

First results from LiF bolometer at Kamioka

(astro-ph/200204441, to appear in Astropart. Phys)

2002 Sep. 5th @York

Department of physics , School of Science, University of Tokyo

Kentaro Miuchi

Group Members

M. Minowa

Y. Inoue^A, W. Ootani^A, Y. Ootuka^B, Y. Shimizu,

H. Sekiya, A. Takeda, K. Miuchi

University of Tokyo, ICEPP^A, Tsukuba University^B



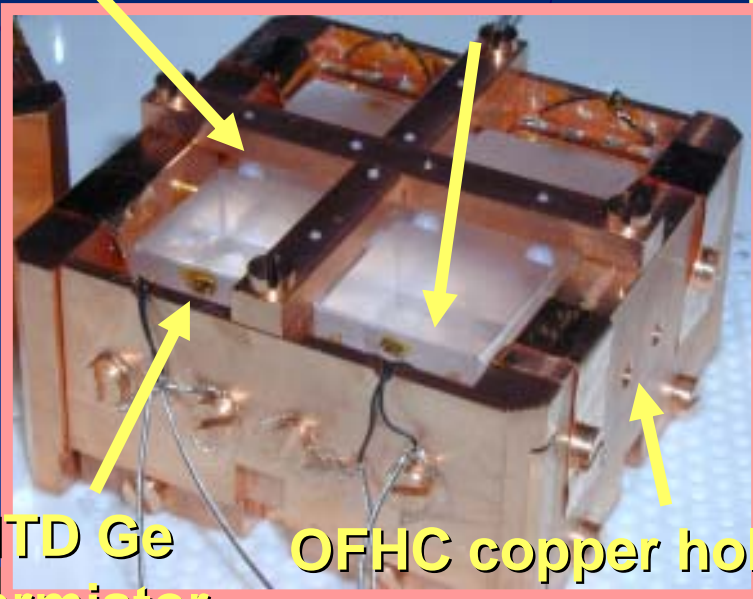
1 Detector

1-1 LiF Bolometer array

- For SD WIMPs
- 8pcs x 21g LiF crystals ($2 \times 2 \times 2\text{cm}^3$)

delrin balls

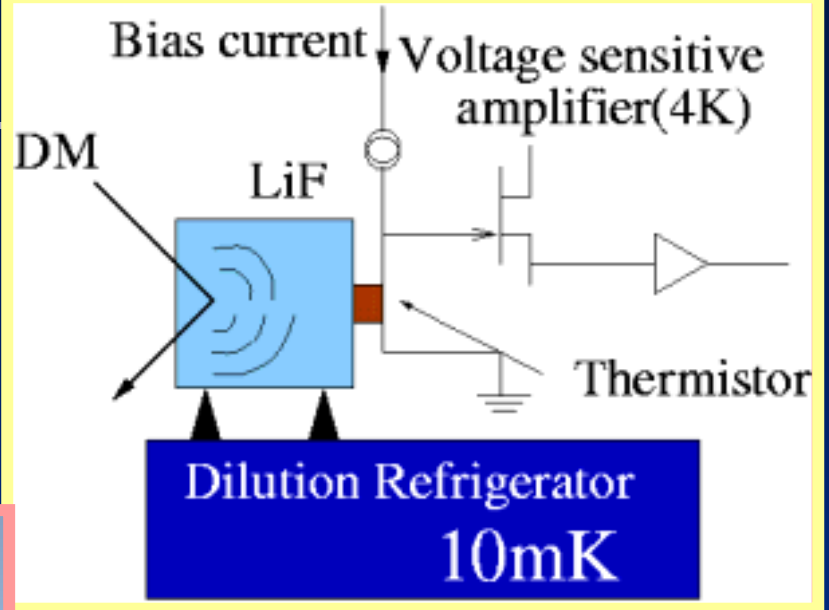
LiF crystal



NTD Ge thermistor OFHC copper holder

LiF bolometer array (1 of 2 stages)

schematics of the bolometer

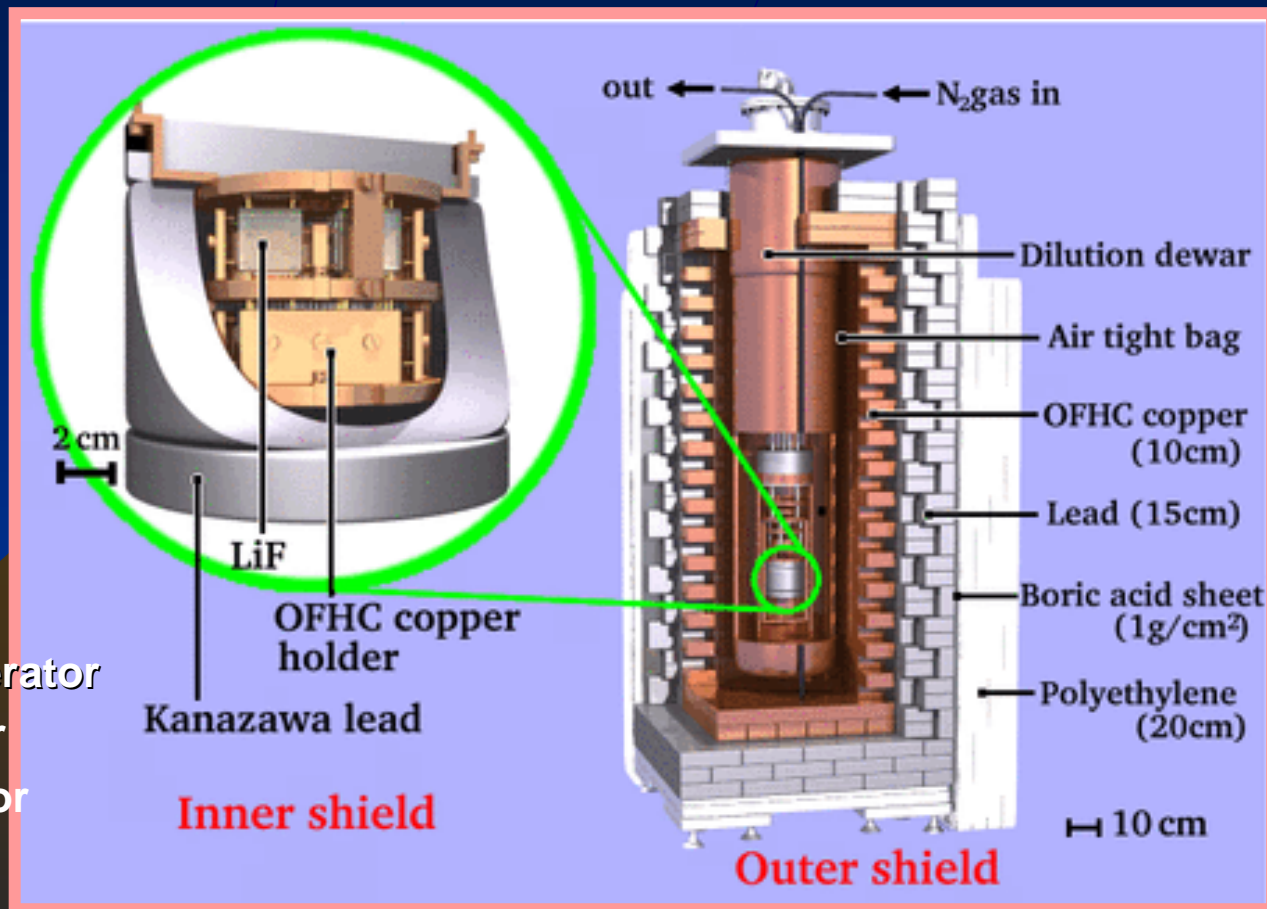


Spin factors of the materials used for DM search

| Isotope | abundance | $^2J(J+1)$ |
|-------------------|-----------|------------|
| ^7Li | 92.5% | 0.411 |
| ^{19}F | 100% | 0.647 |
| ^{23}Na | 100% | 0.041 |
| ^{73}Ge | 7.8% | 0.065 |
| ^{127}I | 100% | 0.023 |
| ^{129}Xe | 26.4% | 0.124 |

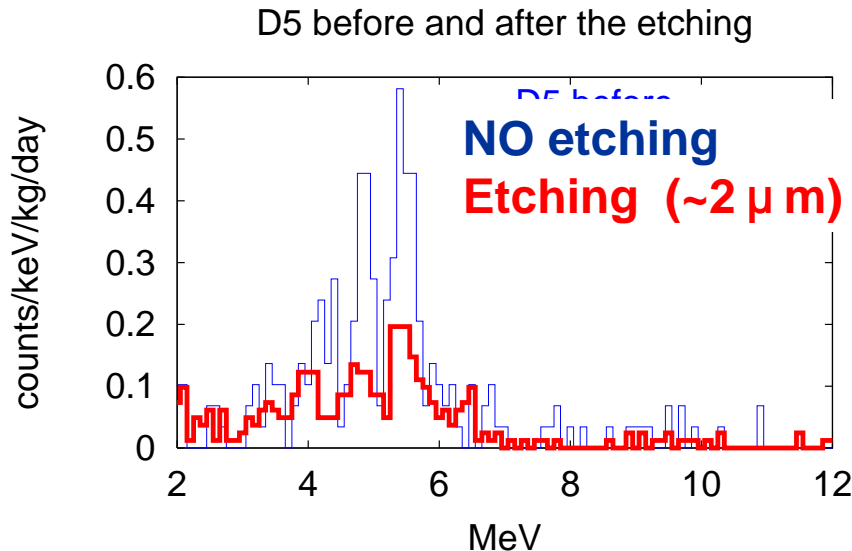
1-2 Detector set up

- Kamioka Mine (~2700m.w.e)
- **Shielding**
- **Cryogenics**
 - Dilution refrigerator
 - Helium liquefier
 - Liq. N₂ generator
- **Bolometer array**



2 Recent improvements

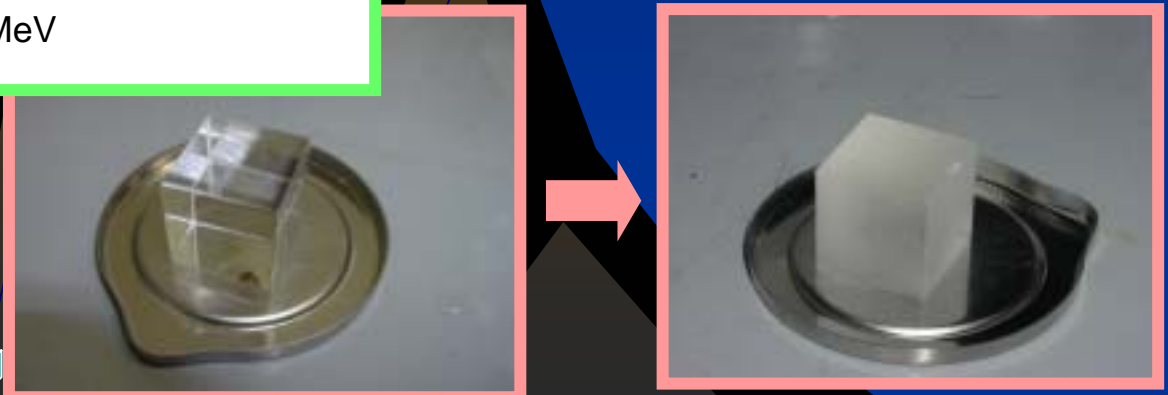
2-1 Chemical etching of the LiF crystals



– Results of the preliminary run at Kamioka

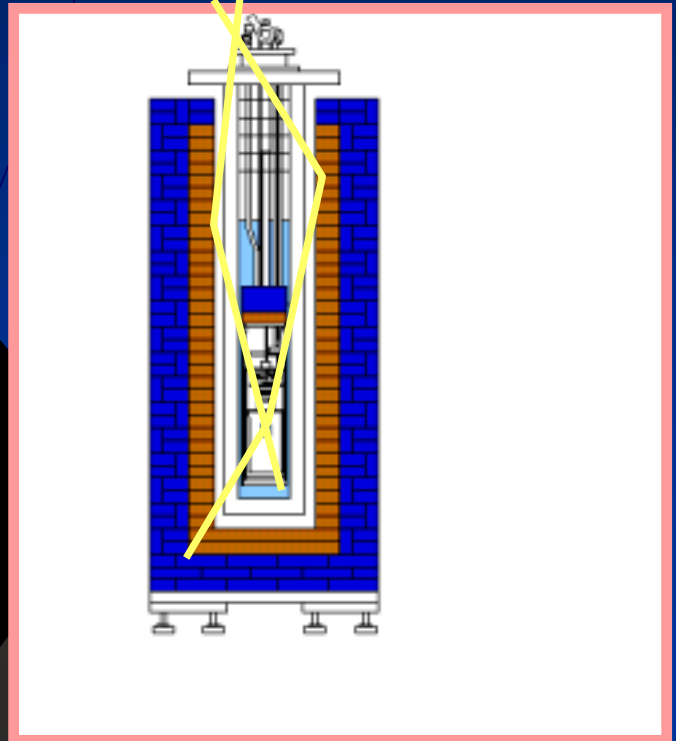
– peaks are suppressed by etching the surface ($\sim 2 \mu\text{m}$) with perchloric acid

– $20 \mu\text{m}$ etchings for this measurement.



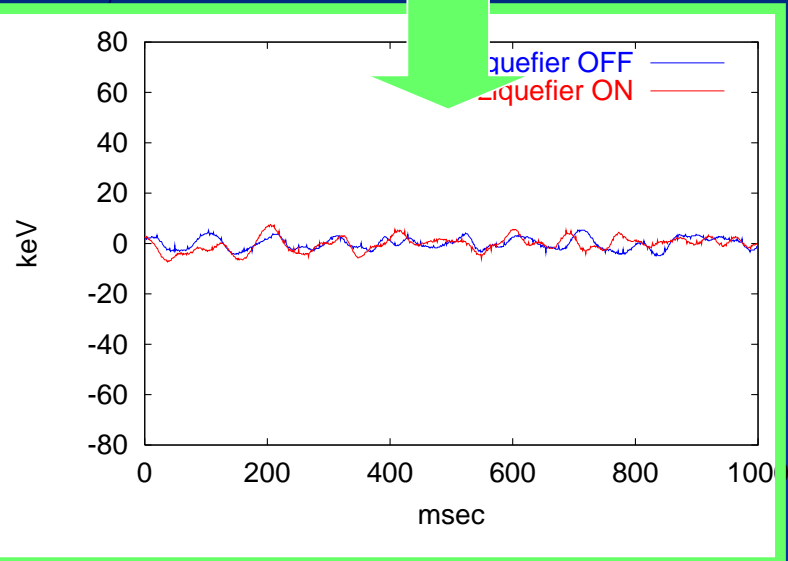
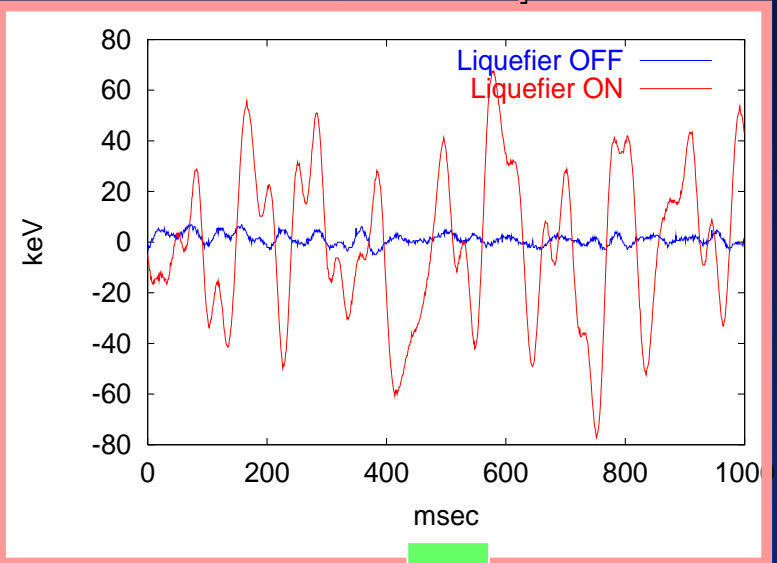
2-2Inner shield

- Back-scattered γ -rays would be the serious BG source
 - install 2cm lead shield in the cold stage
- Problem: ^{210}Pb (half life=22.3 y)
 - old lead (>200 y) from Kanazawa castle



2-3 Eliminating the microphonic noise

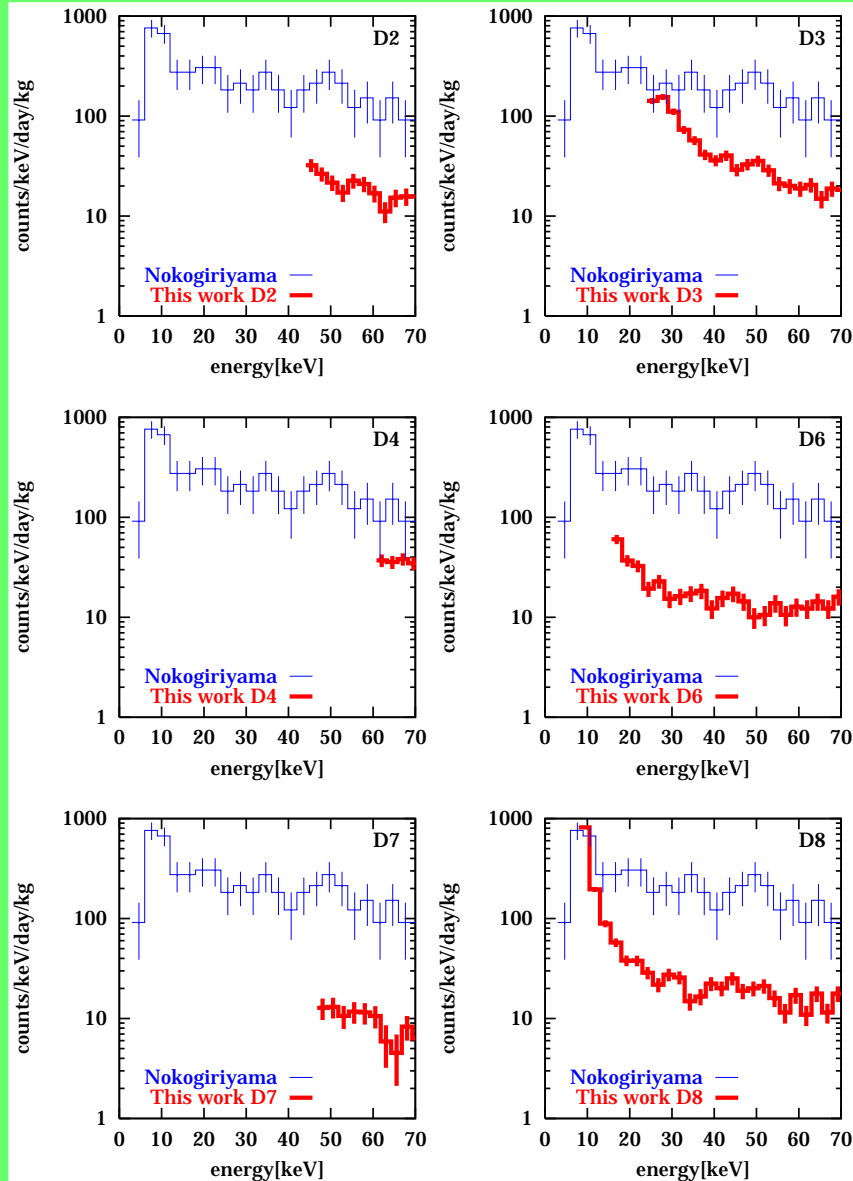
He liquifier noise



3 Measurement

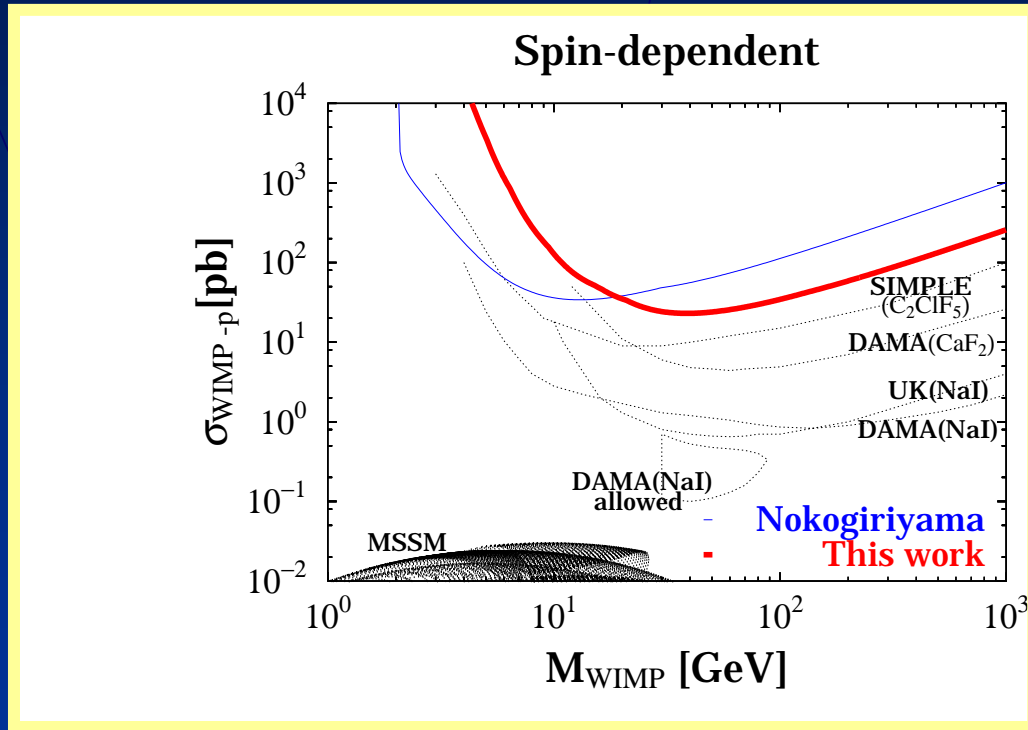
3-1 Run data

- Total exposure: **4.1 kg·days** (50 days)
- **6 of 8 detectors** was used for the analysis
- **10-30 counts/keV/kg/day**
- **BG rate : × 1/10** from the pilot run



3-2 SD WIMP-p limits

- Assumption : all events are WIMP- proton events.



| | |
|--------------------|--------------------------|
| DM | 0.3 GeV cm ⁻³ |
| v ₀ | 220 km s ⁻¹ |
| v _{esc} | 650 km s ⁻¹ |
| v _{Earth} | 217 km s ⁻¹ |

Astrophysical parameters

- Limits to $M_{WIMP} > 30\text{GeV}$: $\times 10$ improved (from the pilot run)
- Best limits SD : 23 pb (for 40GeV)



3-3 SD limits in the a_p - a_n plane

– SD WIMP- N cross section

$$\sigma_{\text{WIMP-N}}^{\text{SD}} = 4G_F^2 \mu_{\text{WIMP-N}}^2 \left(a_p \langle S_{p(N)} \rangle + a_n \langle S_{n(N)} \rangle \right)^2 \frac{J+1}{J}$$

(D.R. Tovey et. al. Phys. Lett. B 488(2000)17)

(contributions of both proton and neutron are considered)

a_p, a_n : WIMP-nucleon couplings

To be measured

$\langle S_{p(N)} \rangle$: proton spin contribution in the nucleus

| Isotope | unpaired | $\langle S_{p(N)} \rangle$ | $\langle S_{n(N)} \rangle$ |
|---------------------|----------|----------------------------|----------------------------|
| ${}^7\text{Li}$ | p | 0.497 | 0.004 |
| ${}^{19}\text{F}$ | p | 0.441 | -0.109 |
| ${}^{23}\text{Na}$ | p | 0.248 | 0.020 |
| ${}^{73}\text{Ge}$ | n | 0.009 | 0.372 |
| ${}^{127}\text{I}$ | p | 0.309 | 0.075 |
| ${}^{129}\text{Xe}$ | n | 0.028 | 0.359 |

${}^{19}\text{F}$ has opposite sign of

$\langle S_{p(N)} \rangle / \langle S_{n(N)} \rangle$

compared to ${}^{23}\text{Na}, {}^{73}\text{Ge}, {}^{127}\text{I}$

→ COMPLEMENTARY

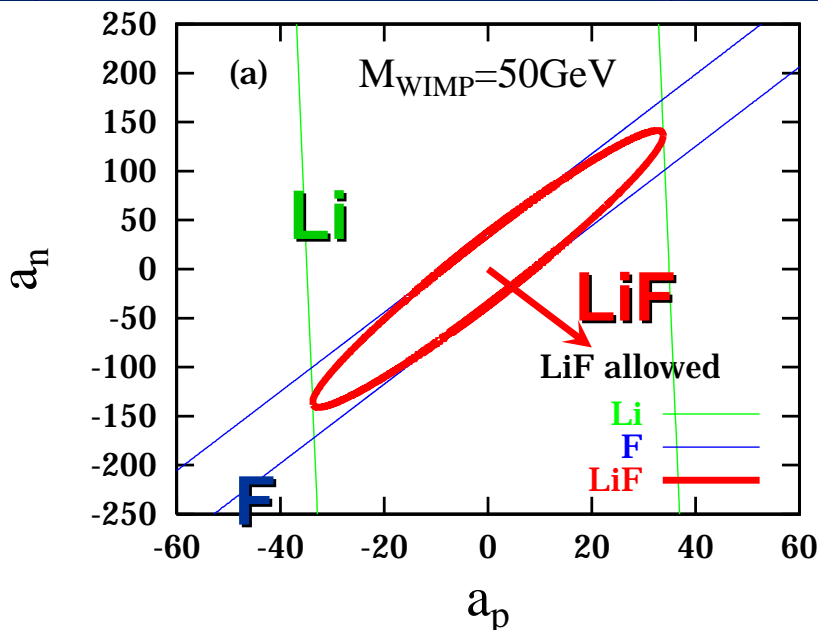
for a_p, a_n determination

$\langle S_{p(N)} \rangle$ and $\langle S_{n(N)} \rangle$ values
(calculated by shell models)

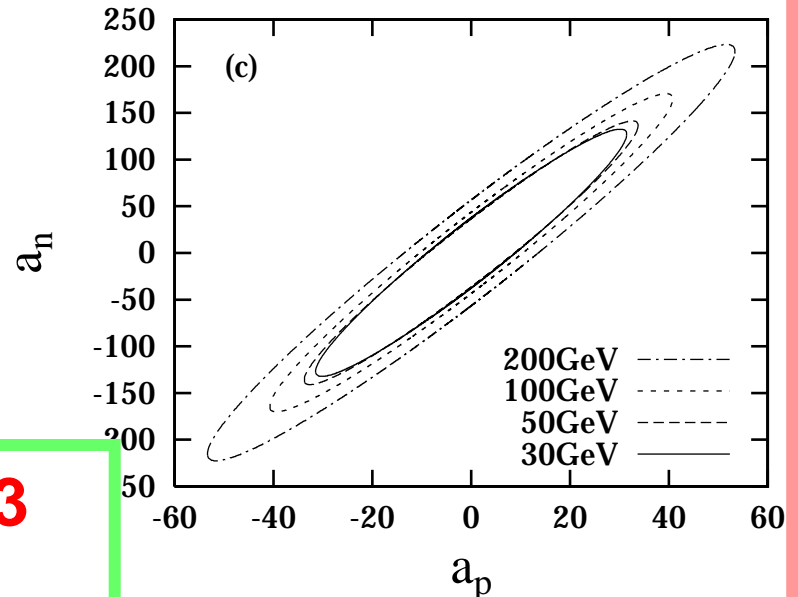
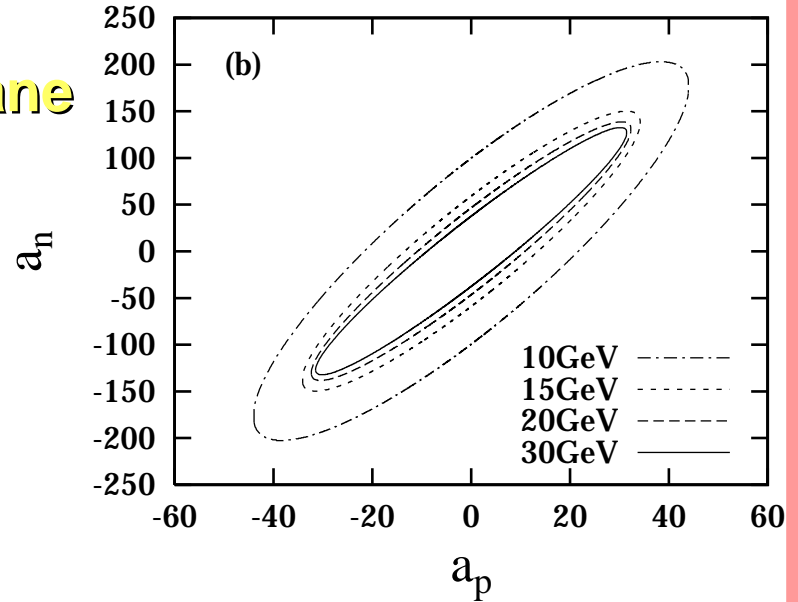
SD limits in the a_p - a_n plane

- Derived the a_p - a_n limits from the cross section limits.

$$\sigma_{\text{WIMP-N}}^{\text{SD}} \propto \left(a_p \langle S_{p(N)} \rangle + a_n \langle S_{n(N)} \rangle \right)^2 \frac{J+1}{J}$$



$|a_p| < 32$, $|a_n| < 133$
for 30GeV WIMP

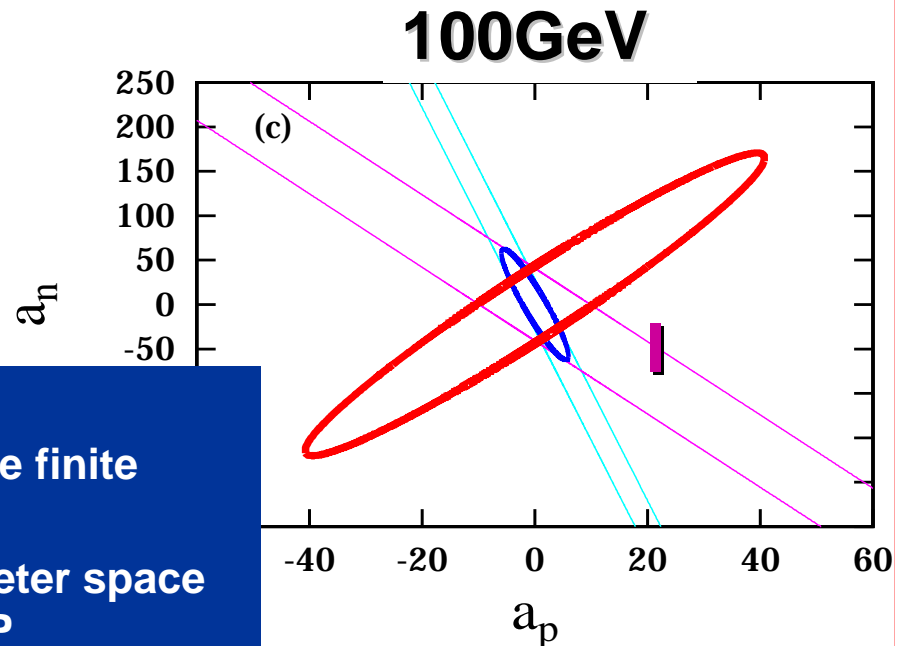
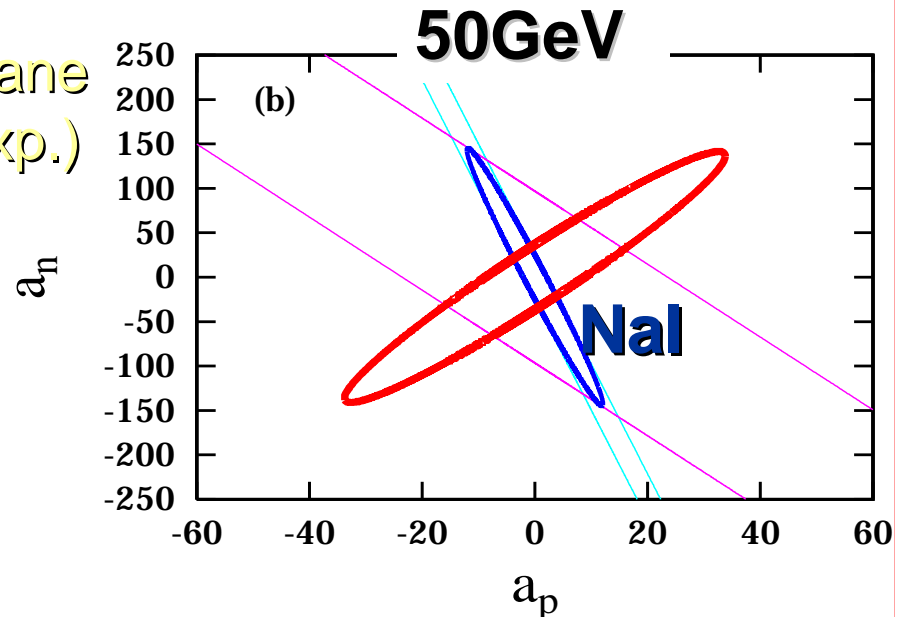
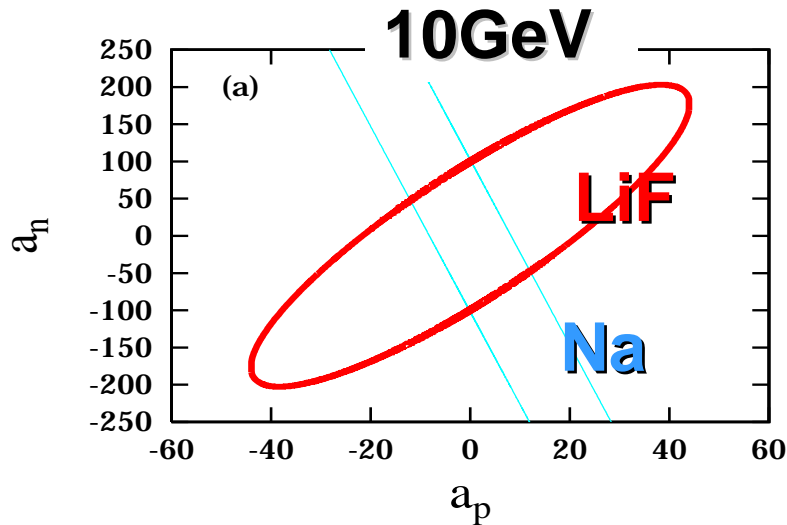


□ SD limits in the a_p - a_n plane
(comparison with other exp.)

– Comparison with the UKDMC

LiF sets the limits

COMPLEMENTARY to NaI



– This is the **only experiment** to set the finite limits for 10 GeV WIMP

– **Killed more than 2/3** of the parameter space allowed by UKDMC for 50 GeV WIMP

4 Discussions, prospects

□ Main background

- ^{40}K contamination in LiF crystals
- U, Th in the holder and Kanazawa lead
- ^3H from neutron capture of ^6Li

□ prospects

- Etching the crystals more deeply
- Ta bolometer as an active shield
- NaF bolometer
- Directional detectors (stilbene scintillator)

Y. Shimizu et. al. astro-ph/0207529

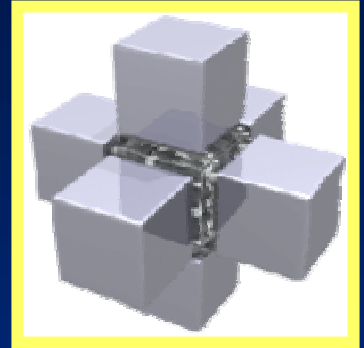
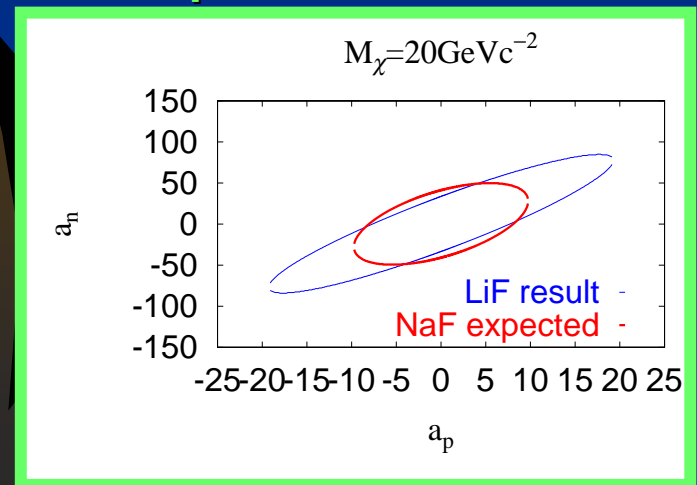


Image of Ta active shield

Expected limits with NaF





5 Conclusions

□ **LiF bolometers** with special care for

- low threshold
- low background

was used for the measurement.

□ SD limits in the **a_p - a_n plane** are shown.

□ A large part of the parameter space in the a_p - a_n plane for light WIMPs **allowed by UKDMC** experiment **was excluded** by this experiment.

