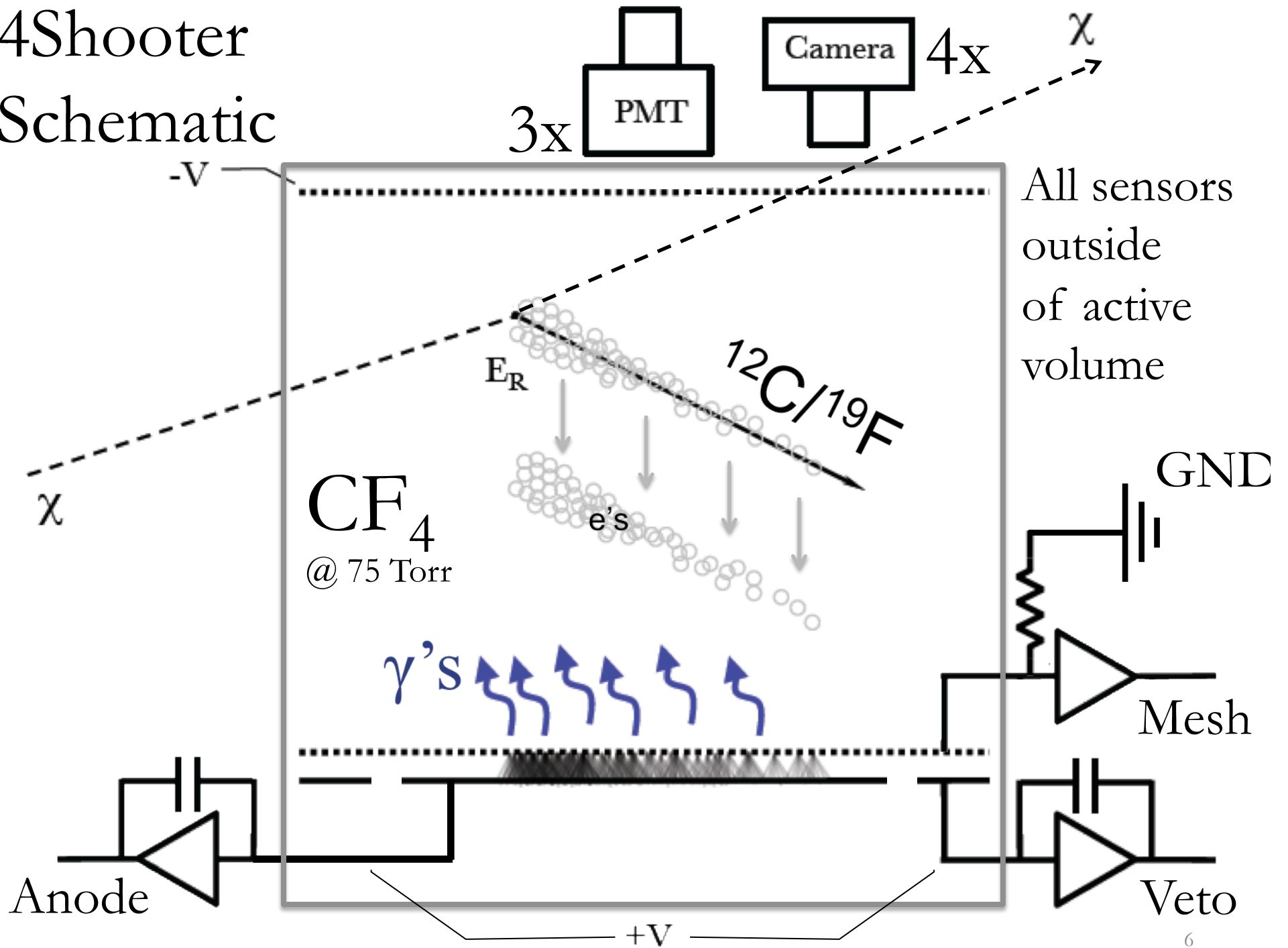
Under development

4Shooter Overview

- 20L (4.5g F)
- Higher vacuum (10^{-5} torr) between fills
- Material selection (OFHC copper, acetal, stainless steel, G-10)
- 4 CCD cameras, 3 PMTs, 3 charge readout channels

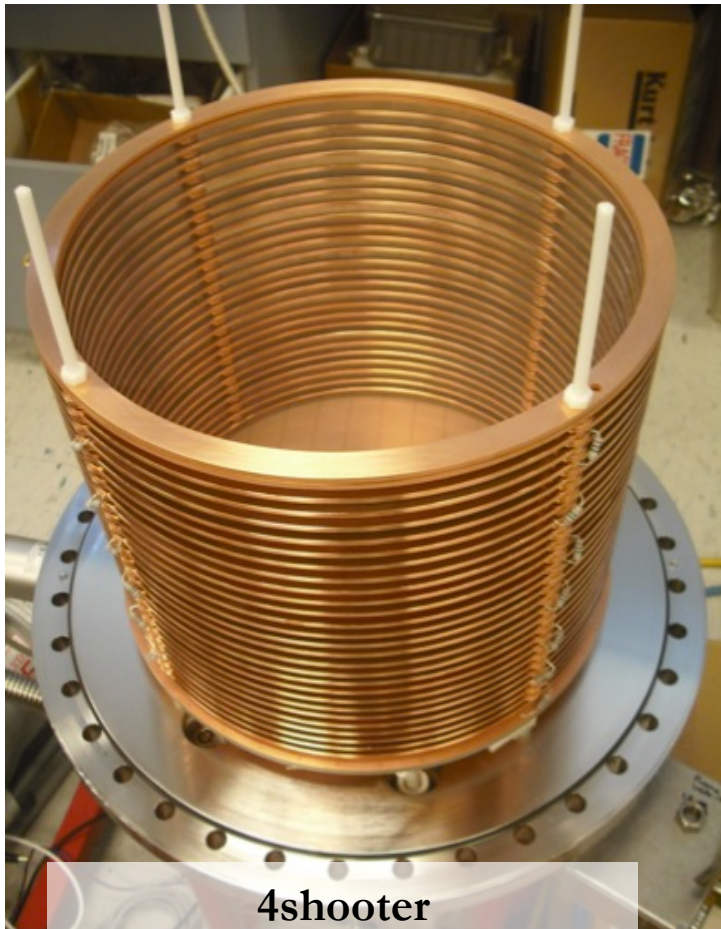


4Shooter Schematic



4Shooter Construction

Improved Field Cage (spacers), and amplification region fabrication scheme → More gain, lower spark rate. Repeatable amplification region construction technique in a clean-room environment.



vs.



Lower alpha background rate & better tagging capability

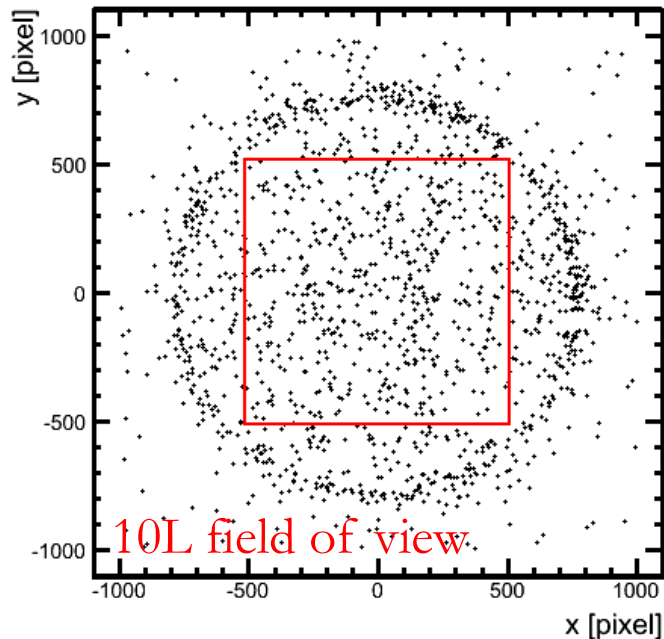
Raw α rate is 19x lower

10L 210 mHz

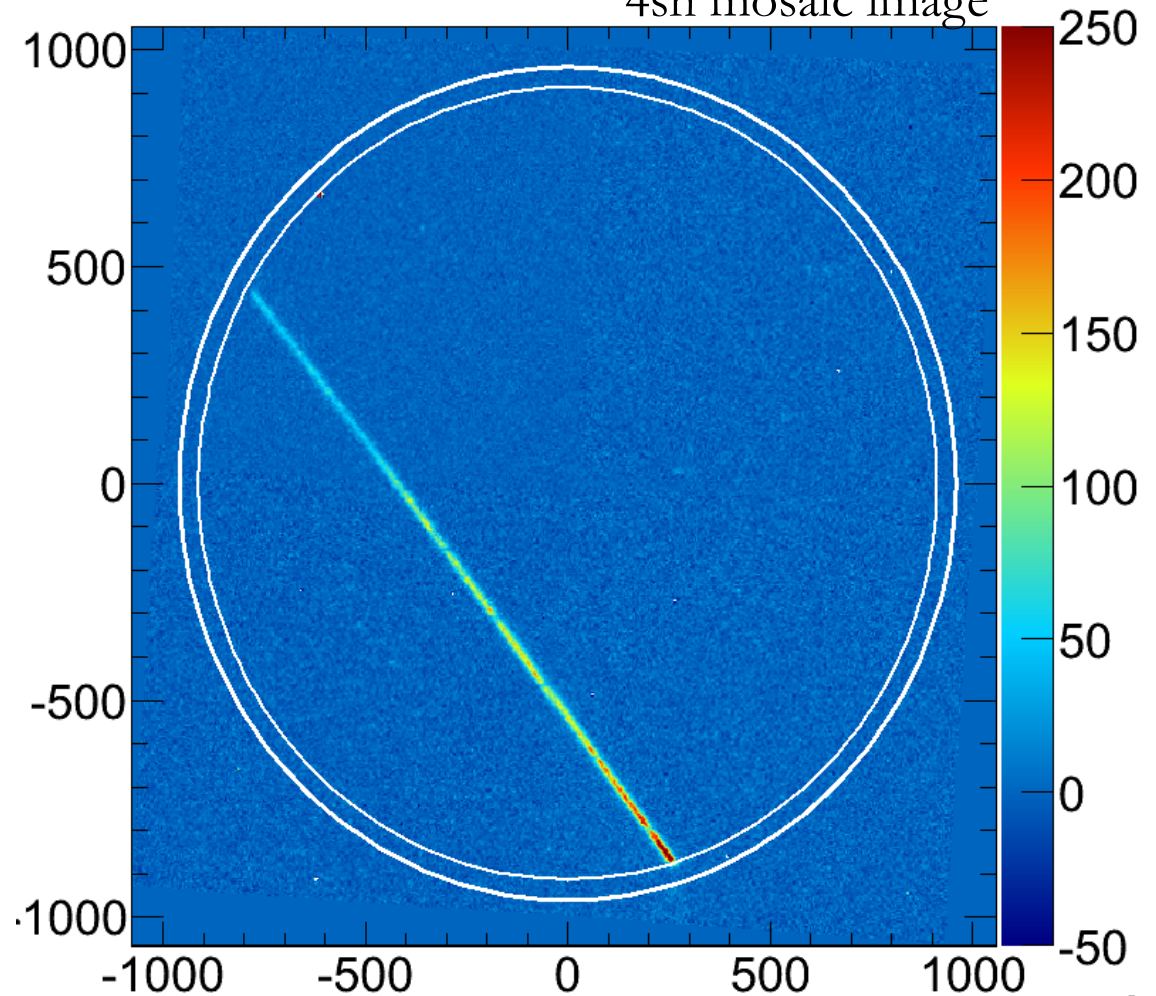
4sh 11 mHz

And 4sh field cage has 3x more internal surface area

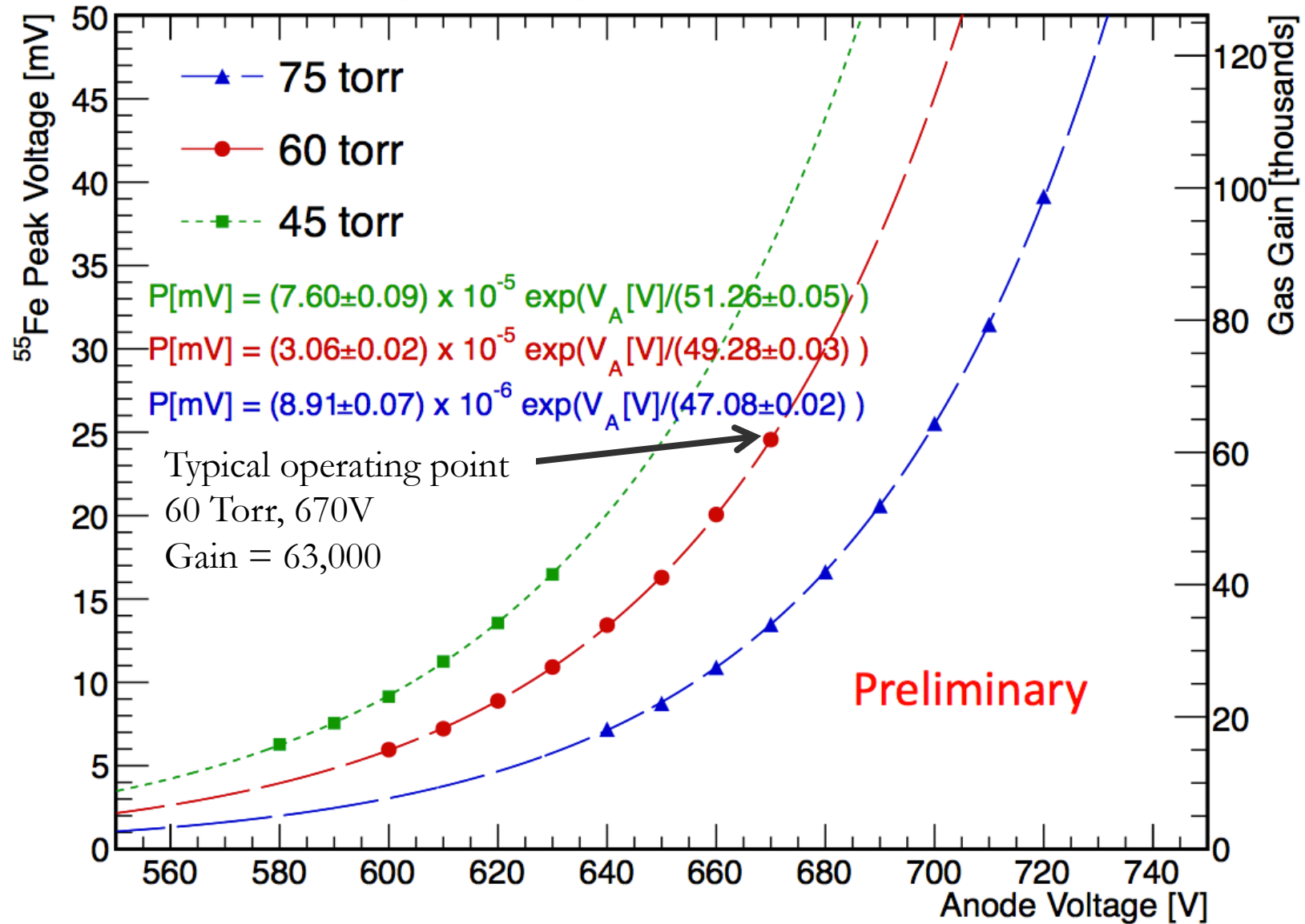
10L: inferred alpha start points



4sh mosaic image

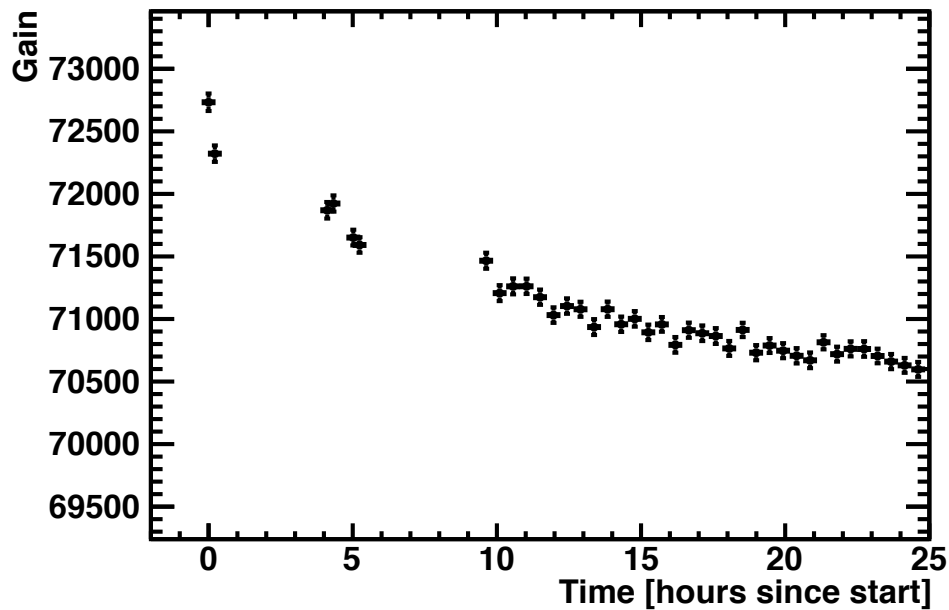


4sh gas gain measurements

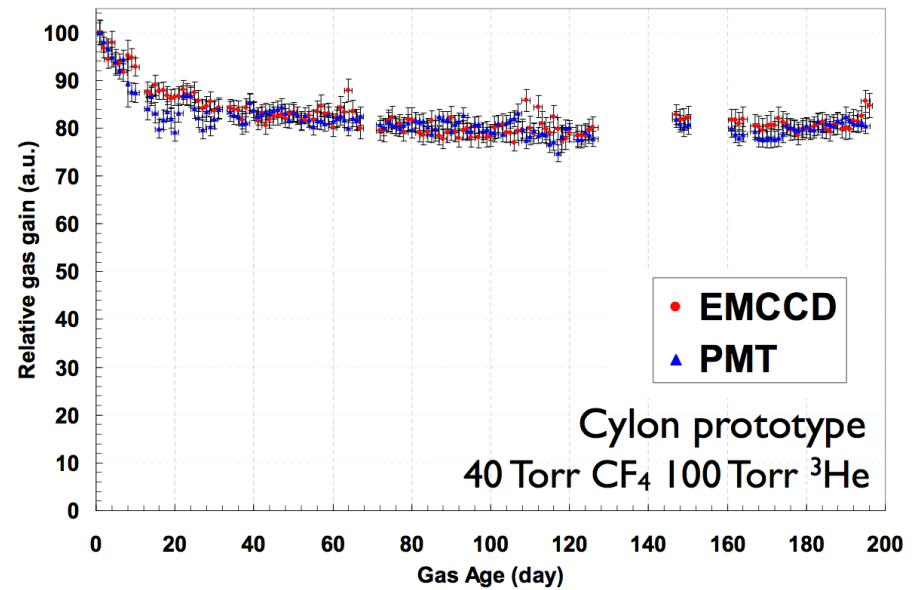


Lesson 1: gain degradation vs. time

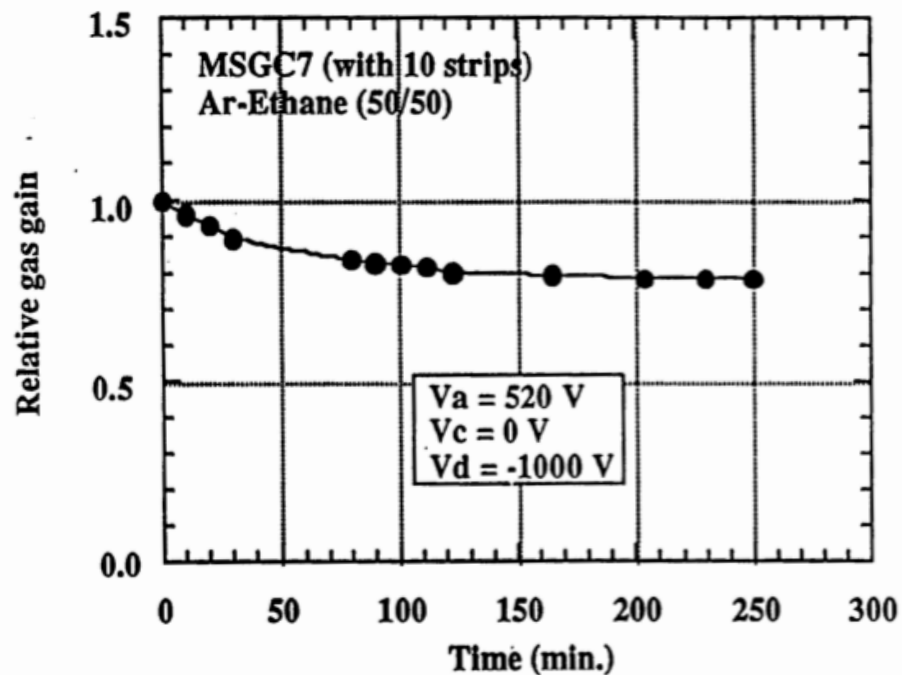
4sh over 1 day ($\sim 3\%$ drop)



Cylon over 200 days
(20% drop then stable)



Gain degradation over time is well-known



---> 20 % gain drop within 200 min.

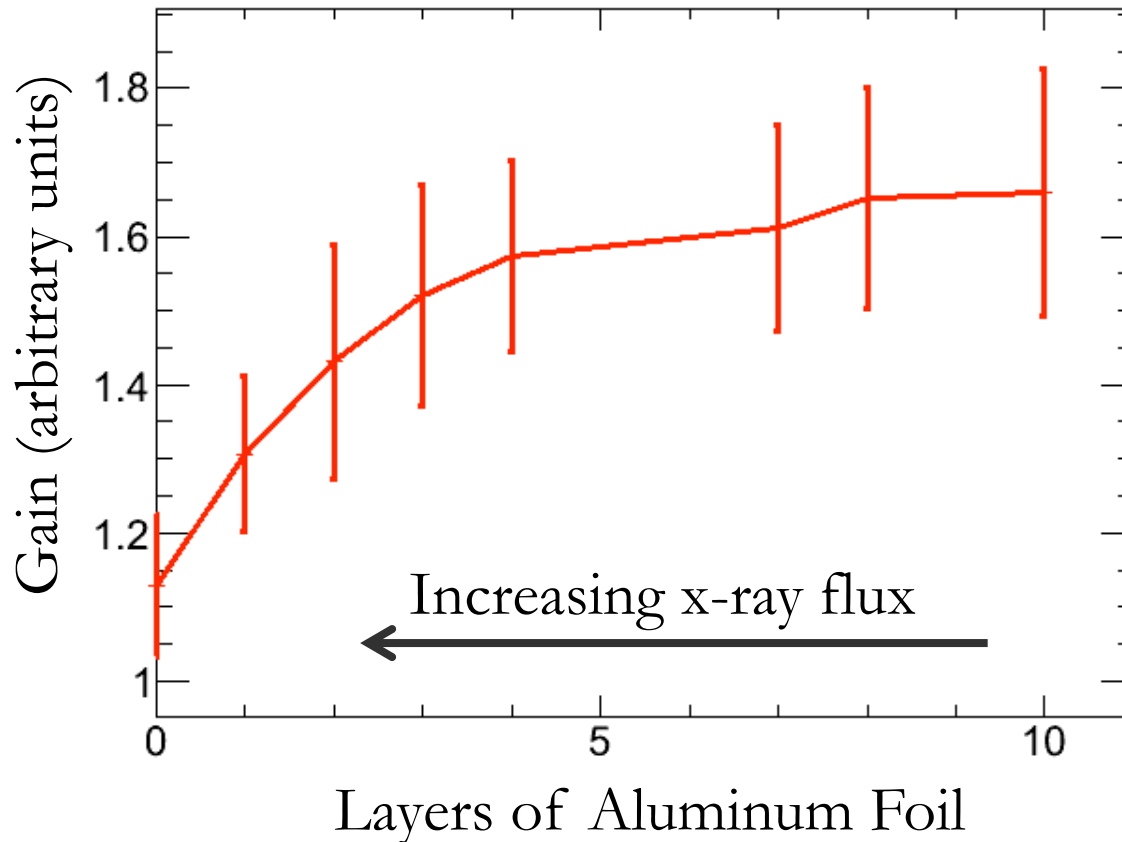
Initial gain decay is well-known and seen in many detectors. For example:

Kadyk 1998 (SLAC Detector Techniques Lectures)

<http://www-group.slac.stanford.edu/sluc/lectures/Detector-Lectures.html>

Lesson 2: gain decreases with x-ray flux

Likely due to space-charge in amplification region

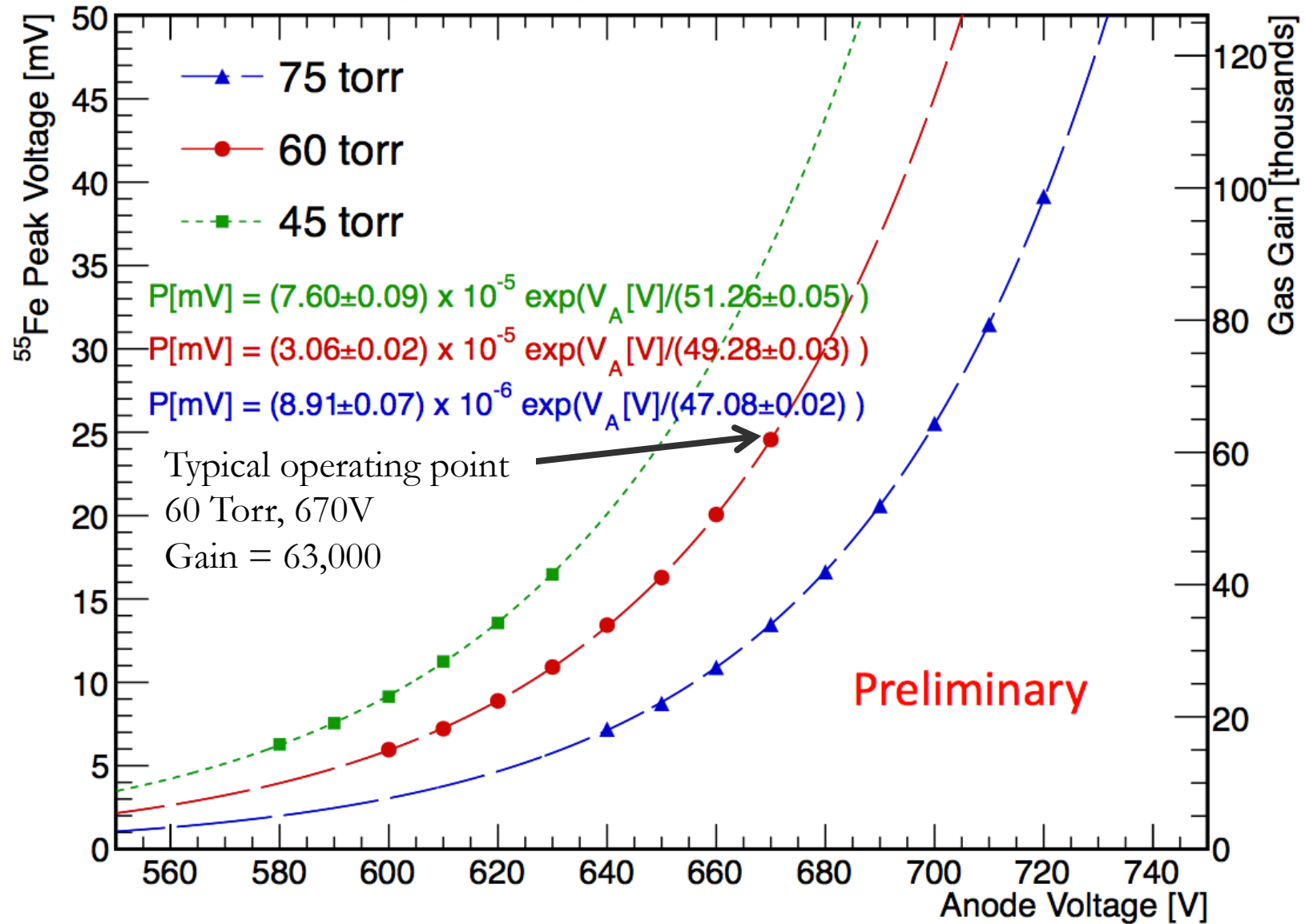


Calibration with a strong source can underestimate the gas gain by $\sim 30\%$.

High rate behavior and discharge limits in micro-pattern detectors

A. Bressan^a, M. Hoch^a, P. Pagano^a, L. Ropelewski^a, F. Sauli^{a,*}, S. Biagi^b, A. Buzulutskov^c,
M. Gruwé^d, G. De Lentdecker^e, D. Moermann^f, A. Sharma^g

Gain limited by streamer discharges

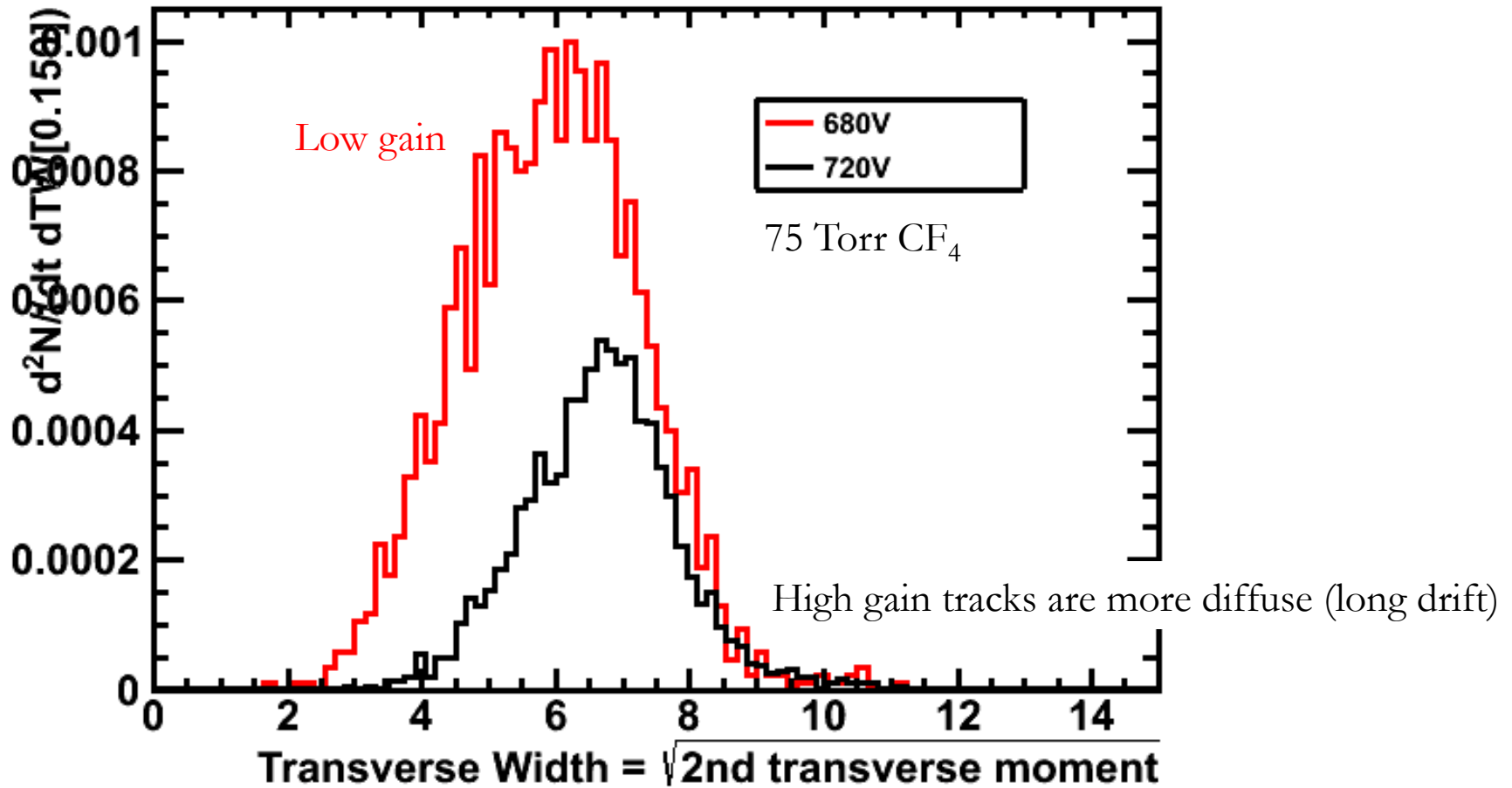


Sparking and Raether limit

- “Raether’s criterion: a spontaneous transition from avalanche to streamer, followed by a discharge, when the avalanche size reaches a value of a few 10^7 .”
- “In multiple structures, where the gain is shared between two devices in cascade, the maximum overall gain under irradiation is increased by at least one order of magnitude; we speculate this to be a consequence of a voltage dependence of Raether’s limit, larger for low operating potentials.”

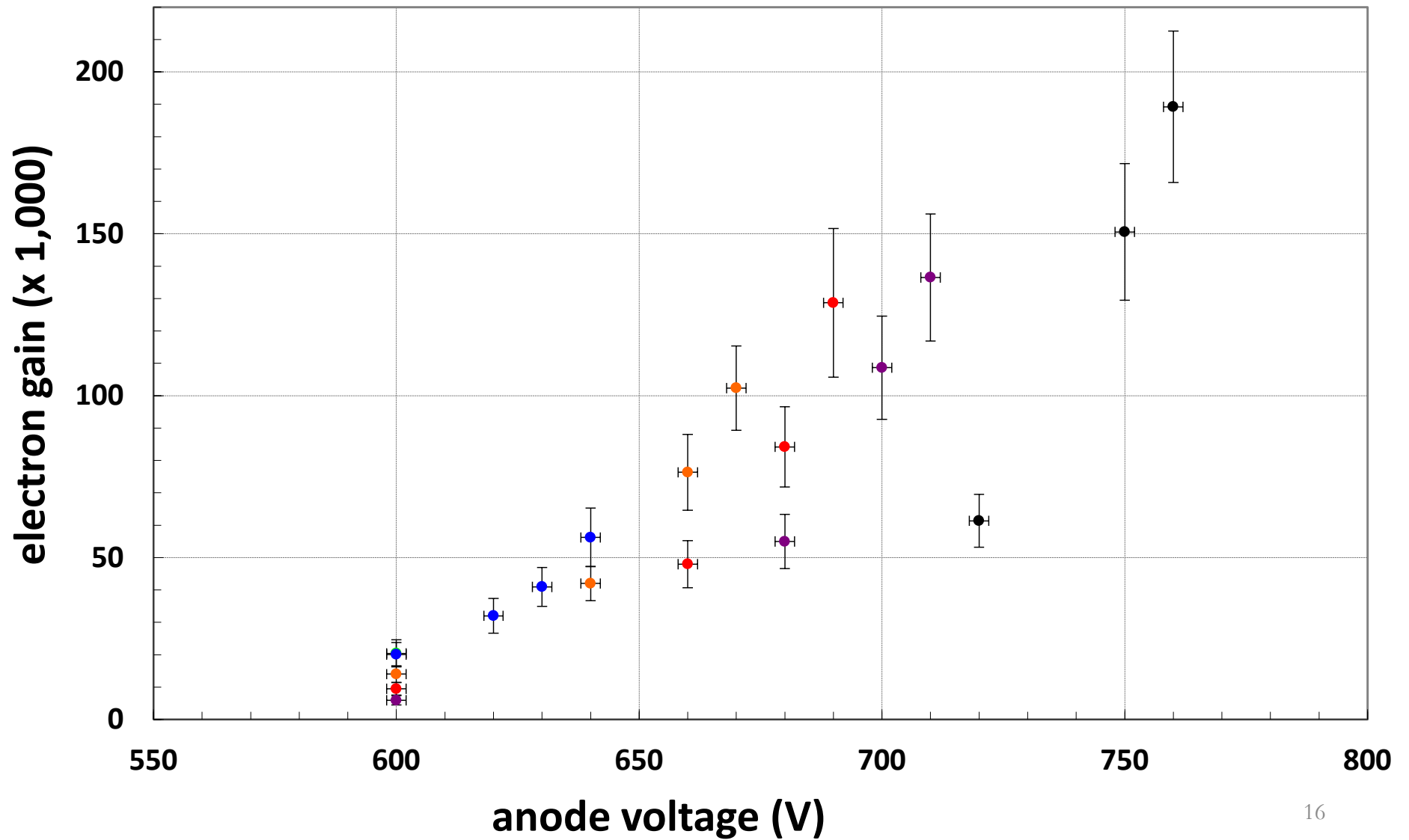
A. Bressan et al., Nuclear Instruments and Methods in Physics Research A 424 (1999) 321—342

Lesson 3: Recoils can trigger sparks



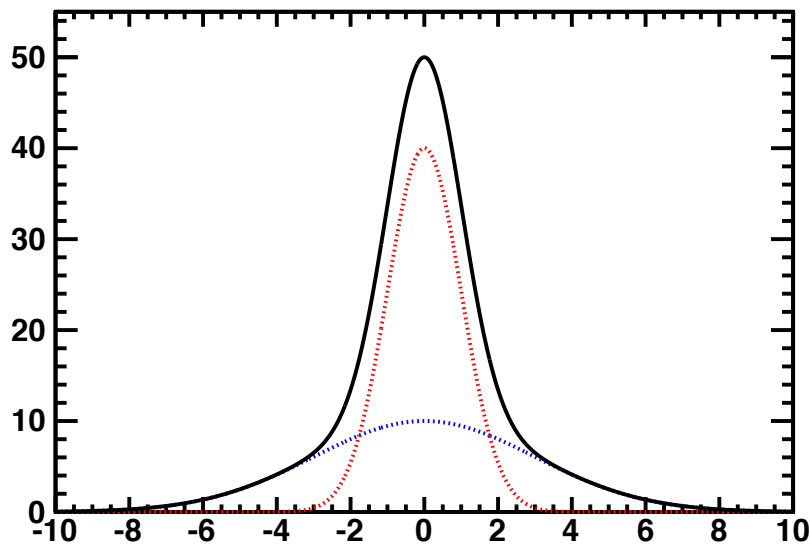
If electron density is too high, the avalanche becomes a streamer, rendering the detector insensitive. In practice, compact (low-drift) tracks preferentially generate streamers/sparks. The lower part of the TPC is not “active” anymore! Also this can lead to an angular dependence as well (vertical tracks lead to sparks)

Electron Gain in the amplification region measured with ^{241}Am α tracks in 40 torr CF_4 gas

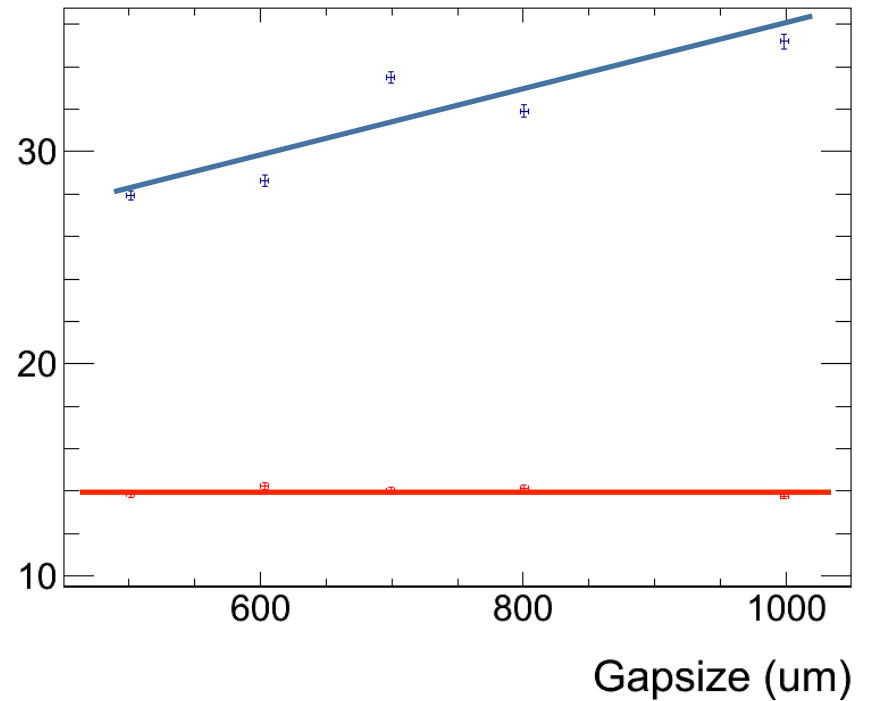


Diffusion vs. amplification gap size

Projected alpha tracks
well fit by double (not single)
gaussian

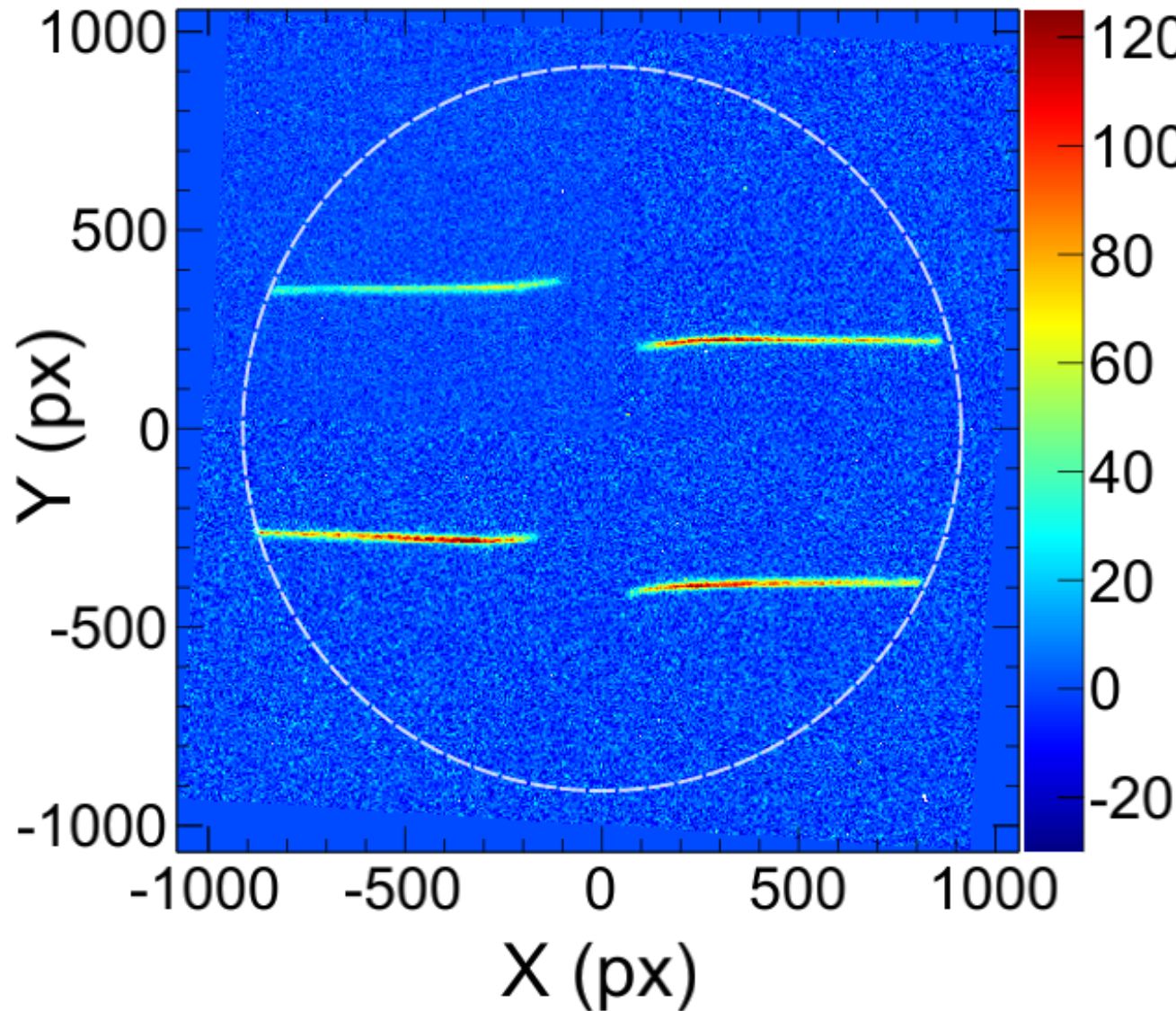


Sigma (pix)



As gap doubles, the **core** is
unchanged, but **wings** widen
by 25%.

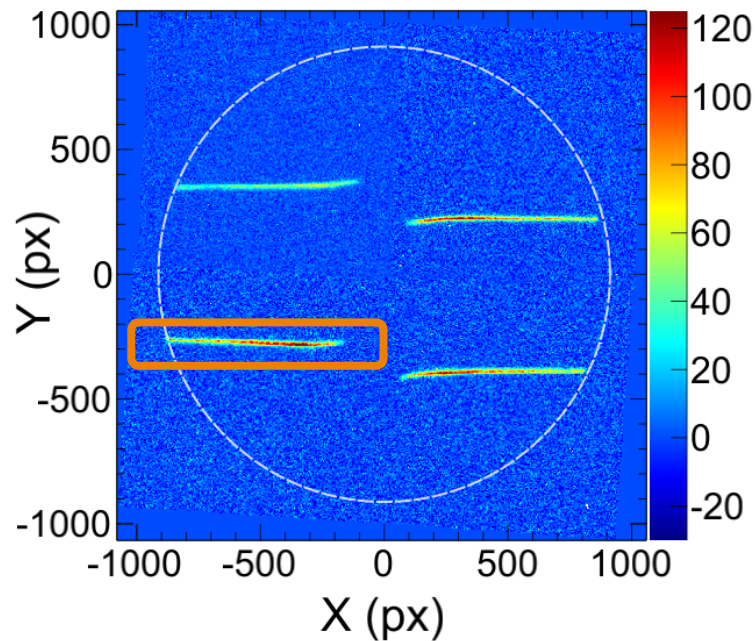
CCD Energy Calibration



One ^{241}Am for each camera.

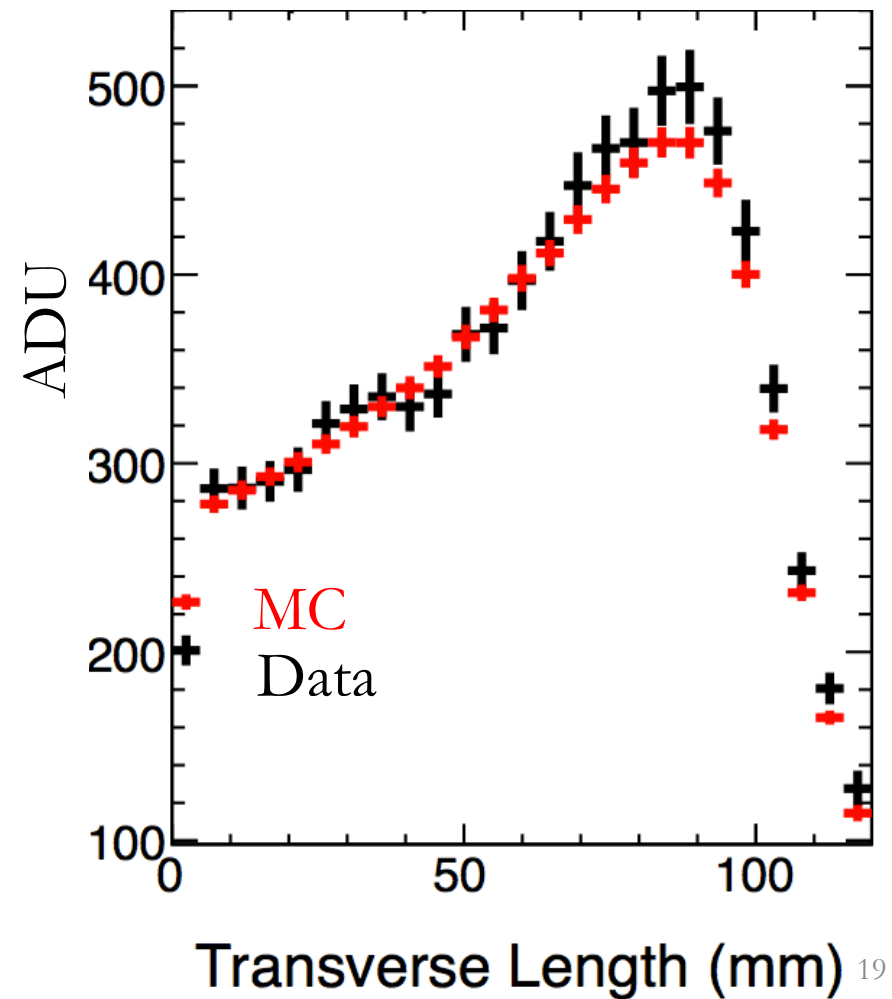
Side question:
do ^{241}Am x-rays
degrade the gain?
(59 keV from ^{241}Am , plus
14, 17, 21 keV lines from
 ^{237}Np)

CCD Energy Calibration



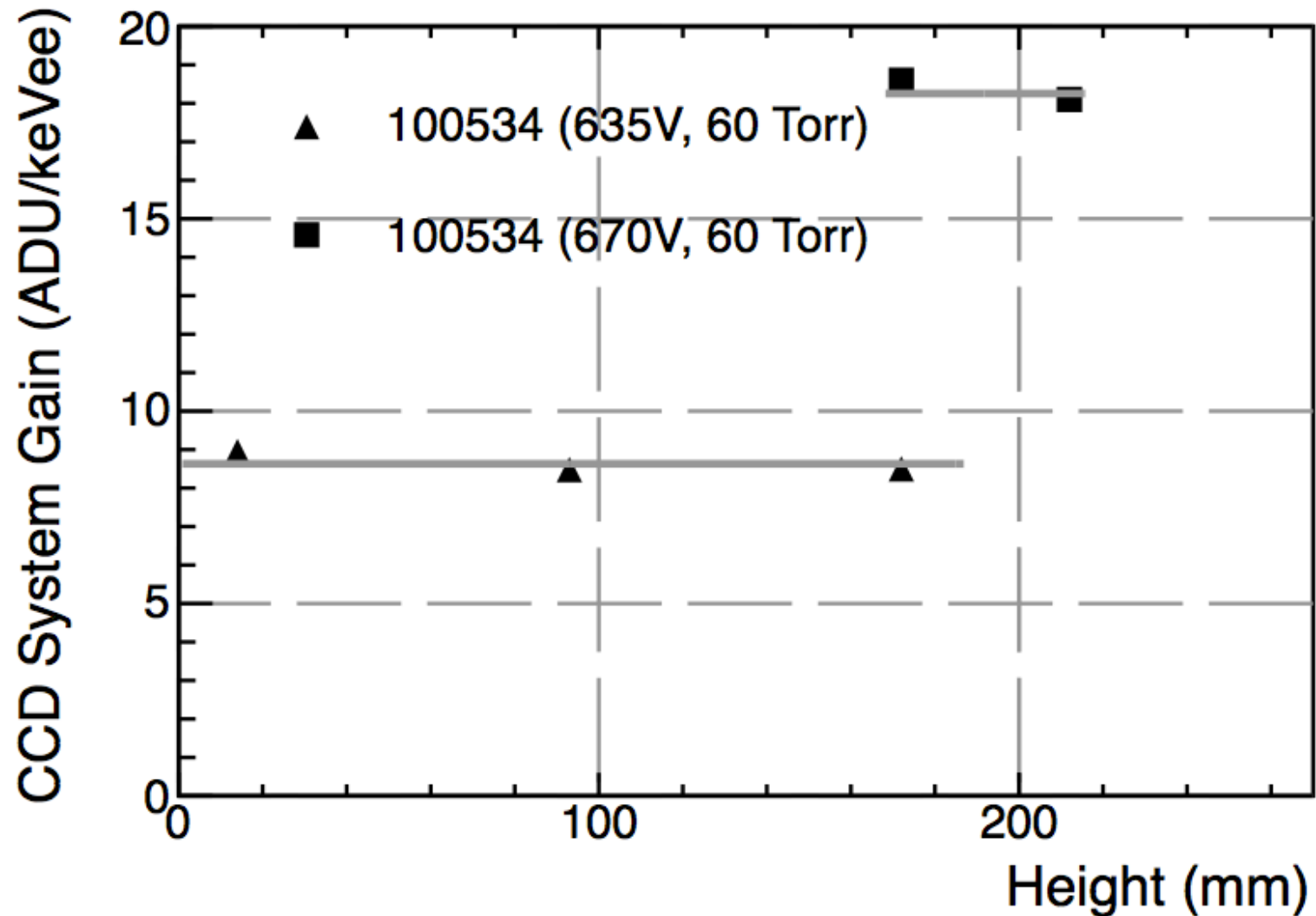
Tune electron diffusion and “gain” (ADU/keVee) until MC matches data (for transverse and longitudinal projections).

Average of projections for many tracks (single camera)

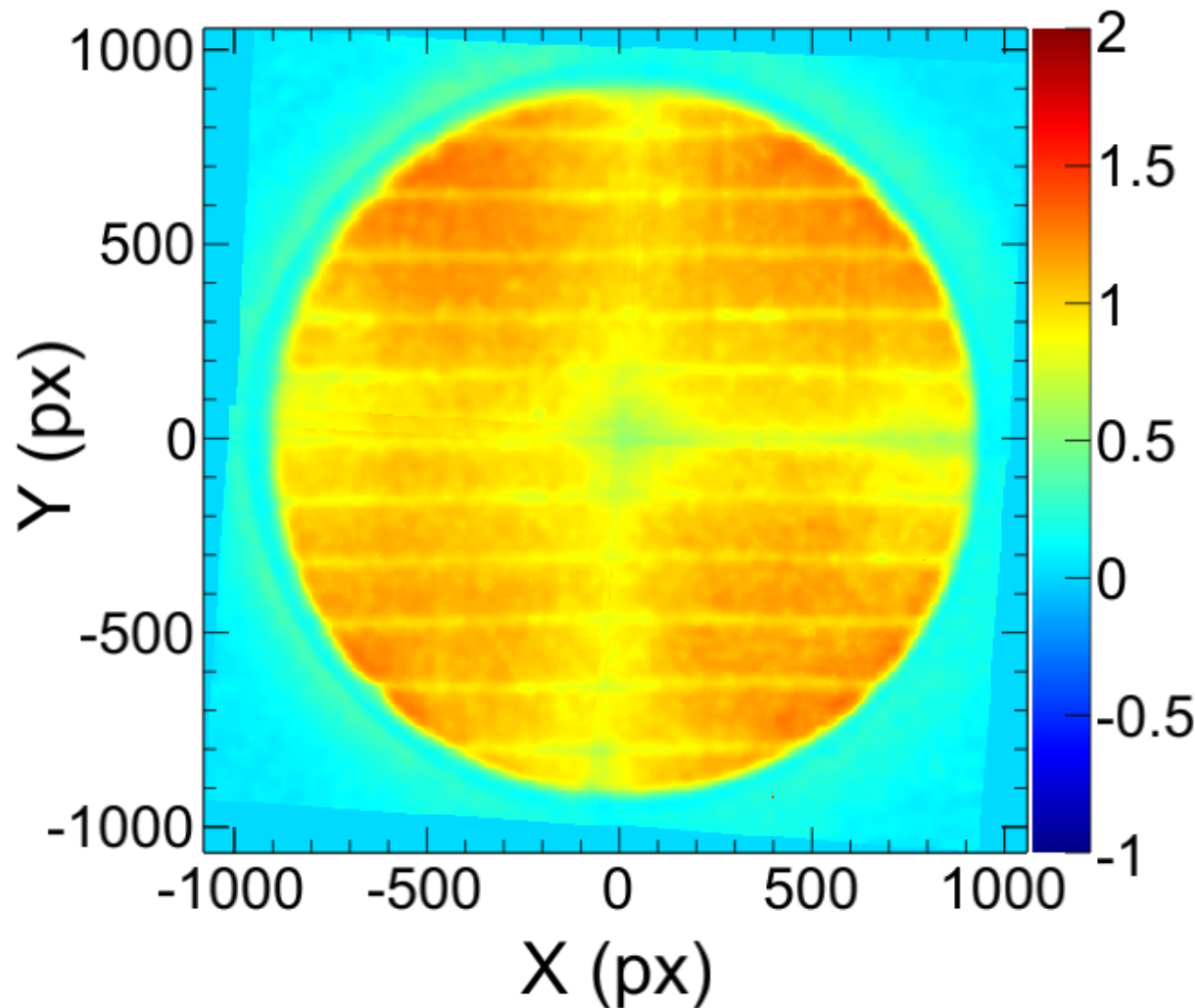


CCD Energy Calibration

As expected, the energy calibration depends on amplification region voltage (gas gain), but not on alpha source height (diffusion)



Spatial variation in detector response



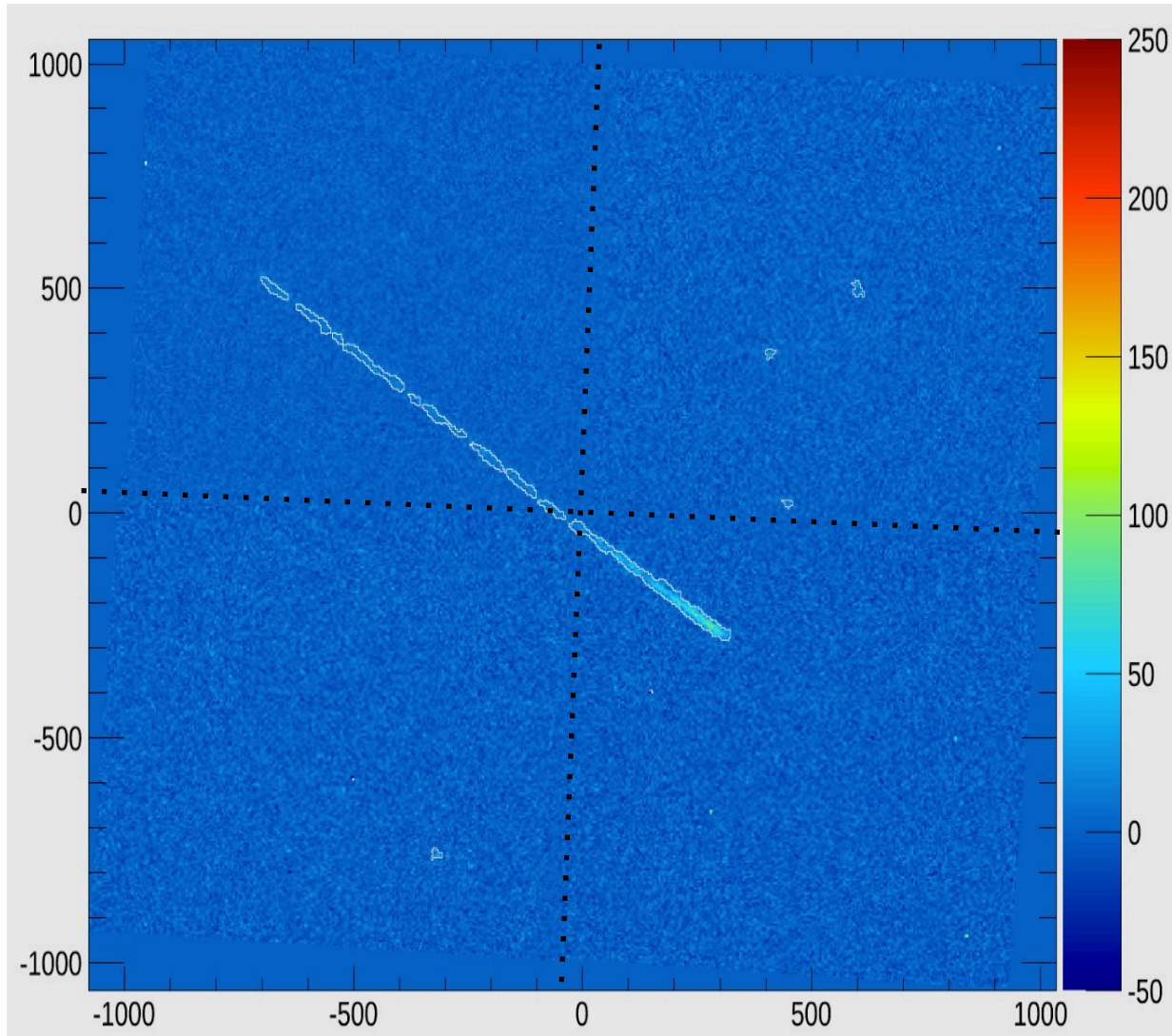
^{57}Co 122 keV γ

MFP = 4 m

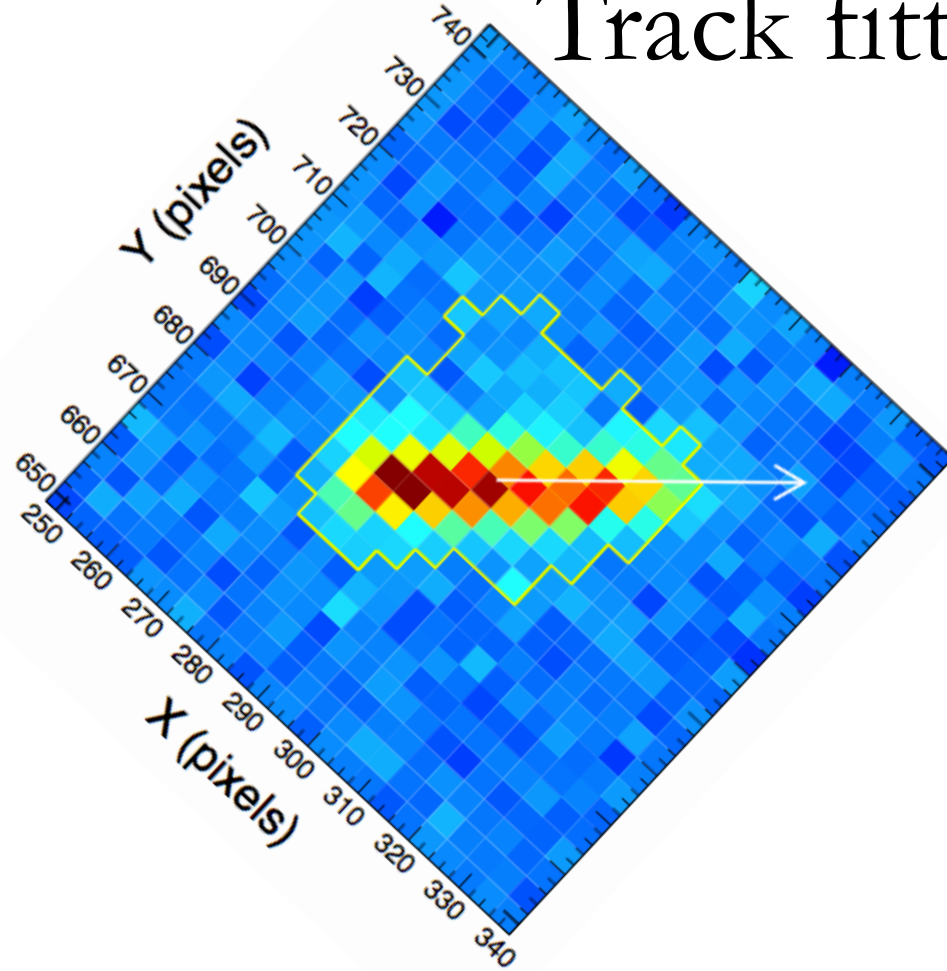
Each of the 4 cameras has a smooth rotational symmetry to the gain pattern.

The amplification region (especially the spacers) shows higher-frequency variations

Track reconstruction in mosaic image



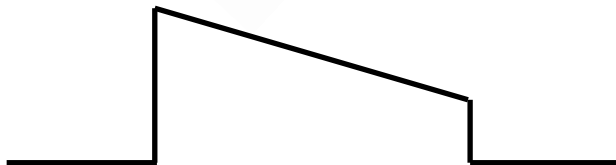
Track fitting with confidence



Make use of the known profile of a NR (from the Bragg curve) to

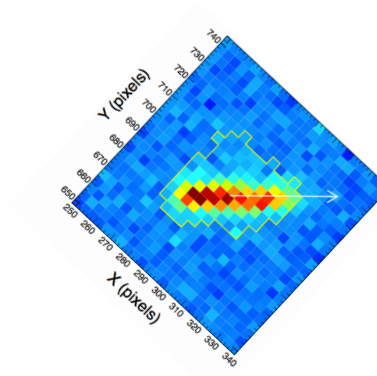
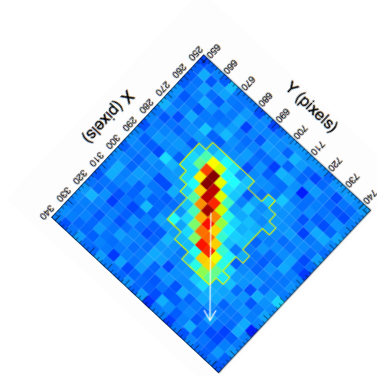
(1) fit for the Head/Tail and

(2) assign a confidence in the H/T determination

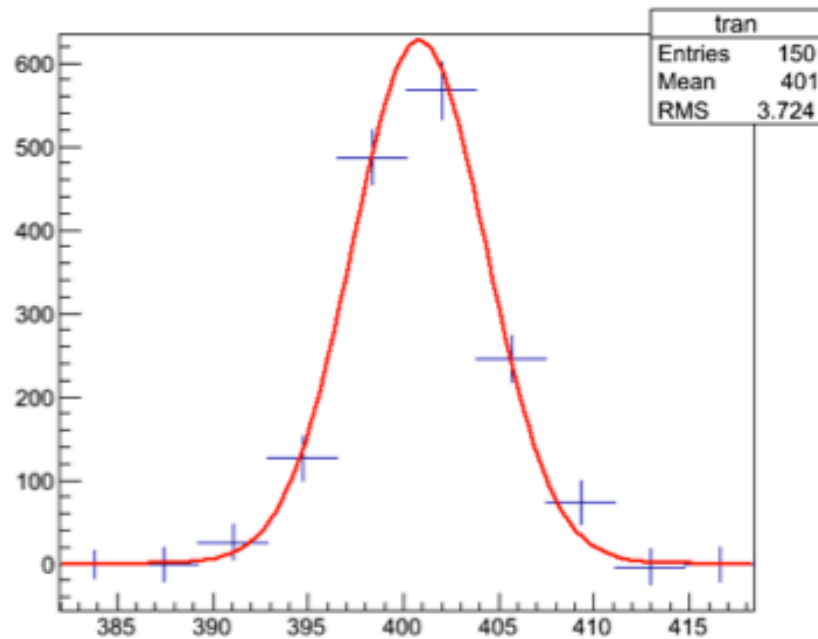


convolved with gaussian

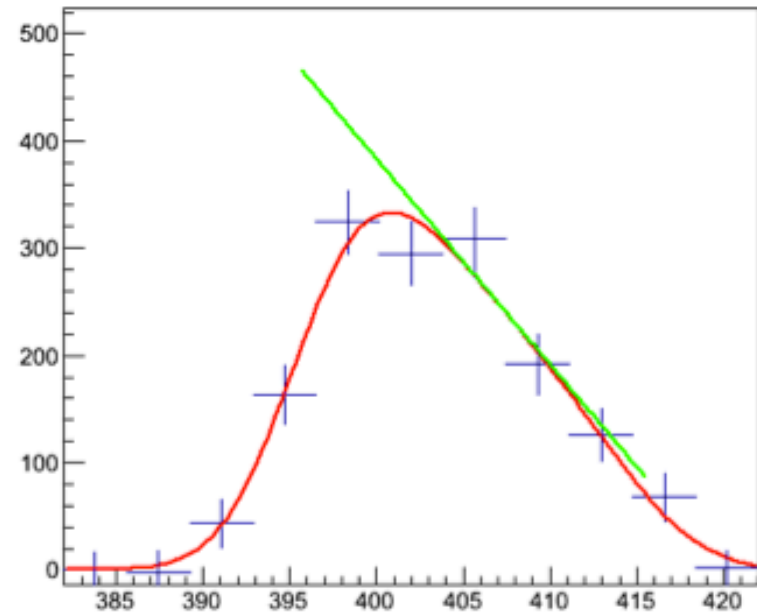
Track fitting with confidence



Transverse Fit



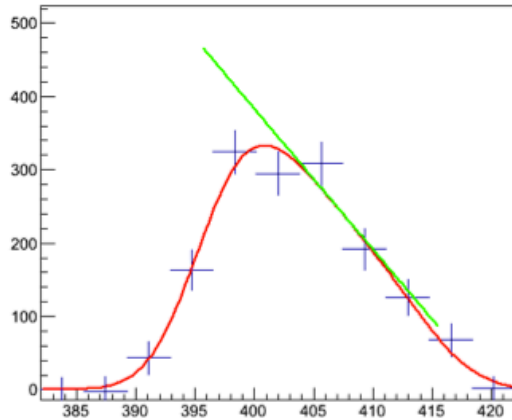
Longitudinal Fit



(CCD image is just to illustrate point)

Track fitting with confidence

Longitudinal Fit



$\chi^2/\text{ndof} : 4.352148/10 = 0.435215$

$x0: 395.668893 \pm 0.625993 (-0.664133, 0.598328)$

$y0: 466.210150 \pm 51.341684 (-56.980963, 46.037852)$

$y1: 85.021626 \pm 103.864438 (-64.085506, 99.278767)$

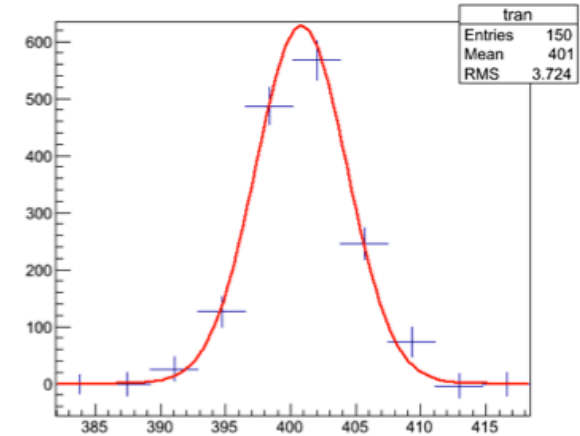
E: 1505.203272 \pm 9.356560 adu (62.716803 \pm 0.389857 keVee)

Range: 19.912867 \pm 8.374463 pixels (3.186059 \pm 1.339914 mm)

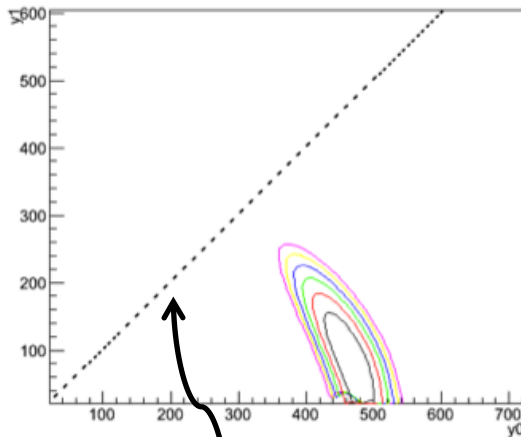
sin θ (from range): 0.888838 (from min): 1.237788

Implied Probability: 0.967850

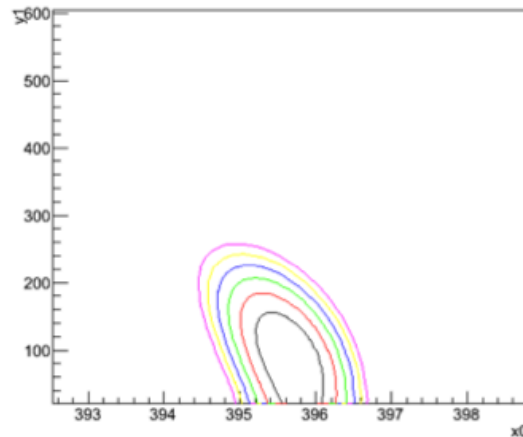
Transverse Fit



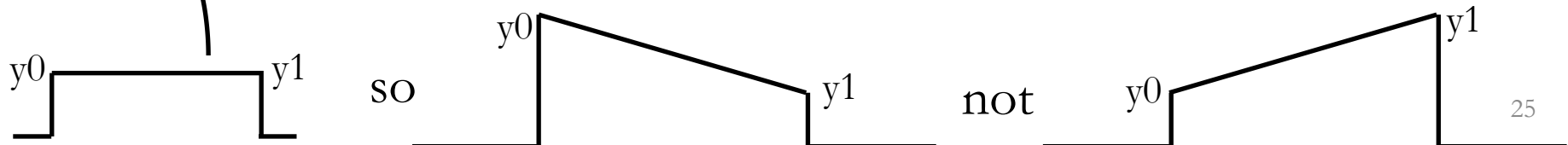
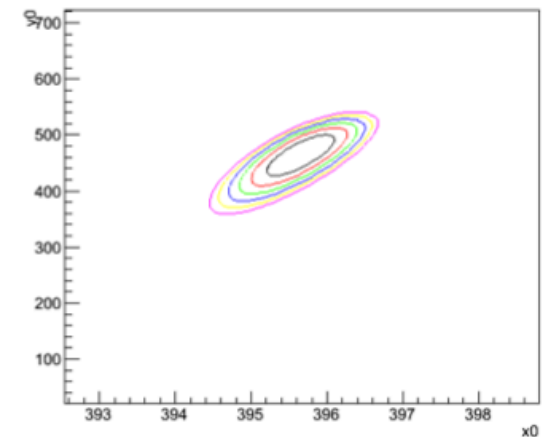
y0 vs y1



x0 vs y1

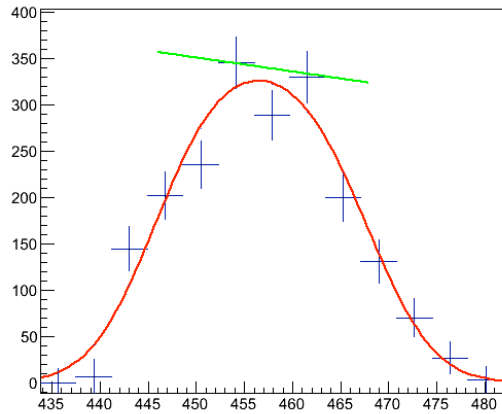


x0 vs y0



Track fitting with confidence

Longitudinal Fit



$\chi^2/\text{ndof} : 15.027260/12 = 1.252272$

$x0: 445.931635 \pm 1.318754 (-1.449950, 1.237641)$

$y0: 356.582855 \pm 99.450537 (-102.434698, 97.579702)$

$y1: 323.325790 \pm 106.787317 (-114.698638, 102.019991)$

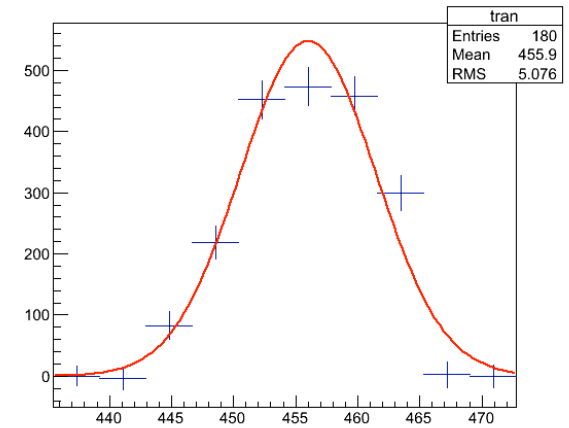
E: 2018.897334 \pm 6.212286 adu (80.755893 \pm 0.248491 keVee)

Range: 21.993059 \pm 9.441429 pixels (3.518889 \pm 1.510629 mm)

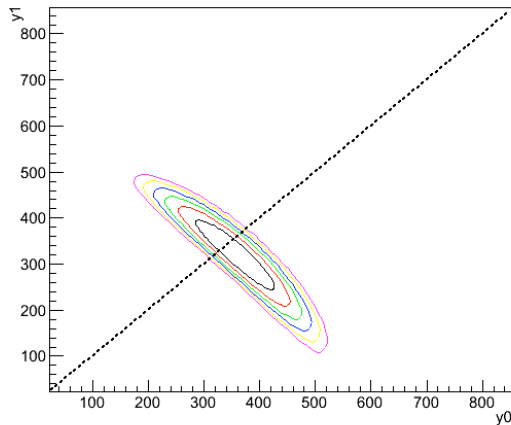
sin θ (from range): 0.843616 (from min): 0.348456

Implied Probability: 0.579762

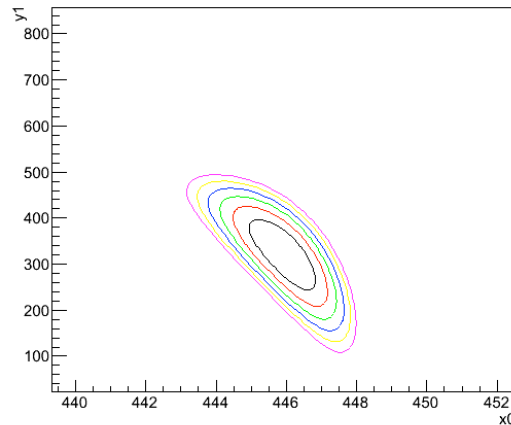
Transverse Fit



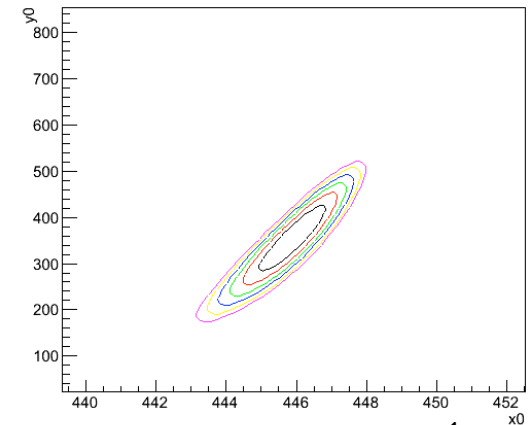
y0 vs y1



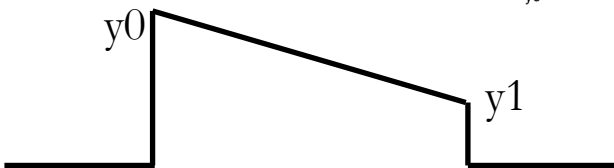
x0 vs y1



x0 vs y0



y0



y1

indistinguishable from

y0

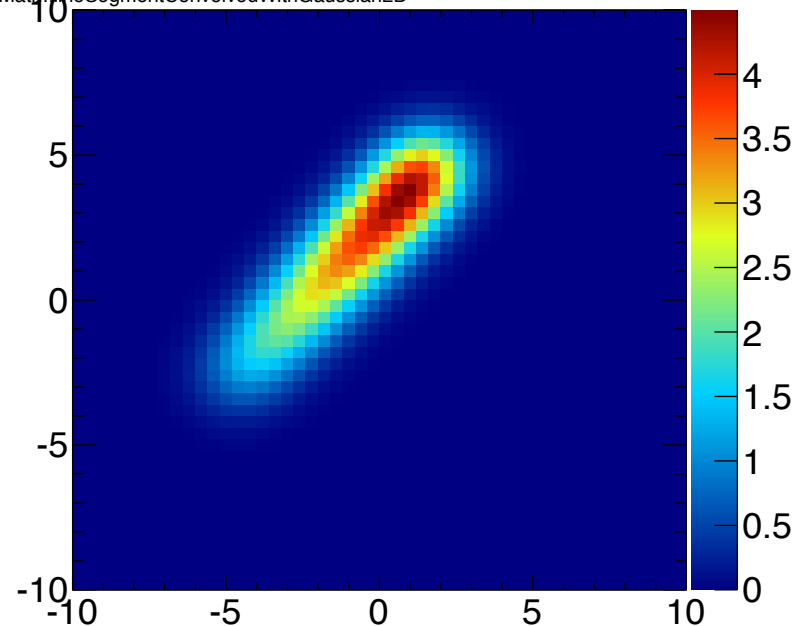


y1

Track fitting with confidence: next steps

- Evaluate benefits of restricting DM analysis to “high confidence” tracks only?
- Extend the track fitting to 2D

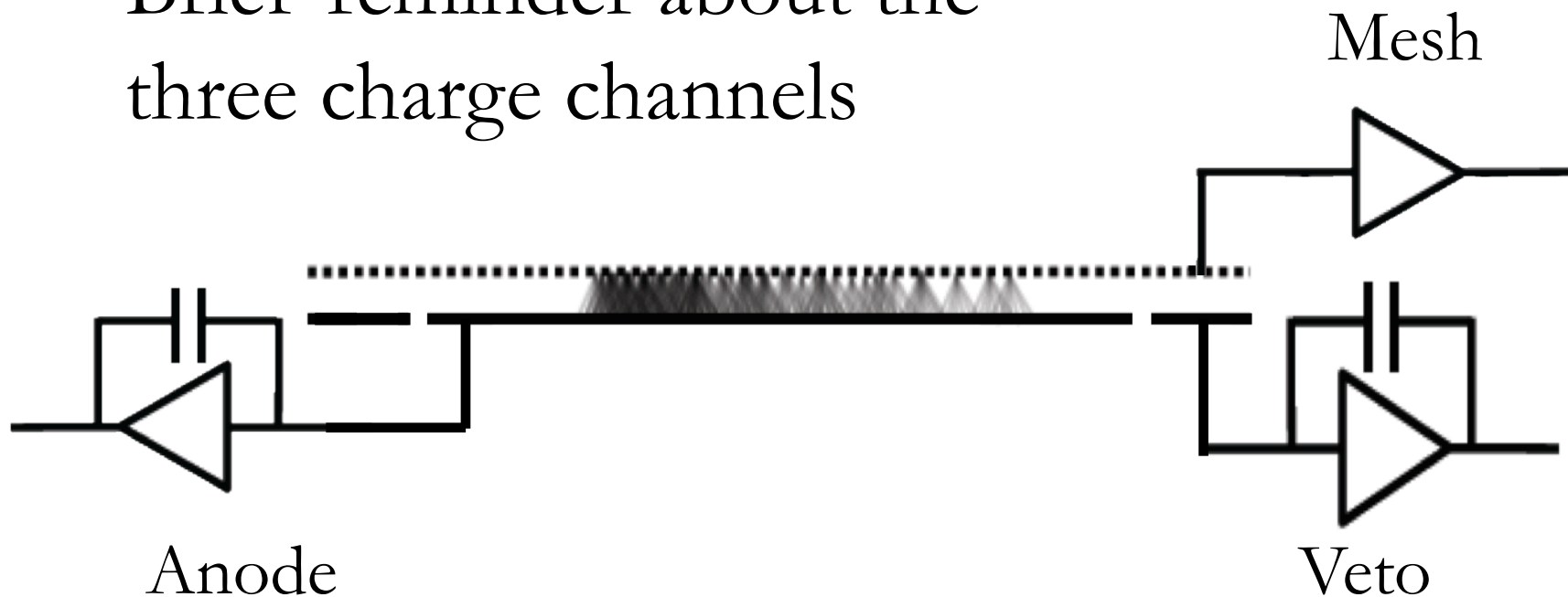
DmtpcMath::lineSegmentConvolvedWithGaussian2D



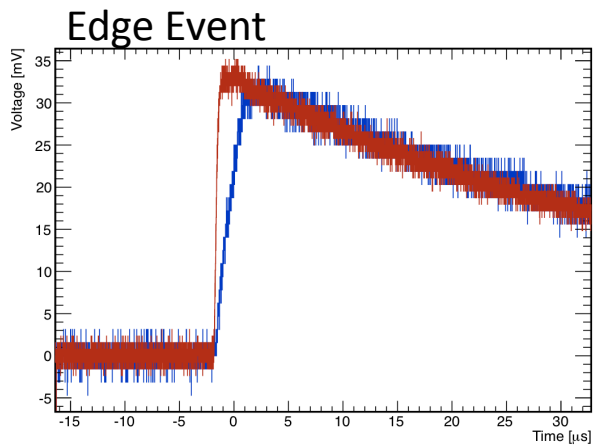
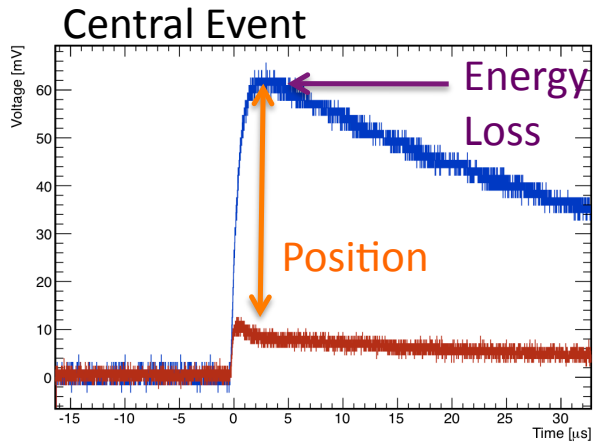
Nuclear recoil model

Charge readout: going beyond energy reconstruction

Brief reminder about the three charge channels

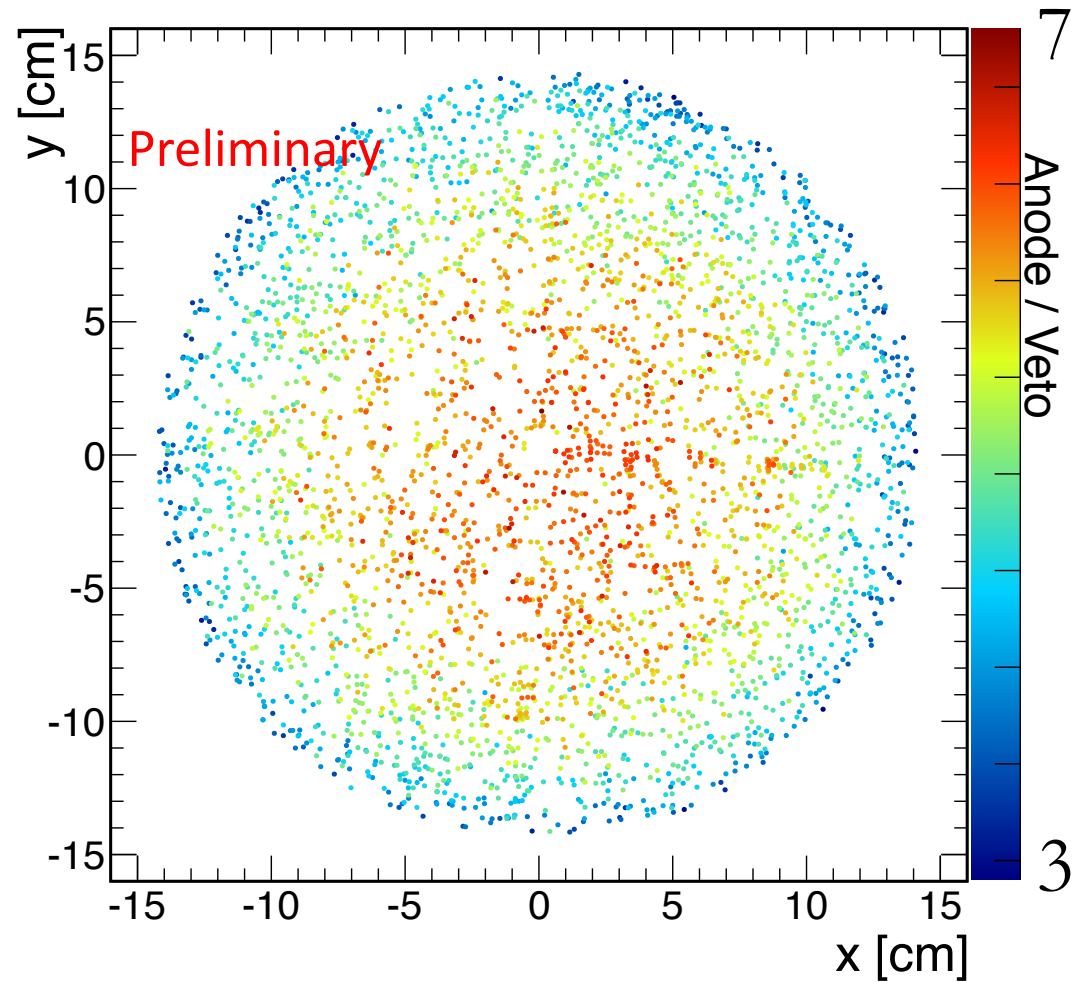


x-y position information from charge alone (useful in ccd/charge matching)

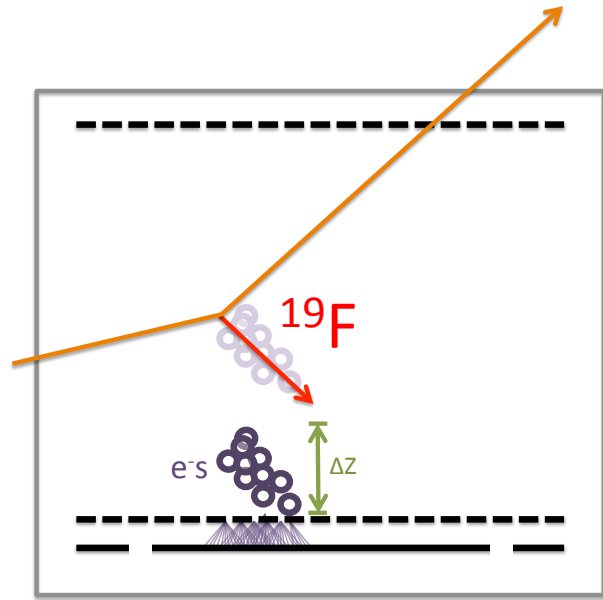


Central anode channel

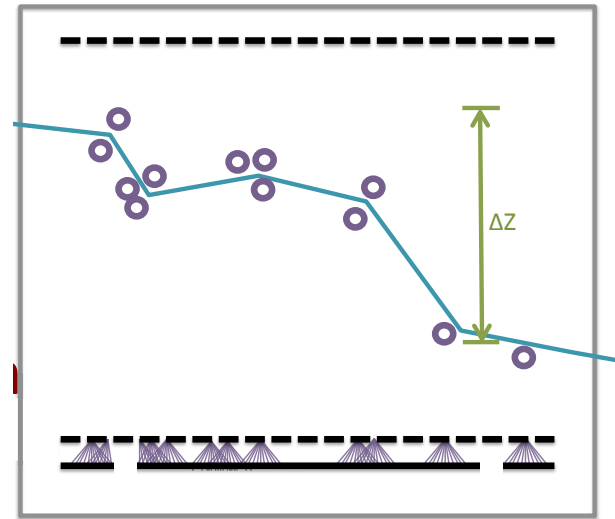
Veto channel



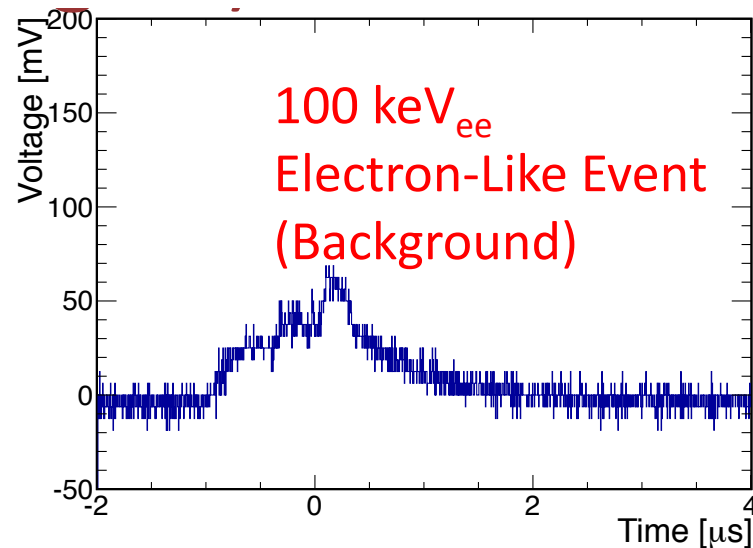
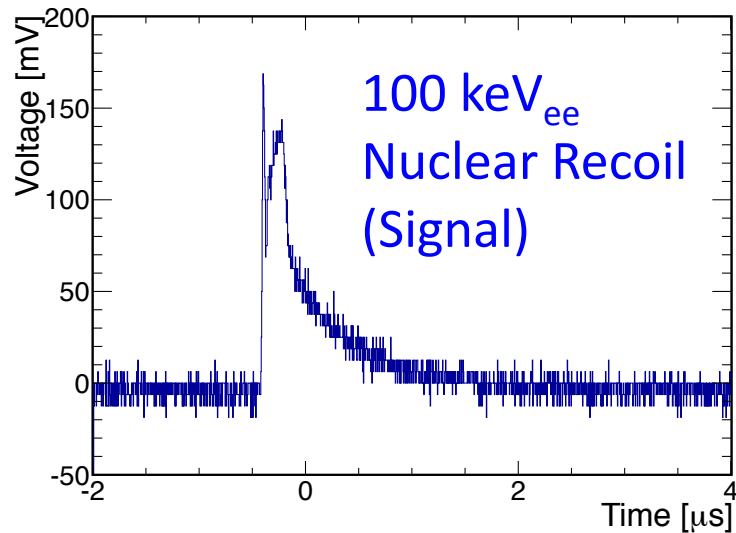
Event discrimination based on mesh pulse shape



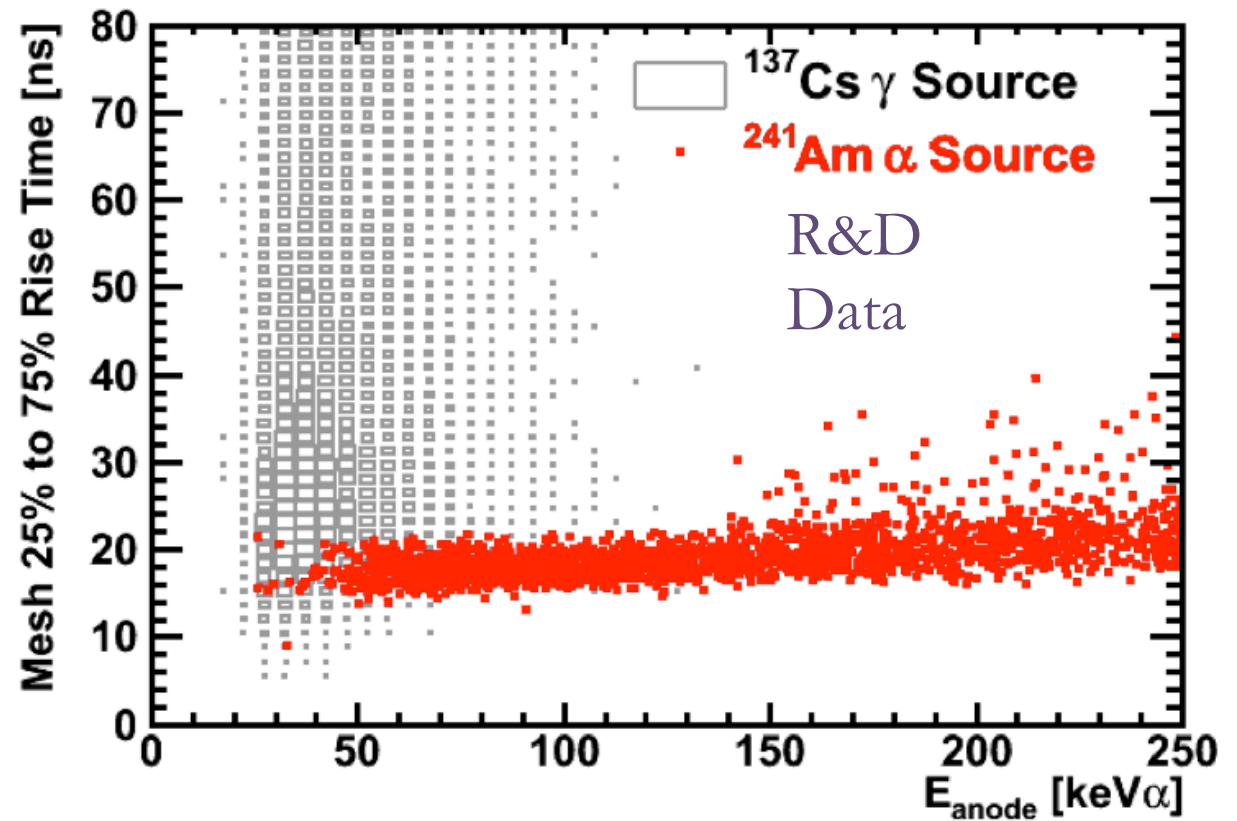
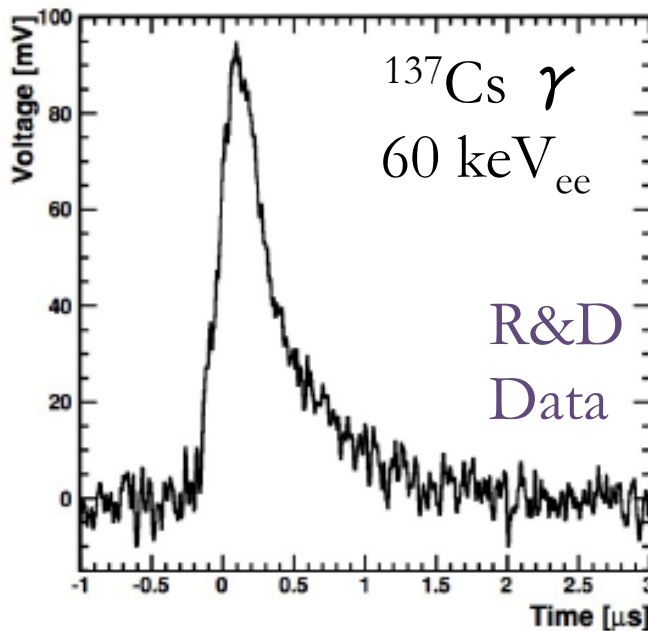
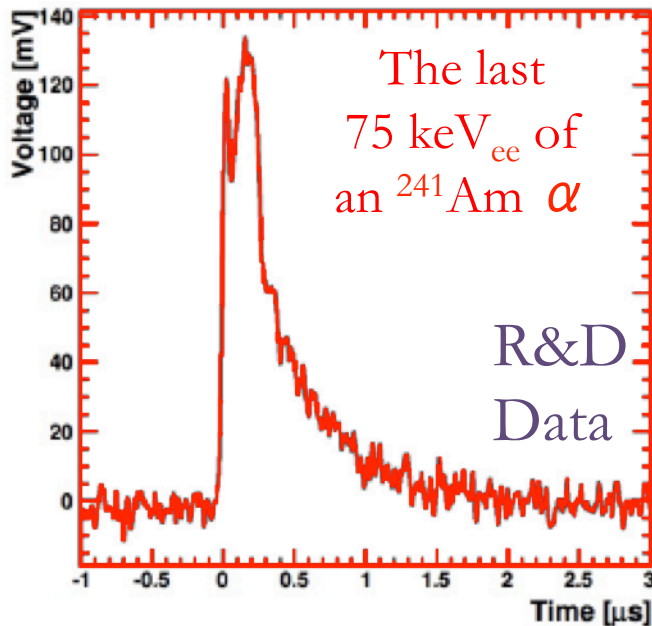
Nuclear recoil



Electron recoil

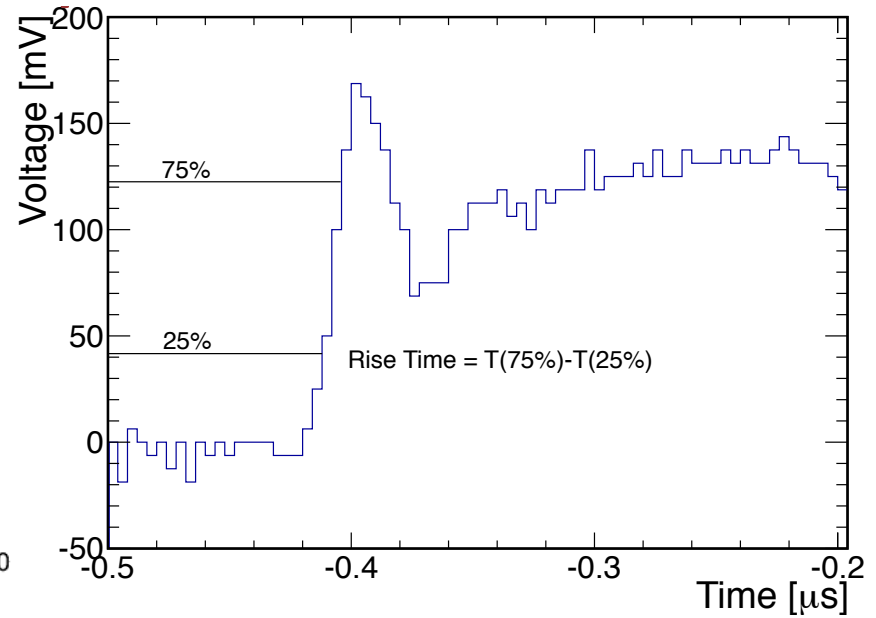
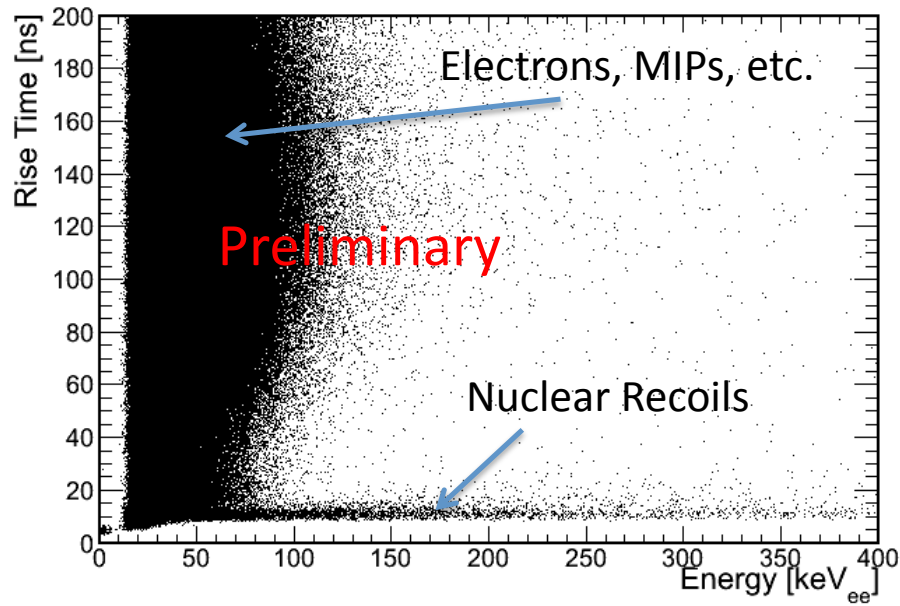


PID with Mesh Readout (different detector)

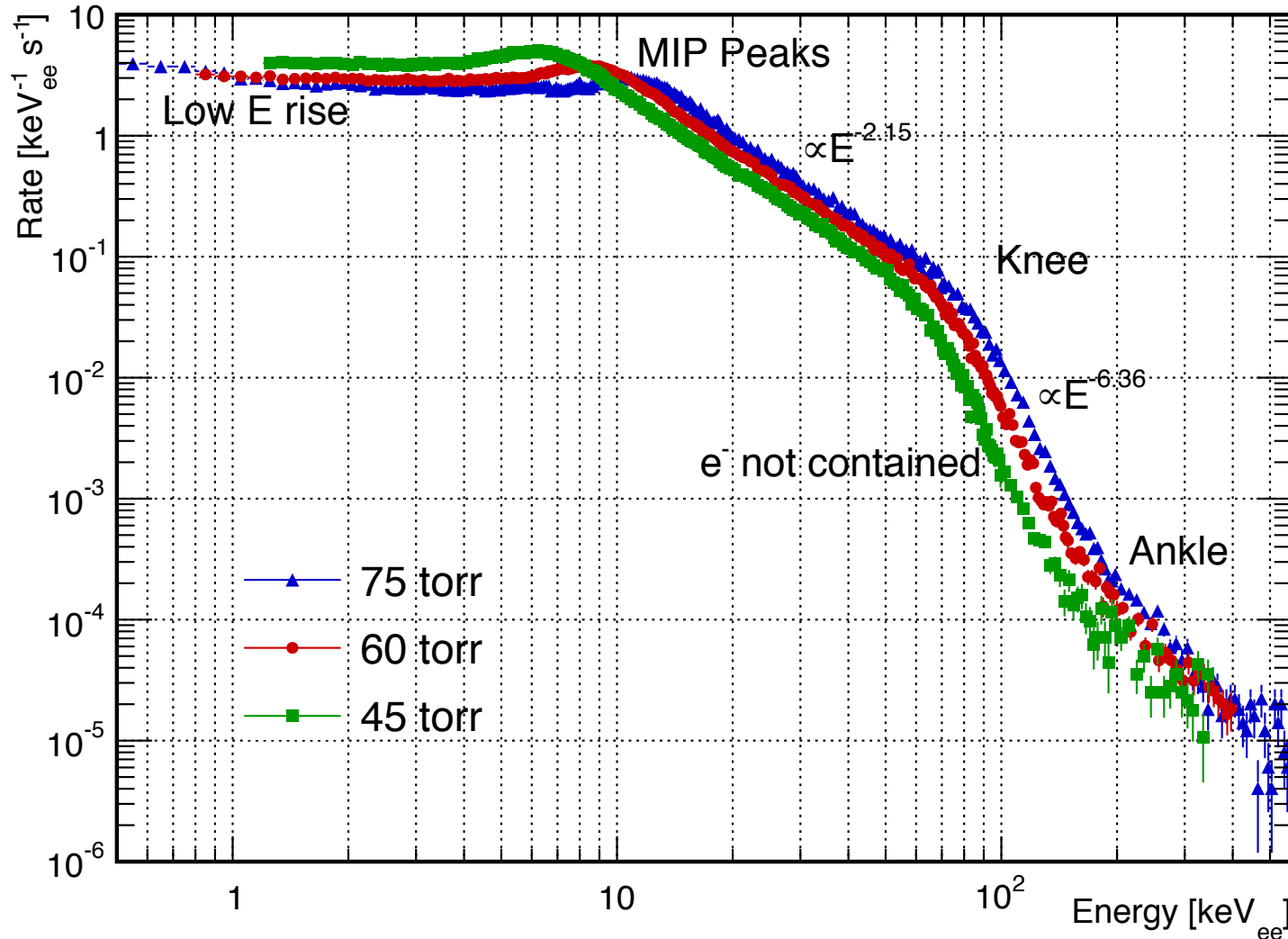


Demonstrated rejection of ¹³⁷Cs γ's
between 40 keV_{ee} < E < 200 keV_{ee} of
10⁵ (90% CL upper-limit) using
CCD+veto+mesh

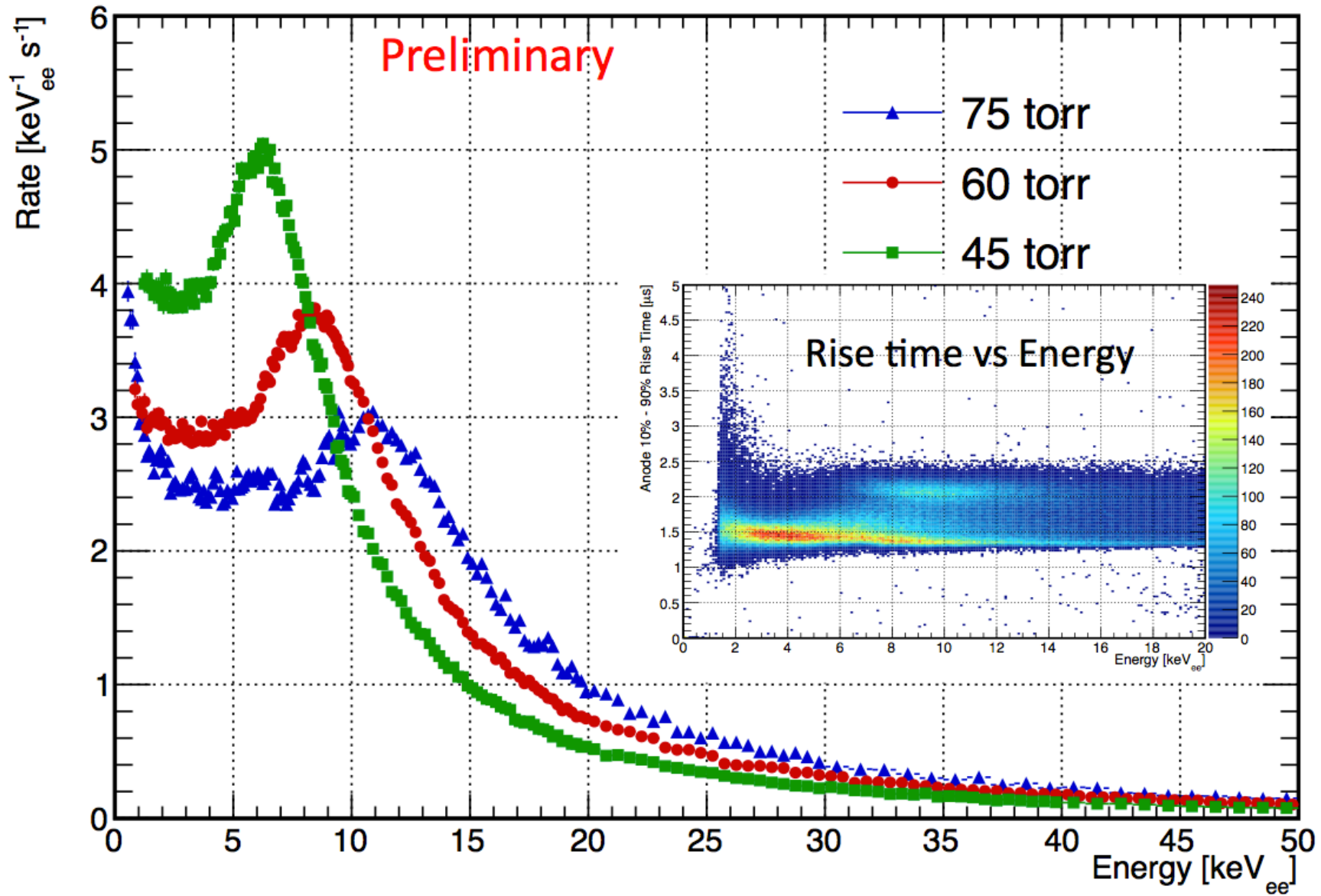
4sh data looks even better (though analysis ongoing)



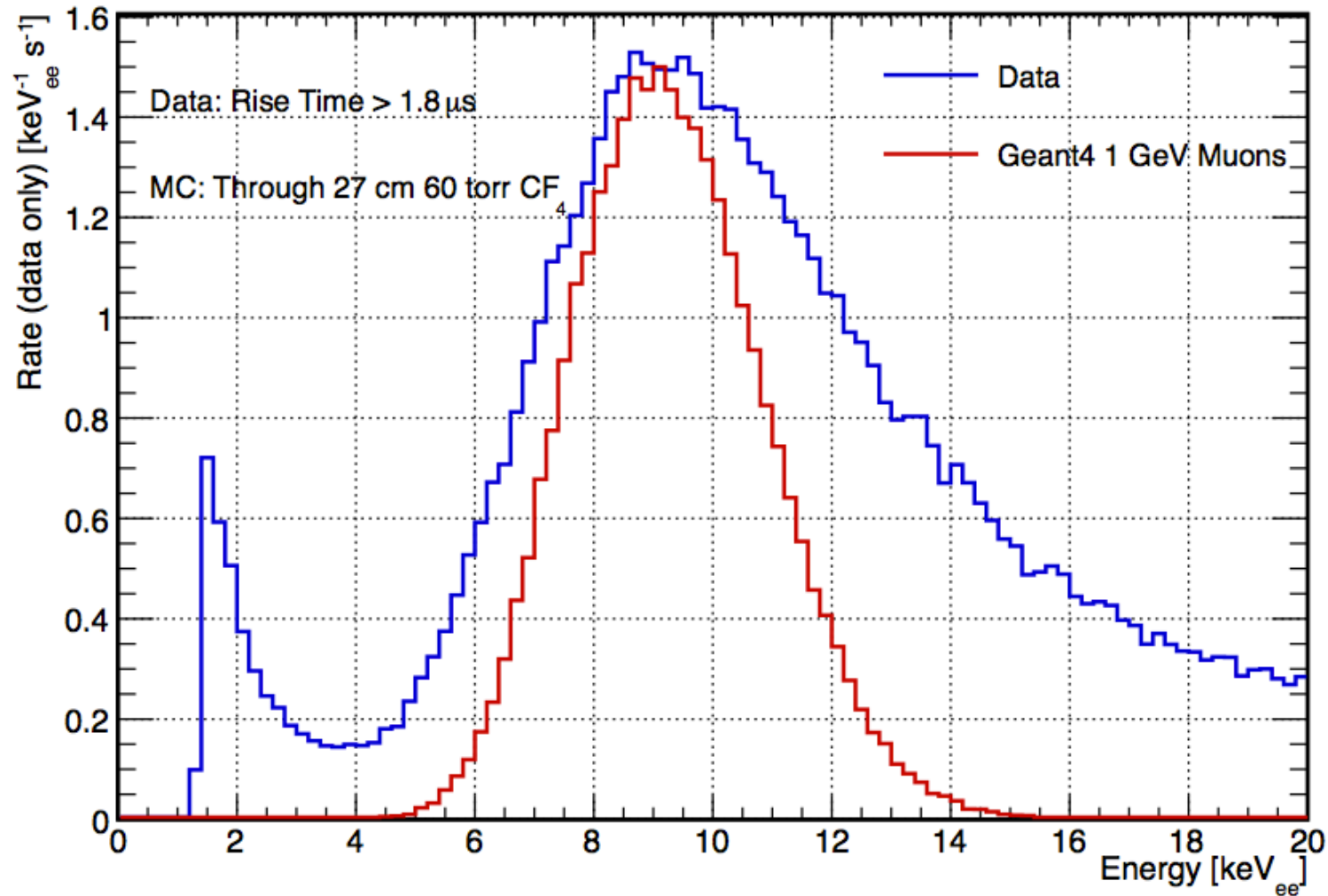
Charge background spectra



Charge background spectra



See MIPs in high rise-time events



Data expected to be wider due to detector resolution, angular distribution, and possible multi-particle events. Simulation is vertical only.

Nuclear recoil directional sensitivity
analysis underway (with AmBe source)

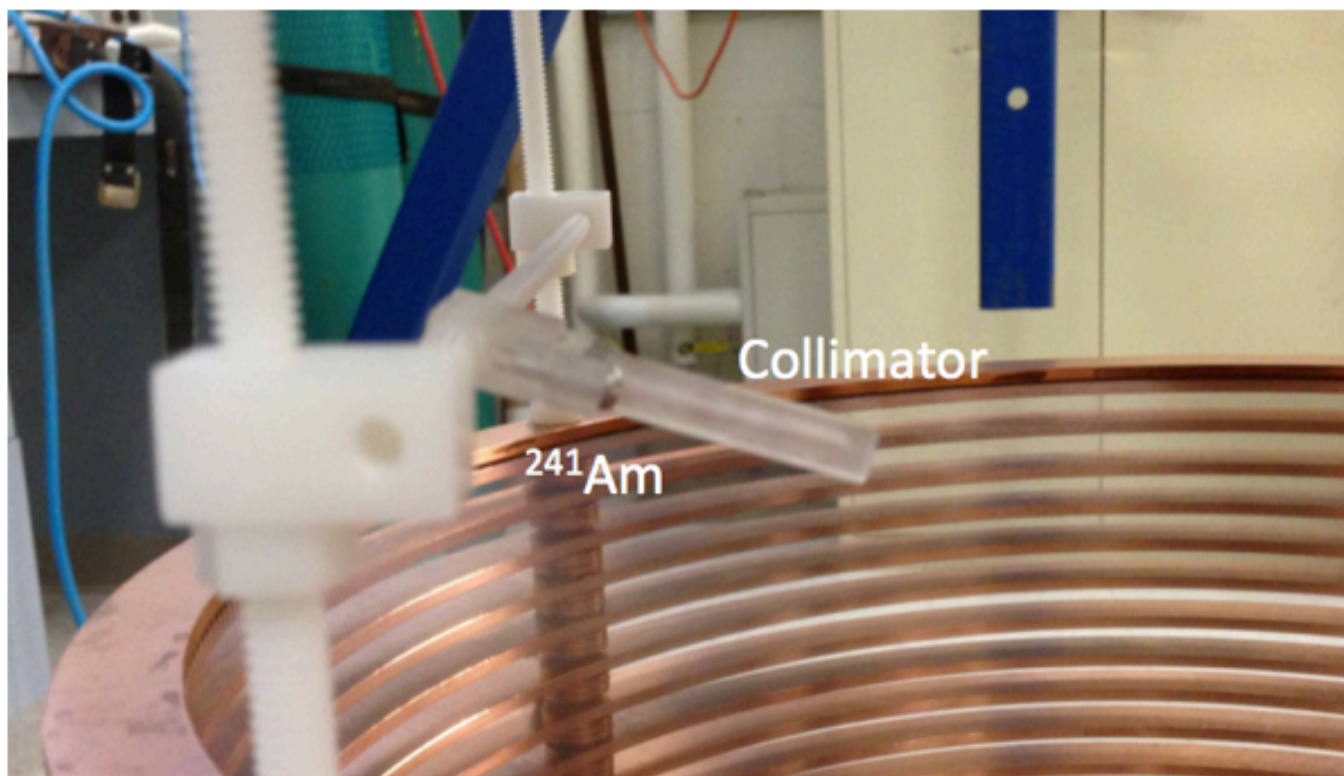
**An Assessment of the Sensitivity of a Low Pressure
Time Projection Chamber to the Direction of
WIMP-Induced Nuclear Recoils**

by

Shawn Wesley Henderson

(MIT Ph. D. thesis to be submitted in August)

Direction reconstruction at low energy

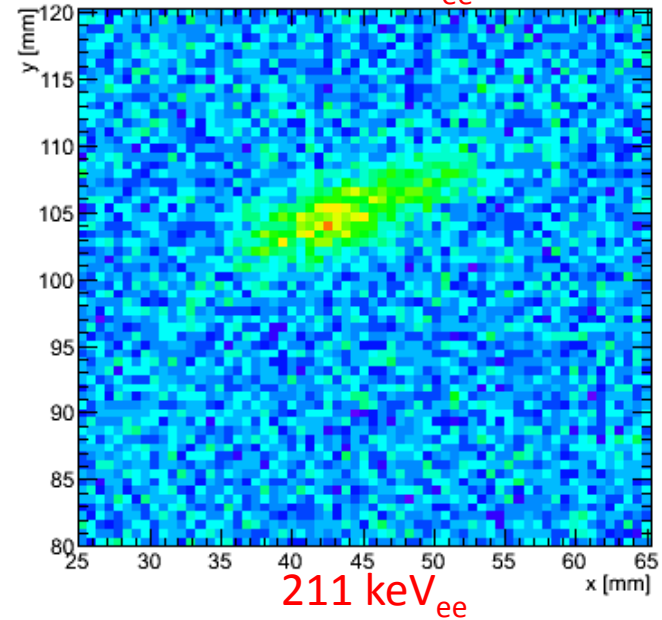
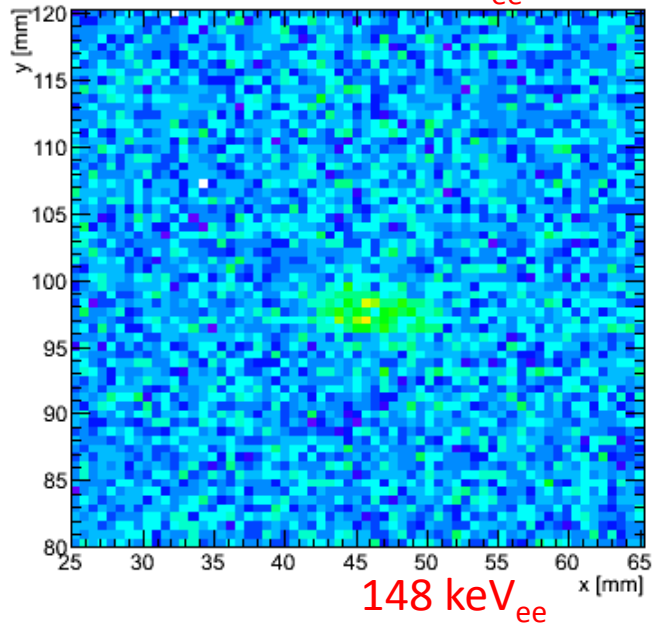
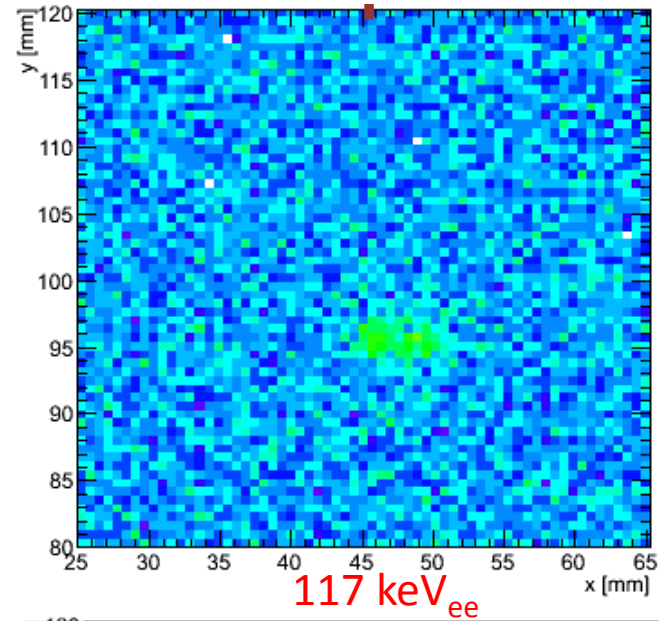
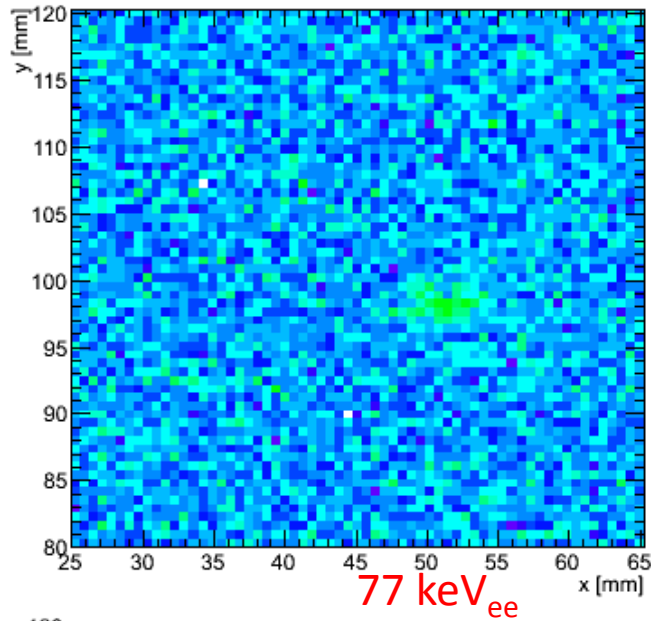


All tracks drift full length

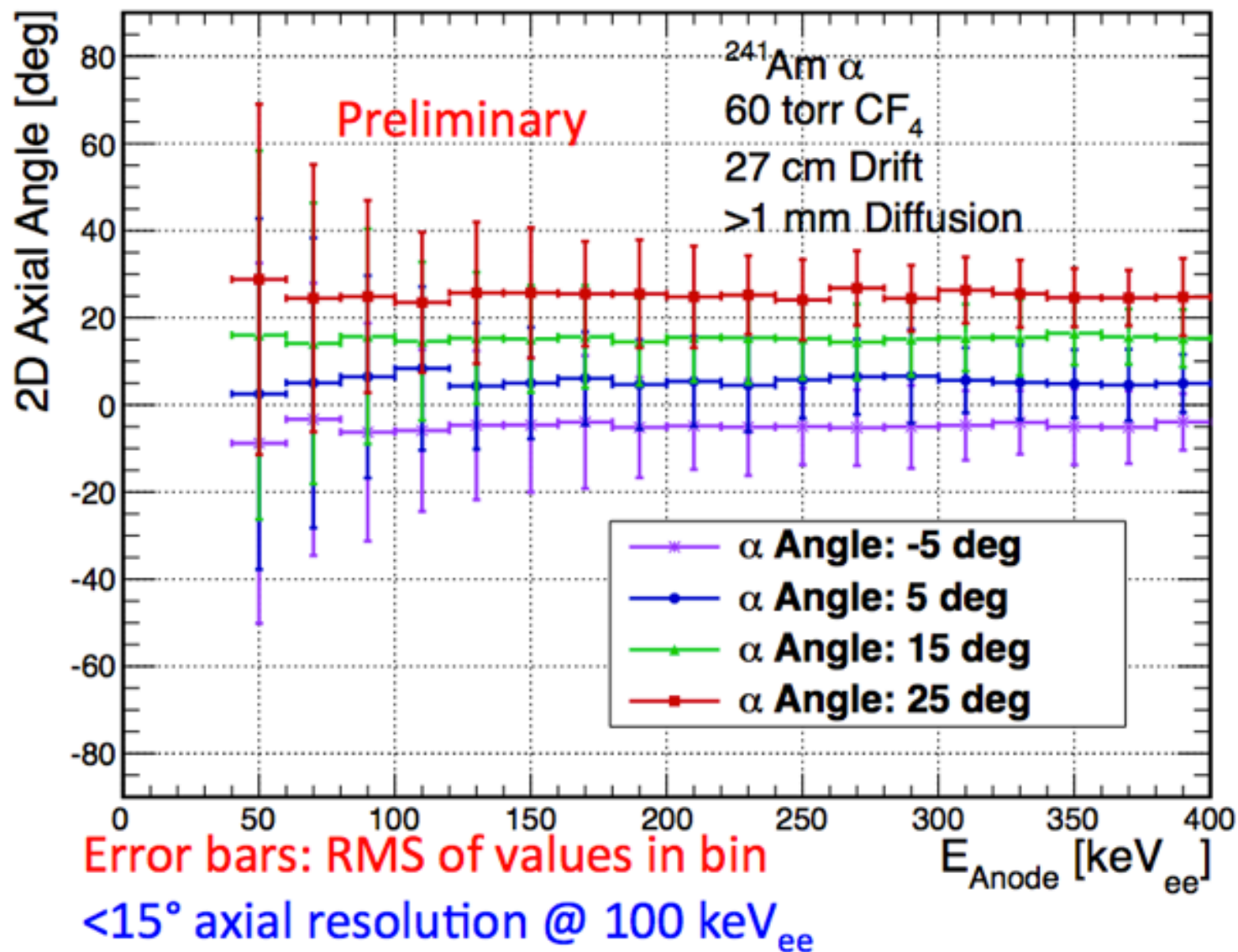
Smaller dE/dx than NR

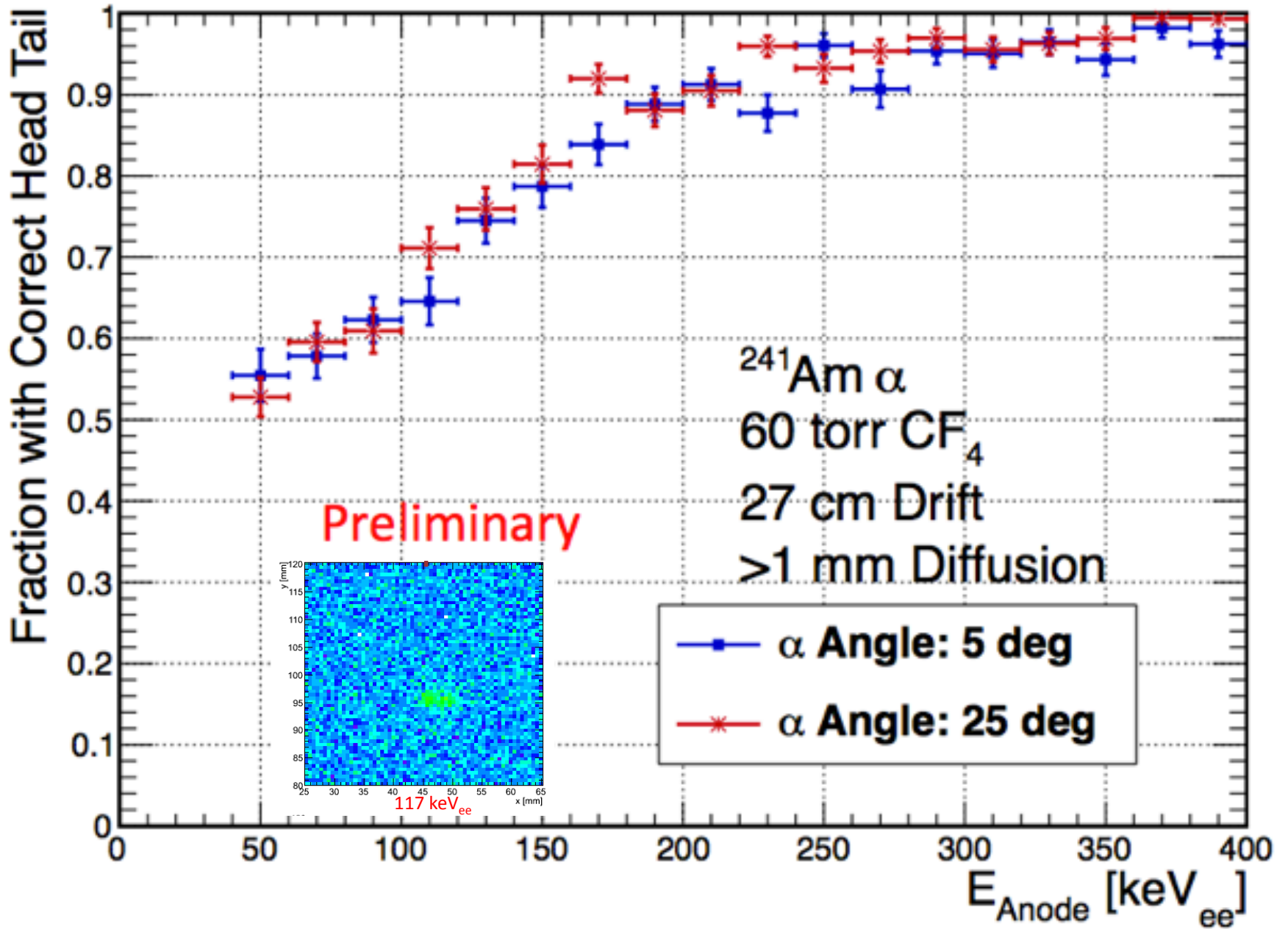
Only the tail end of the alpha track makes it into the active region → low energy recoils.

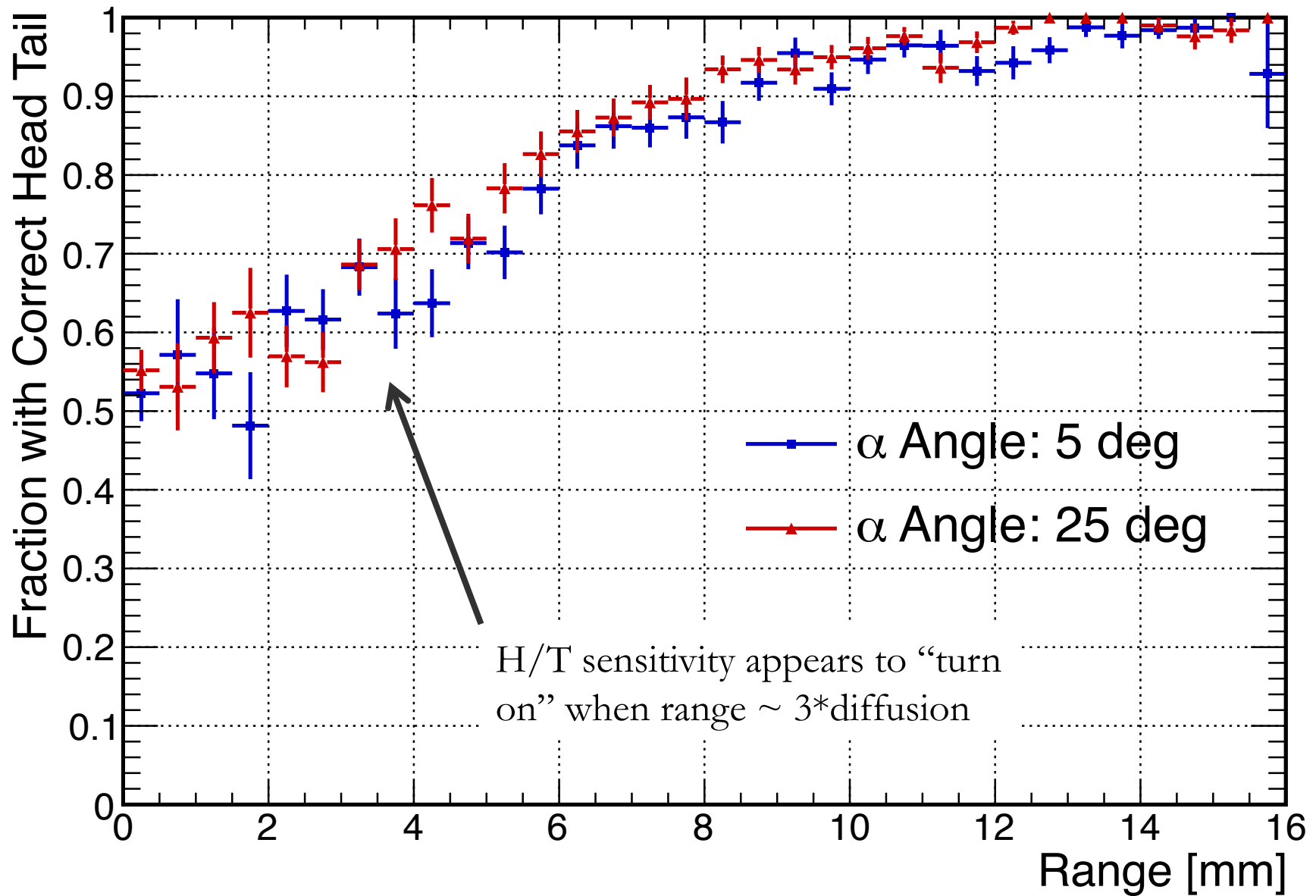
Sample low-energy alpha track images

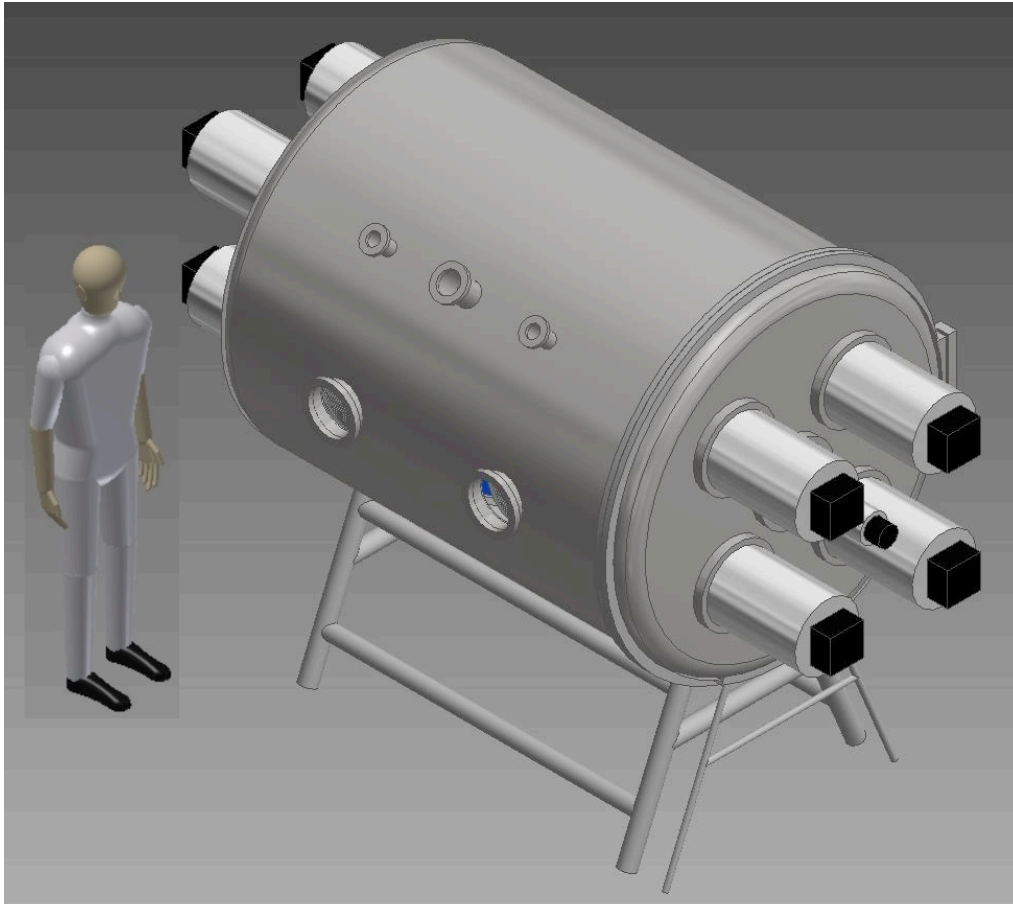


Directionality with low-energy alpha tracks

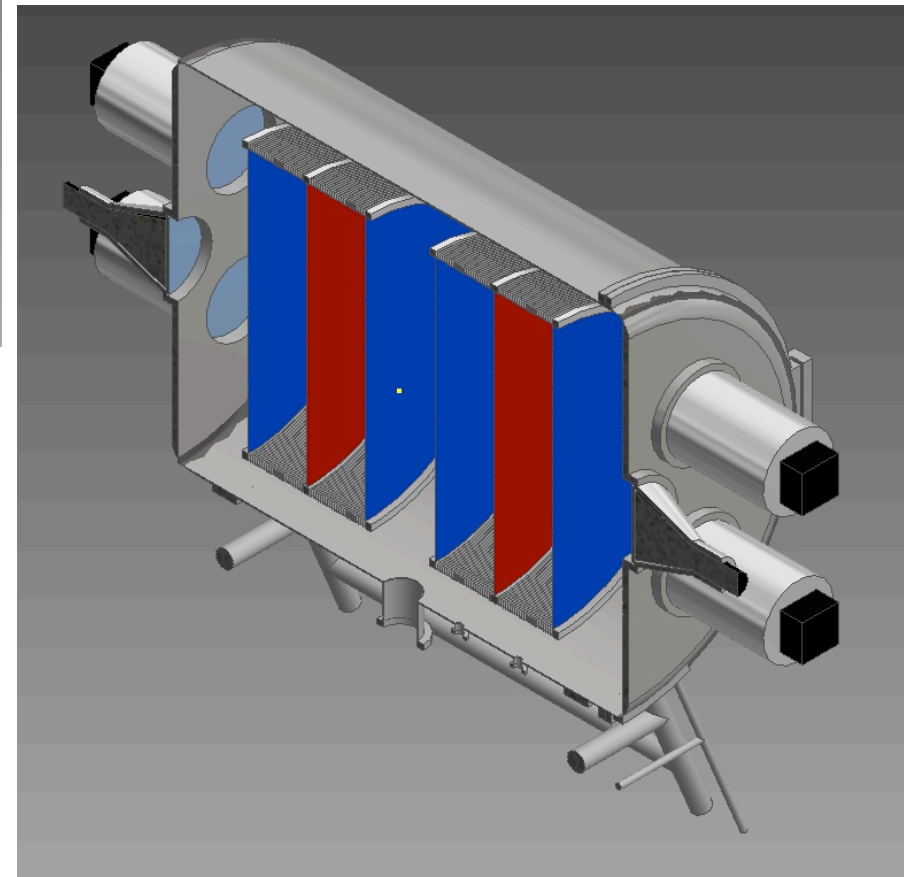








DMTPCino (1 m³)



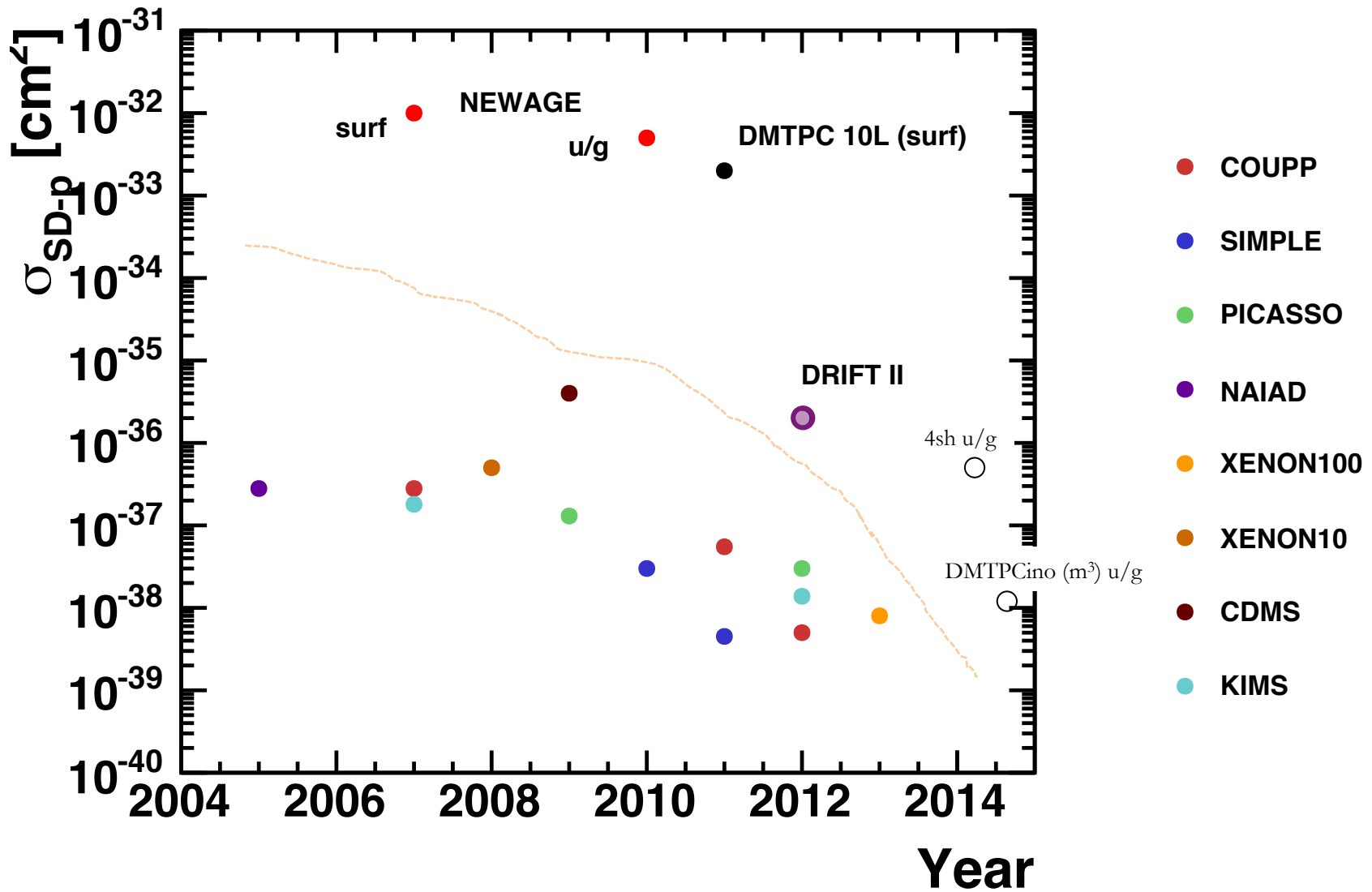
Amplification region = triple-mesh

One camera images two TPCs

Detector will fit in existing underground laboratory at WIPP

Triple mesh prototype built and under test now.
Vessel fabrication expected in the Fall.

Spin-dependent (proton) limits vs. time



Thank you