



# NEWAGE

(New generation WIMP search with an advanced gaseous tracking device experiment)

## Direction-sensitive direct dark matter search

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### 1. Dark Matter

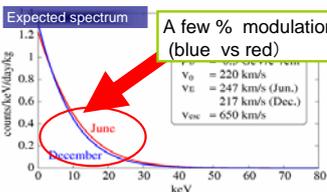


- Promising candidate of dark matter
  - WIMPs (neutralino)
  - Axions
  - Machos
- Dark Matter search
  - Direct search: Search for DM - nuclei elastic scatterings
  - Indirect search: Search for  $\gamma$ , antiproton, e+ ... from DM annihilations
  - Accelerator: Search for missing mass

### 2. "Conventional" Direct Dark Matter Search

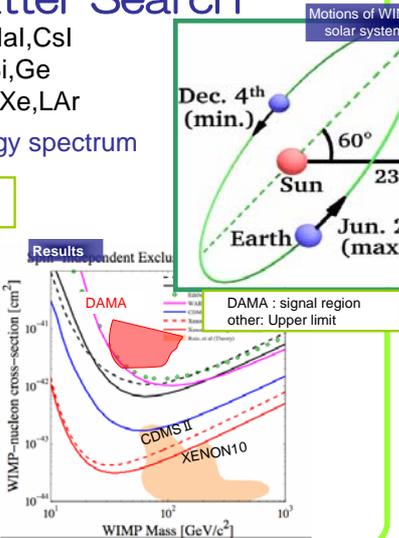
- Large mass
  - Low background
  - High energy resolution
- Nal, Csl  
Si, Ge  
LXe, LAr

Annual modulation in energy spectrum



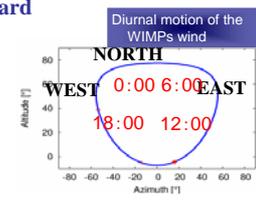
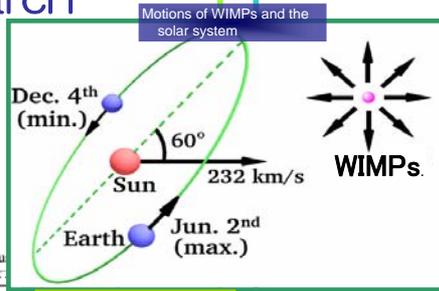
DAMA-signal was observed again. (EPJC56(2008)333)

Other reliable method is required.



### 3. Direction Sensitive Direct Dark Matter Search

- Detect the "WIMPs wind"!
- Recoil nuclei scattered forward



Forward events are 10 times larger than backward event

Very strong evidence

Simulation with neutron background in Kamioka mine

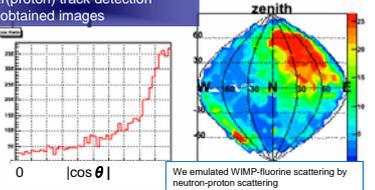
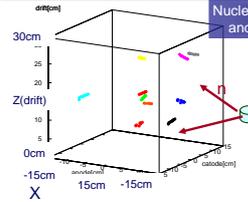
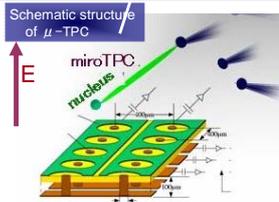
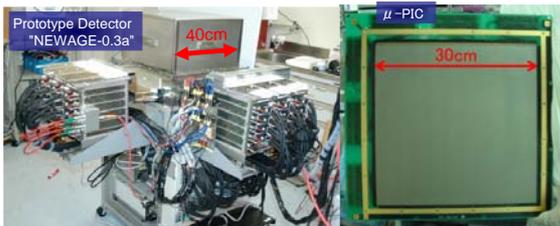
### 4. Detector

Astropart. Phys.31 (2009) 185, PLB 654(2007) 58 Detector Performance and Development plan

- $\mu$ -TPC: micro Time Projection Chamber (3D-tracking device) consists of
  - $\mu$ -PIC: (2D gas detector 400  $\mu$ m pitch gas gain  $\sim$  3000, 30  $\times$  30cm<sup>2</sup>) + 100MHz readout system

	Current	Plan
Detection Volume	30 $\times$ 30 $\times$ 31cm <sup>3</sup>	>1m <sup>3</sup>
Gas	CF <sub>4</sub> 152Torr	CF <sub>4</sub> 30 Torr
Energy threshold	100keV	35keV
Energy resolution(@ threshold)	70%(FWHM)	50%(FWHM)
Gamma-ray rejection(@threshold)	8 $\times$ 10 <sup>-6</sup>	1 $\times$ 10 <sup>-7</sup>
Angular resolution (@ threshold)	55 $^\circ$ (RMS)	30 $^\circ$ (RMS)

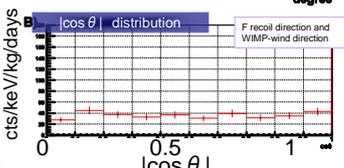
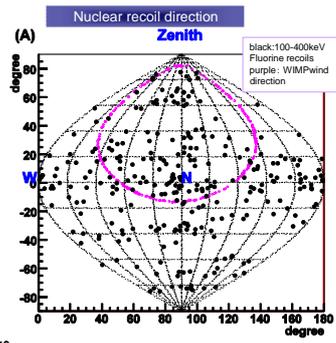
- Fluorine has large spin-dependent cross-section.
- Lower pressure gas provides
  - lower threshold
  - better angular resolution
- Uniform pixels provide
  - better energy resolution
  - better gamma-ray rejection



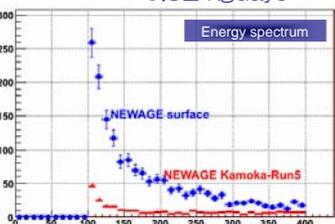
### 5. Underground Measurements



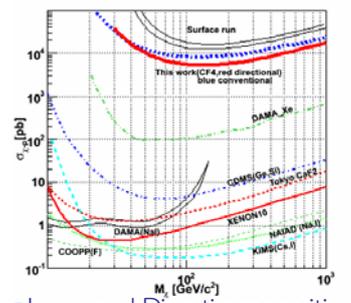
- Since 2007
- Kamioka mine (2700 m.w.e.) (N 36.25 E 137.18)



- After several pilot runs
- Kamioka RUN-5: First Dark Matter run  $\sim$ 0.52 kgdays



### SD 90% WIMP cross section limits



- Improved Direction-sensitive limits
- but..
- Still worse than conventional methods
- Internal radioactive BG restricts the sensitivities
- We are working on to reduce the backgrounds!

### 6. Plans

- Low BG detector by material selection
- Larger volume ( $\sim$  1m<sup>3</sup>)

