



Sep 9th, 2019
TAUP 2019

Toyama, JAPAN



C  GNUS

Kentaro Miuchi
(Kobe University)

- Overview
- Activities
- Highlights
- Summary

科研費
KAKENHI



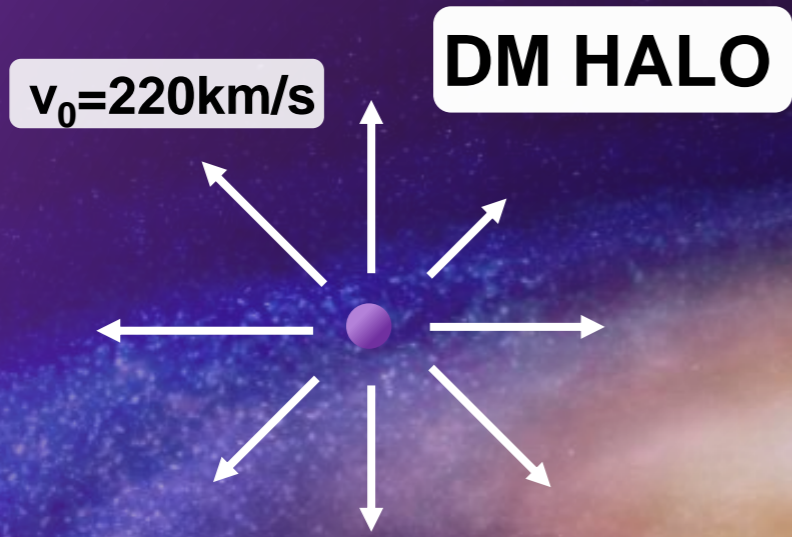
Overview

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“CYGNUS” concept

G. C.

WIMP-wind detection



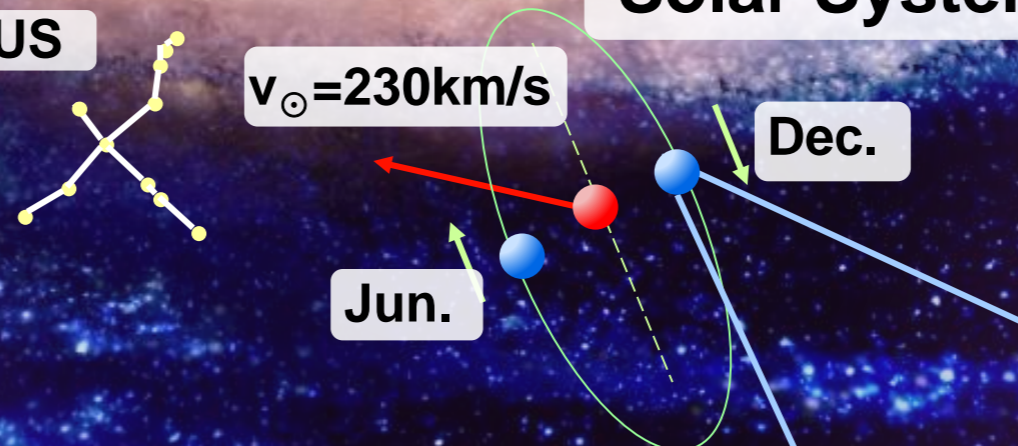
CYGNUS

Solar System

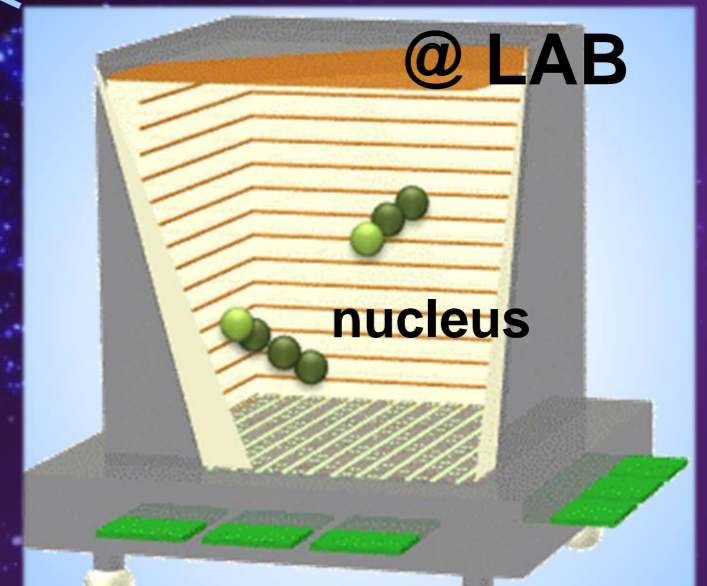
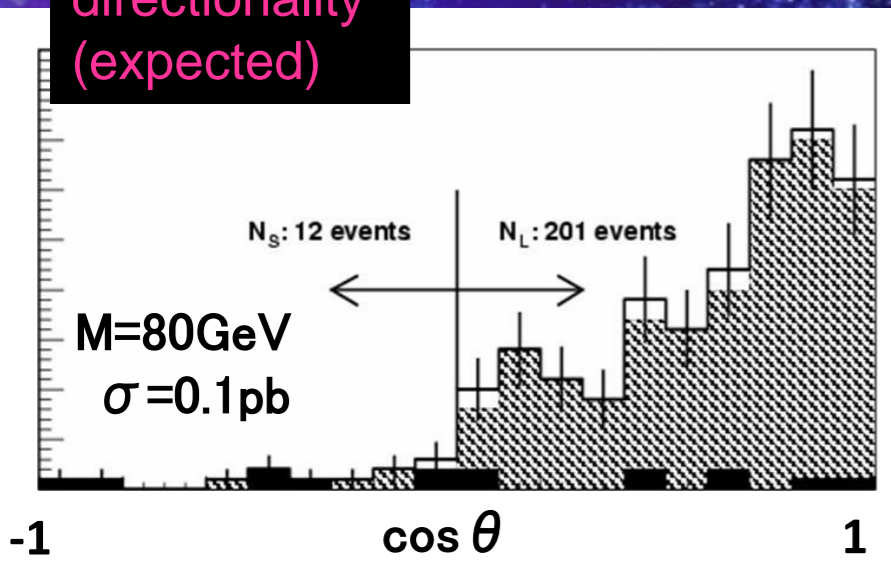
$v_{\odot} = 230 \text{ km/s}$

Dec.

Jun.

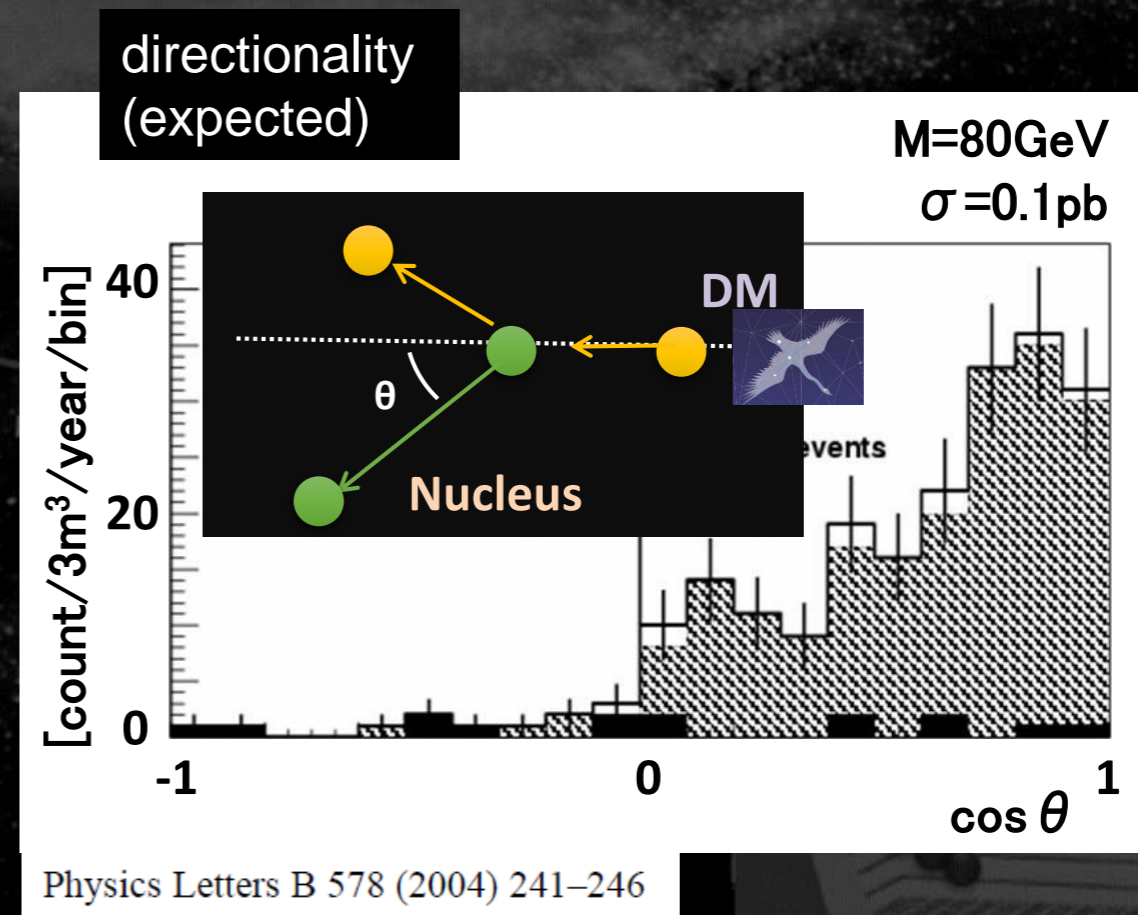
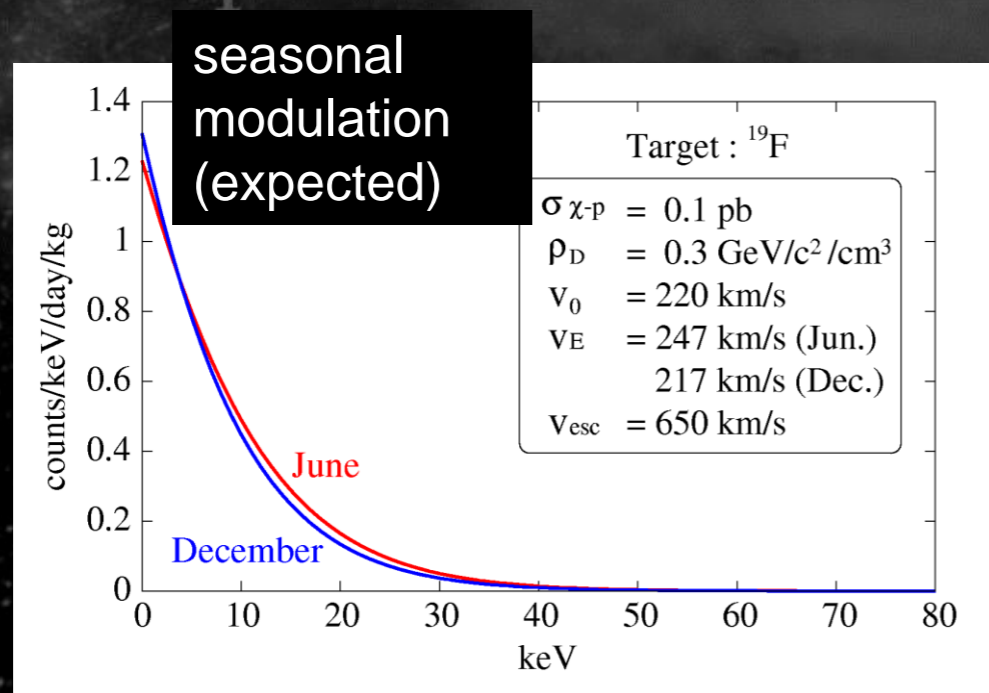


directionality (expected)



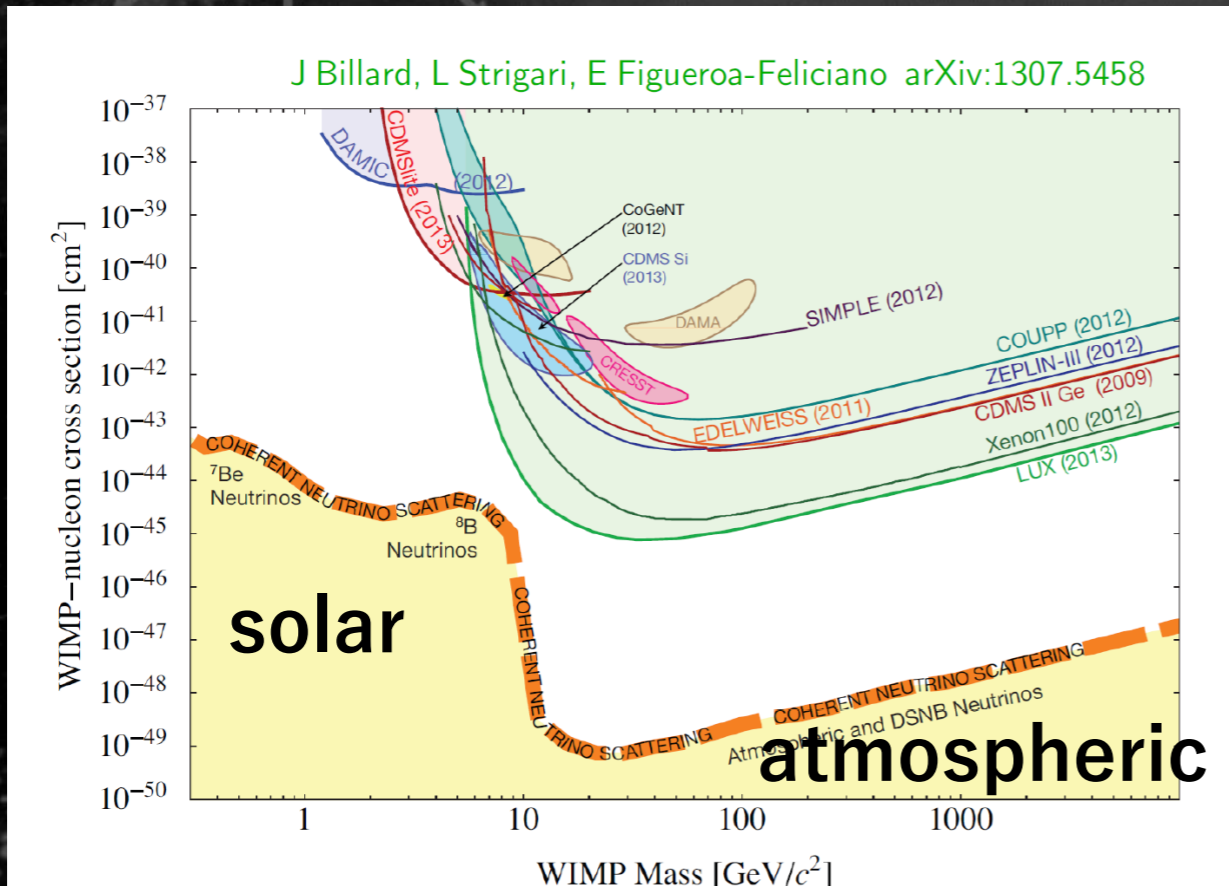
CYGNUS: Directional Detection

- Clear Discovery
+ study the nature of DM after discovery

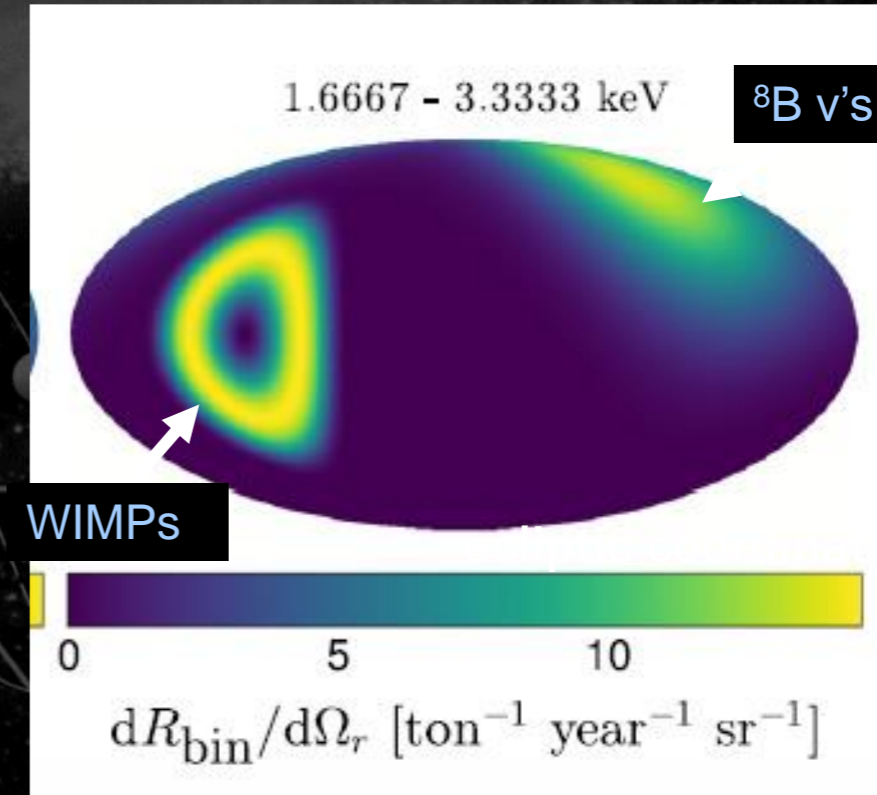


Toward discovery

- Potential to search beyond the “neutrino floor” where large detectors are reaching.



F. Mayet et al. / Physics Reports 627 (2016) 1–49



- distinguishable

CYGNUS: community

- 7 × bi-annual workshops (2007-)

- CYGNUS 2017 Xichang, Sichuan, China June 13 - 16, 2017
- CYGNUS 2015 Occidental College, Los Angeles, California, USA June 2 - 4, 2015.
- CYGNUS 2013 Toyama, Japan June 10 - 12, 2013.
- CYGNUS 2011 Aussois, France June 7 - 10, 2011.
- CYGNUS 2009 Massachusetts Institute of Technology, Cambridge, Massachusetts, USA June 11 - 13, 2009.
- CYGNUS 2007 Boulby Underground Laboratory, Saltburn-by-the-Sea, Cleveland, UK July 22 - 24, 2007.

- 2 × review papers, another is coming



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THE CASE FOR A
DIRECTIONAL DARK MATTER DETECTOR AND
THE STATUS OF CURRENT EXPERIMENTAL EFFORTS

Readout technologies for directional WIMP Dark Matter
detection

Physics Reports 662 (2016) 1–46

J.B.R. Battat^{1,*}, I.G. Irastorza², A. Aleksandrov³,
E. Baracchini⁶, J. Billard^{7,8}, G. Bosson⁷, O. Bourrion⁷, J. Bouvier⁷,
A. Buonaura^{3,9}, K. Burdge^{10,11}, S. Cebrián², P. Colas¹², L. Consiglio¹³, T. Dafni²,
N. D'Ambrosio¹³, C. Decanay^{10,14}, G. De Lellis^{3,9}, T. Decombes⁷

CYGNUS: collaboration

- proto-collaboration (2016-)
 - >50 researchers
 - discussion on-going for actual collaboration



The CYGNUS Galactic Directional Recoil Observatory Proto-Collaboration Agreement

Now that conventional WIMP dark matter searches are approaching the neutrino floor, there has been a resurgence of interest in the possibility of introducing recoil direction sensitivity into the field. Such directional sensitivity would offer the powerful prospect of reaching below this floor, introducing both the possibility of identifying a clear signature for dark matter particles in the galaxy below this level but also of exploiting observation of coherent neutrino scattering from the Sun and other sources with directional sensitivity. There has also been significant progress recently in development of technology able to record the directional information from nuclear recoils at low energy (sub-100 keV) necessary for these goals. This includes progress on improving the sensitivity of low pressure gas time projection chamber technology but also on novel ideas with higher density targets, such as ultra-fine grain emulsions, scintillation materials, columnar recombination with noble gas targets and concepts using nano technology. Such world wide directional expertise, if pooled together and directed

steering committee

E. Baracchini (GSSI)
G. Lane (ANU, Canberra)
K. Miuchi (Kobe)
N. Spooner (Sheffield)
S. Vahsen (Hawaii)

Activities

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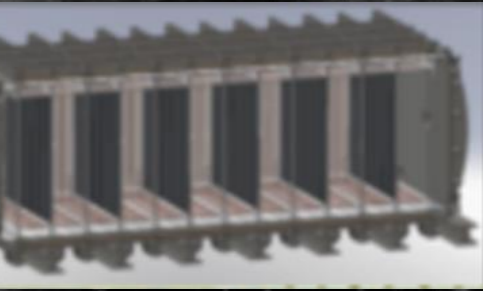
World-wide CYGNUS (ver. TAUP2019)

CYGNUS-10
Boulby, UK
10m³ He:SF₆
GEM + wire readout

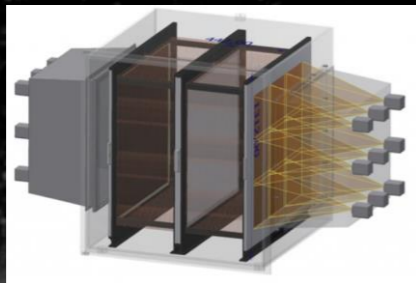


CYGNUS-KM
Kamioka, Japan
SF₆ / CF₄
Strip readout

40cm

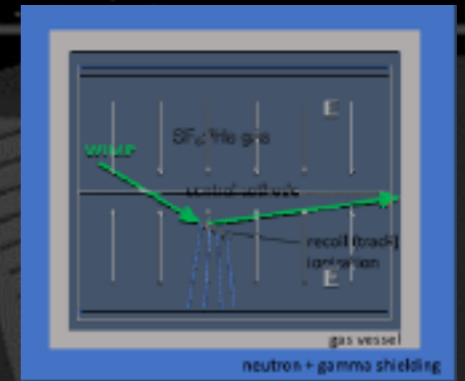


CYGNUS-Initium
Gran Sasso, Italy
He CF₄ (SF₆)
sCMOS+PMT readout



CYGNUS-OZ
Stawell, Australia
R&D leading to 1 m³
Long-term plan 10 m³

CYGNUS-HD10
SURF, USA
He:CF₄:C₄H₁₀
Strip readout



multi-site observatory

• UK / Boulby

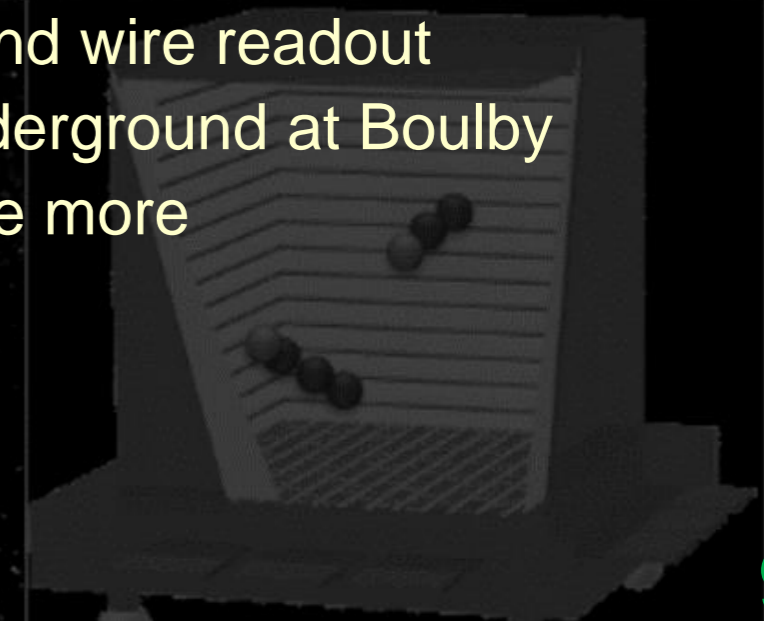
- pioneered this field (DRIFT)
- 1m³ detector running underground (Boulby) for years
- low BG, large volume



Boulby Underground Lab

• 10m³ chamber design ongoing

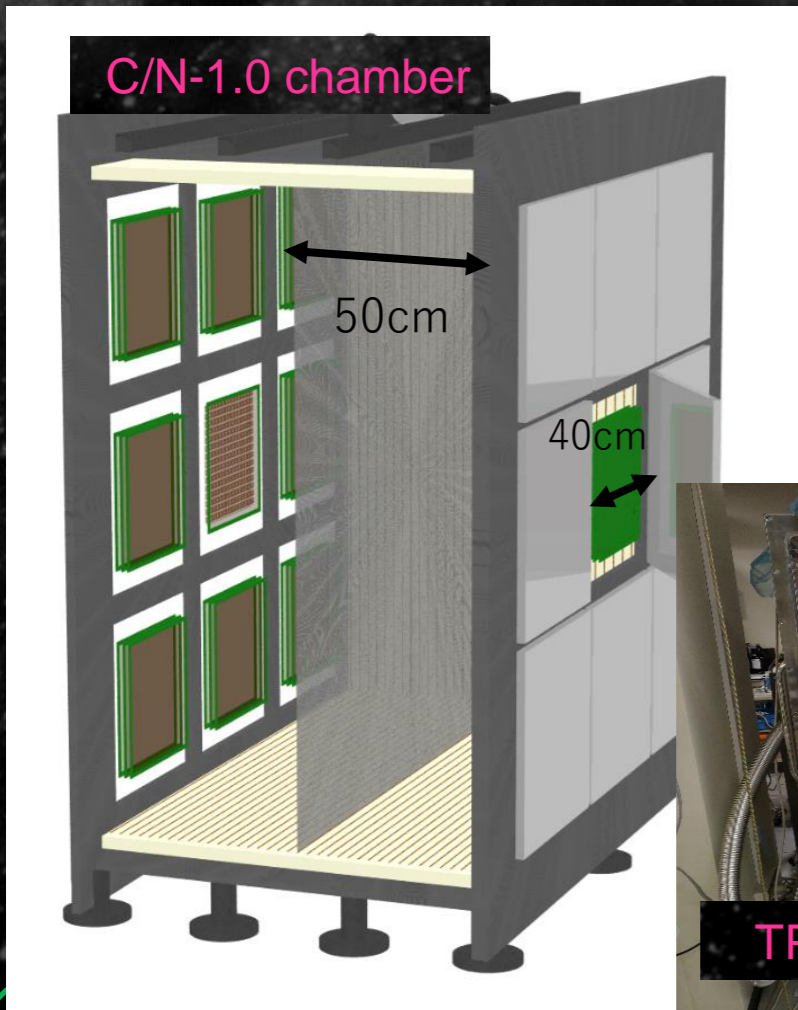
- low BG vessel design w/ simulation
- R&D for GEM and wire readout
- clean space underground at Boulby
- easy to excavate more



• JAPAN / Kamioka

See T.Ikeda's Talk for NEWAGE

- pioneered 3d-tracking (direction sensitive) (NEWAGE)
- C/N-1.0 chamber (18 × 30 × 30 cm² detectors)
 - chamber ready
 - TPC cage (w/ resistive sheet), feedthrough being commissioned



• Negative ion studies

- 3-D tracking
- MPGD gas avalanche simulation



• ASICs for negative ion strip readout

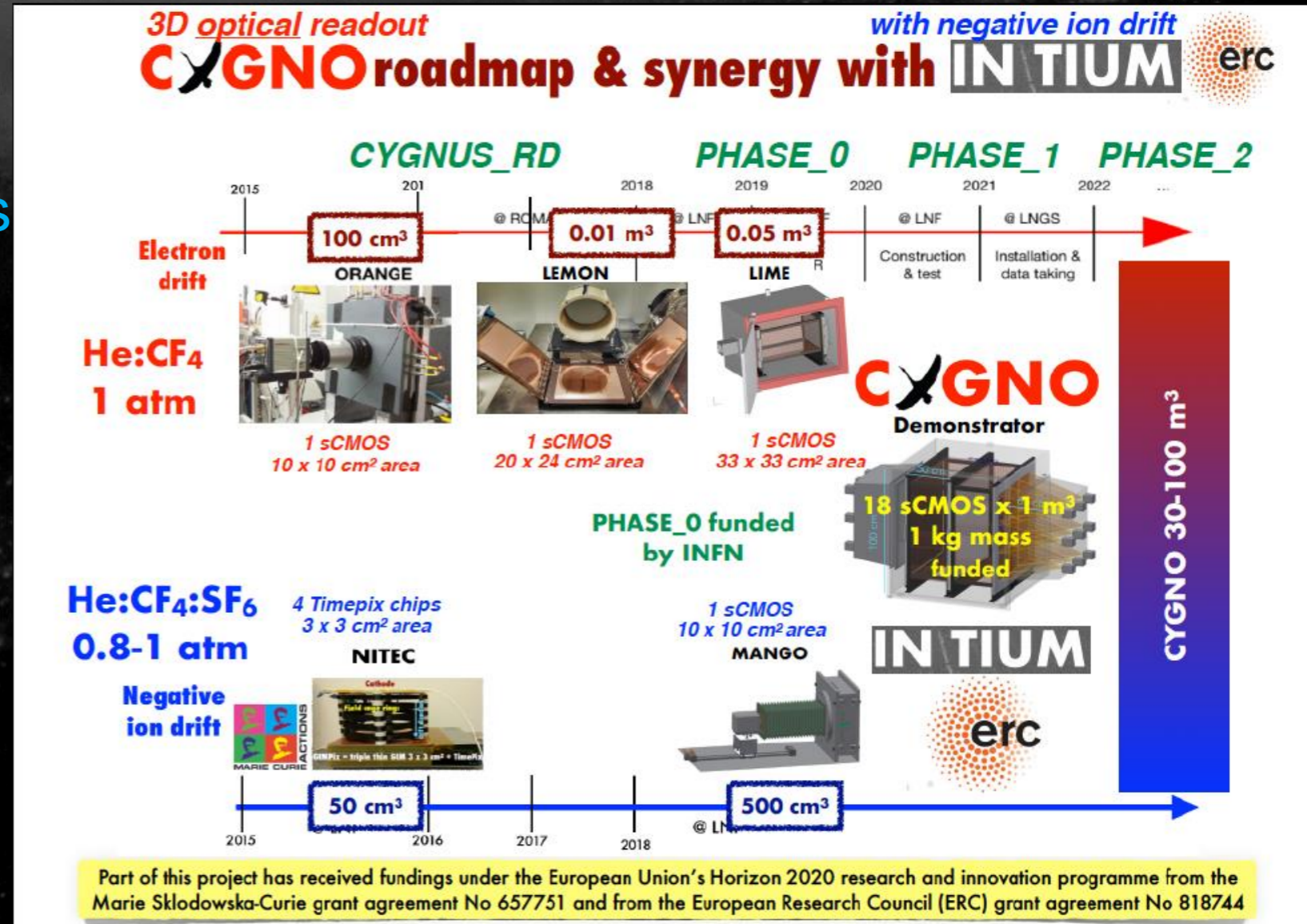
- > 5k channels made
- chip test started



Italy / GranSasso (intended)

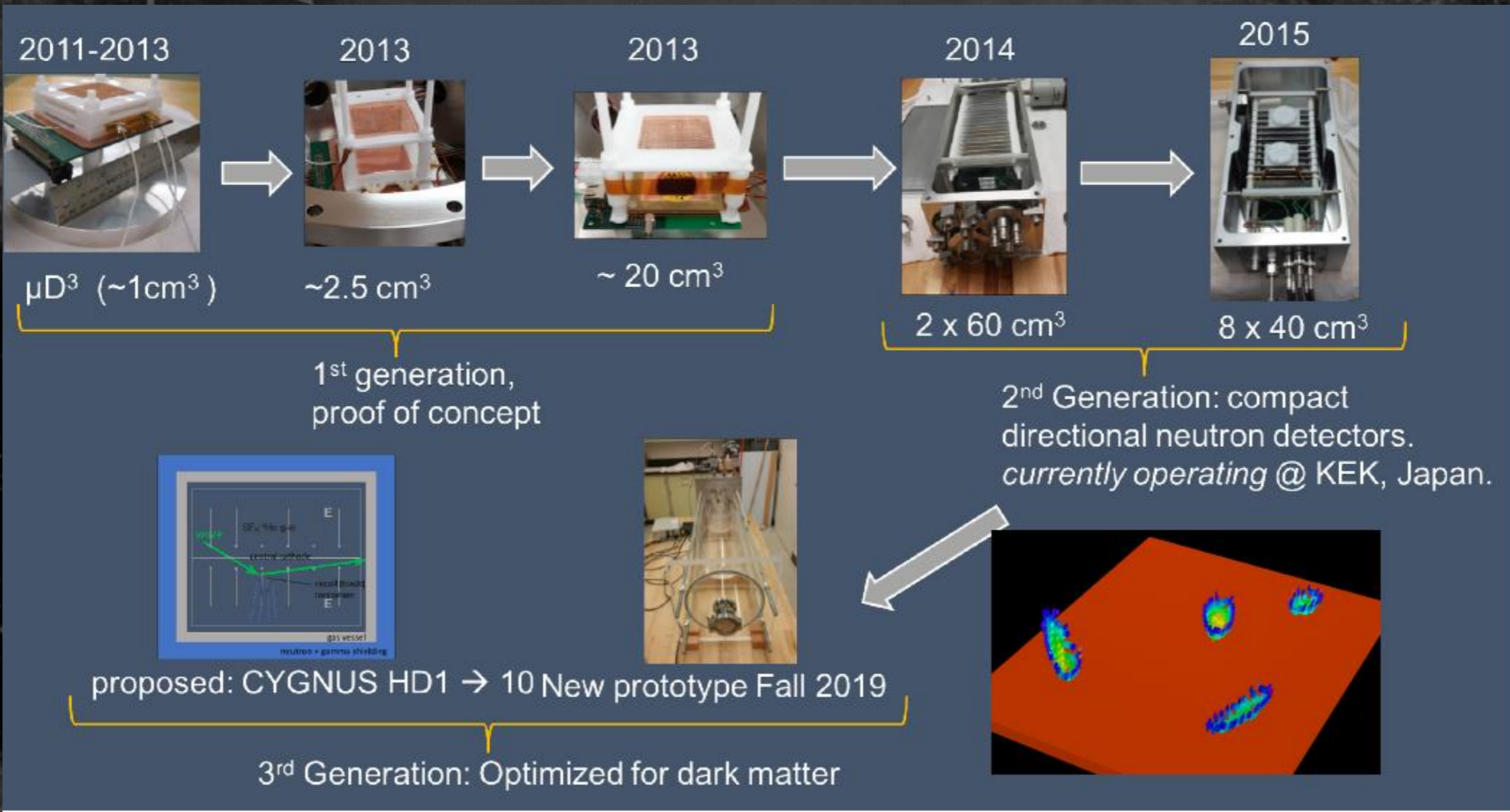
- Focusing optical readout
- Two parallel R&D paths
 - electron drift
 - negative ion drift
- 1m³ scale detector funded as demonstrator for 30-50m³

See E.Barracchini's Talk



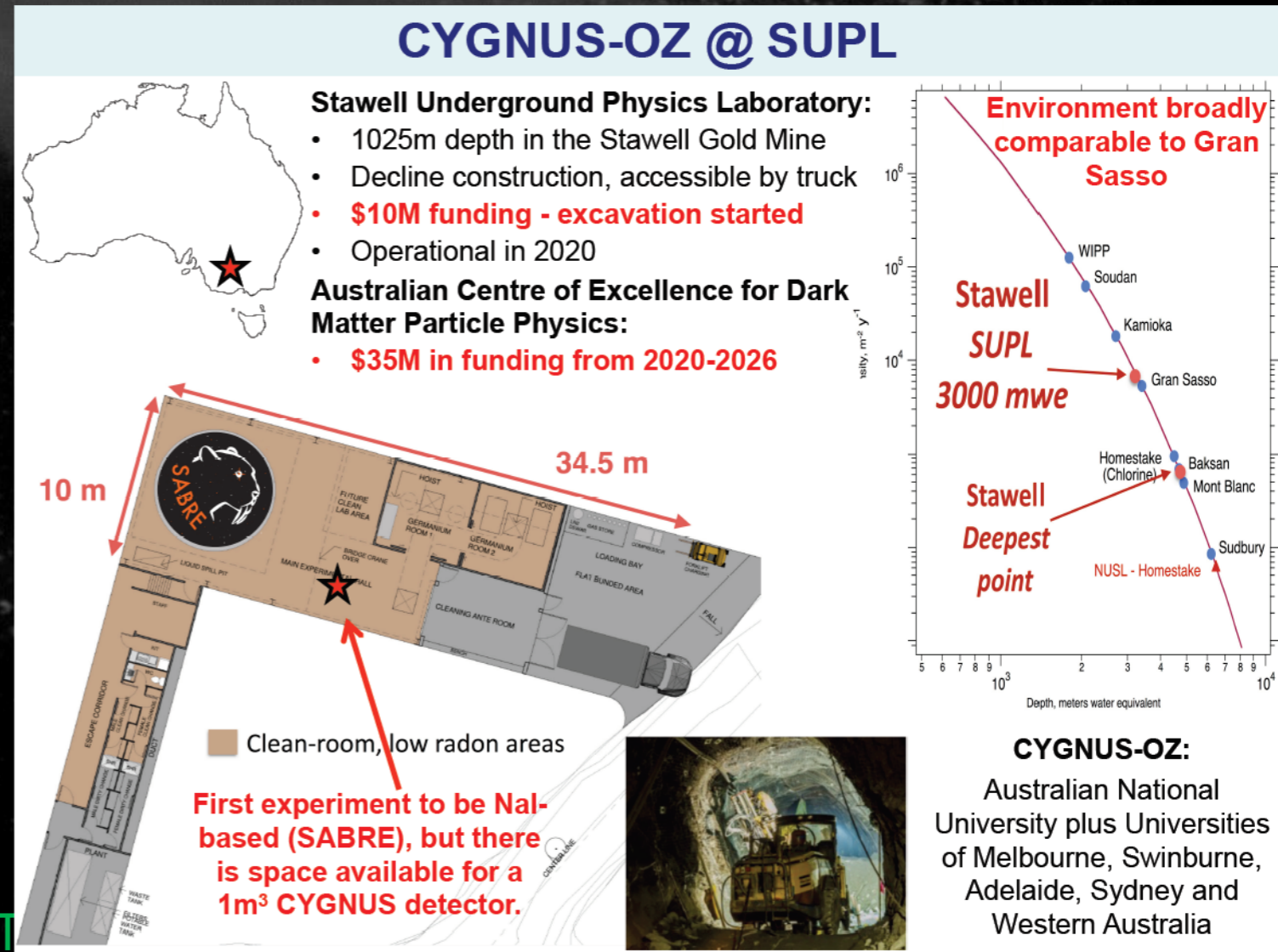
• US / SURF(intended)

- Focusing on pixel, strip readout (HD)
- Extensive prototyping completed
- CYGNUS HD1 1-m³, demonstrator for 10 m³, proposed



• Australia / Stawell

- Excavation of new lab started - operation in 2020
- Space available in 2020 for 1 m³ CYGNUS TPC, 10 m³ in 2025?
- DM community recently funded - includes R&D for CYGNUS





Highlights



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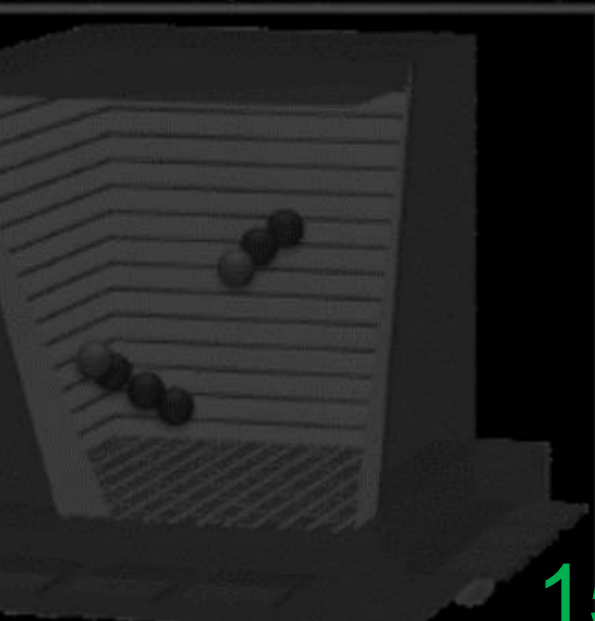
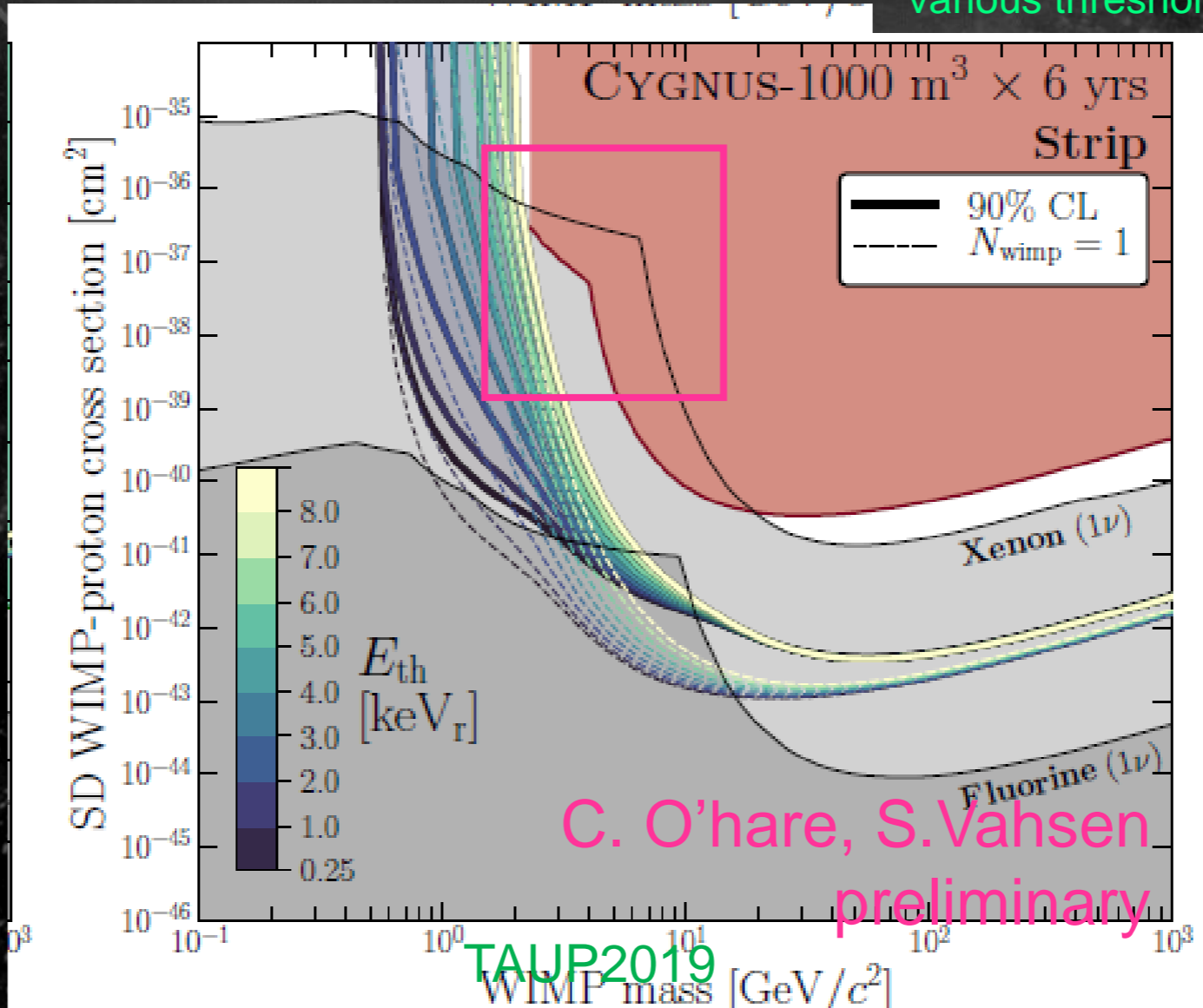
Highlight 1: Feasibility Study

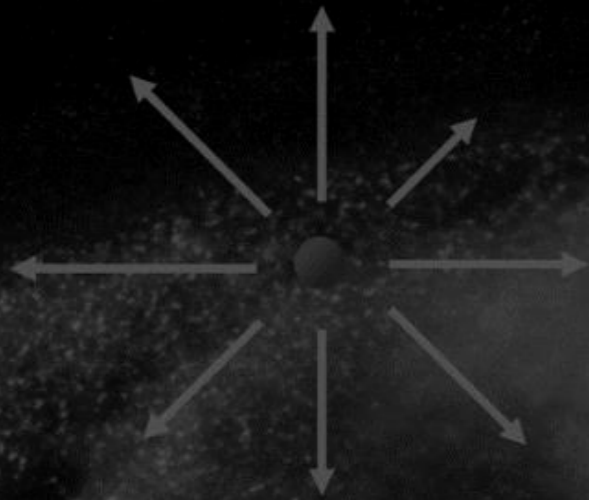
Paper in internal review

- Realistic simulation (strip readout)

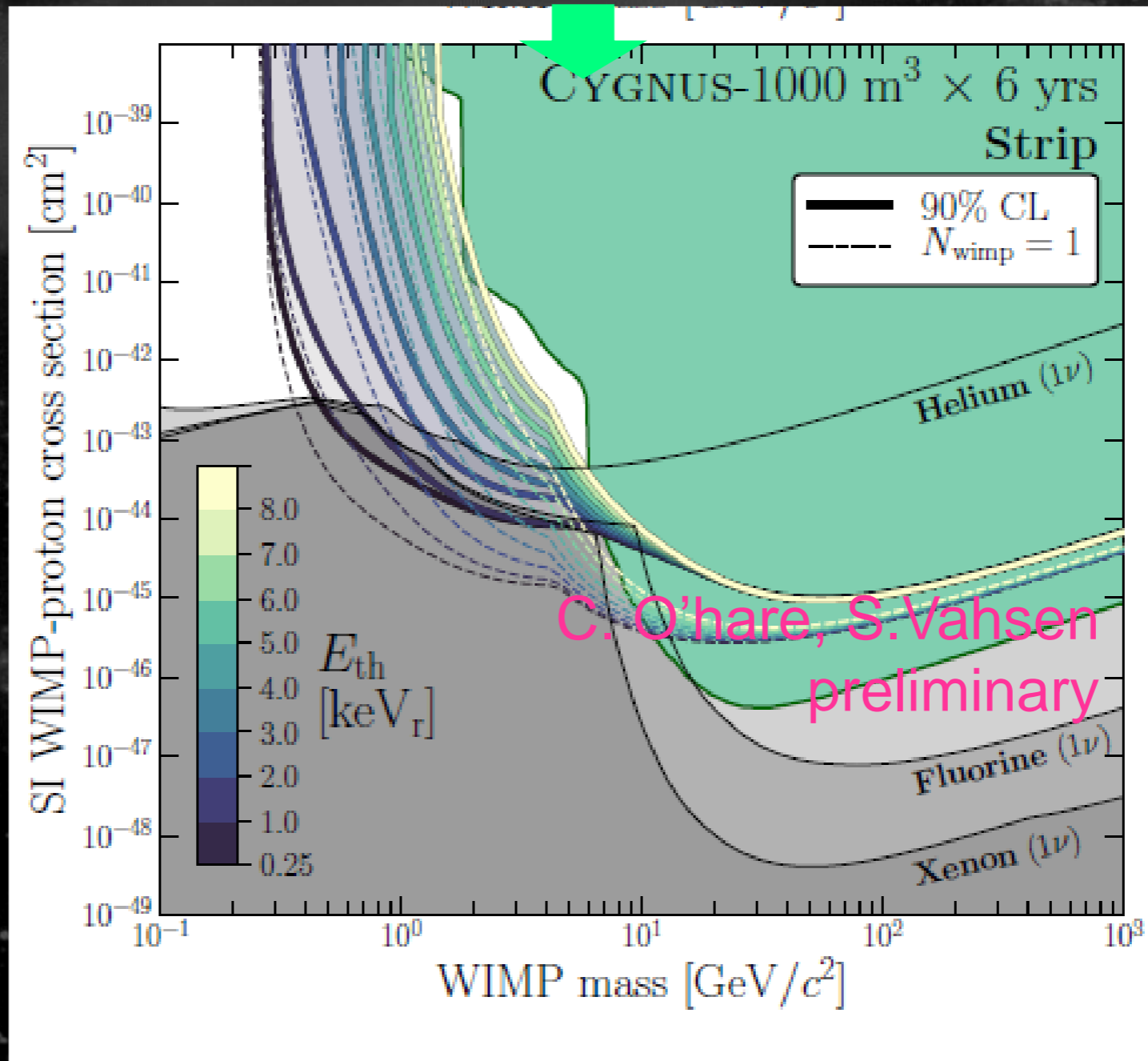
even 10m³ detector
(3 order magnitude higher than
the shown curves) can start
exploring Xe neutrino floor

1000m³
strip readout with
various threshold



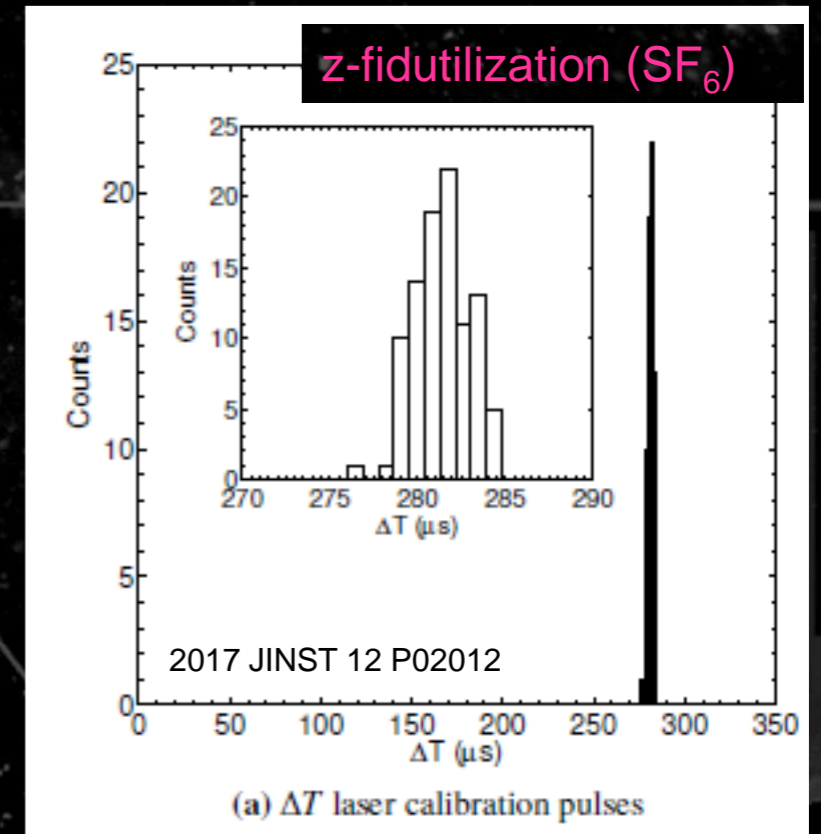
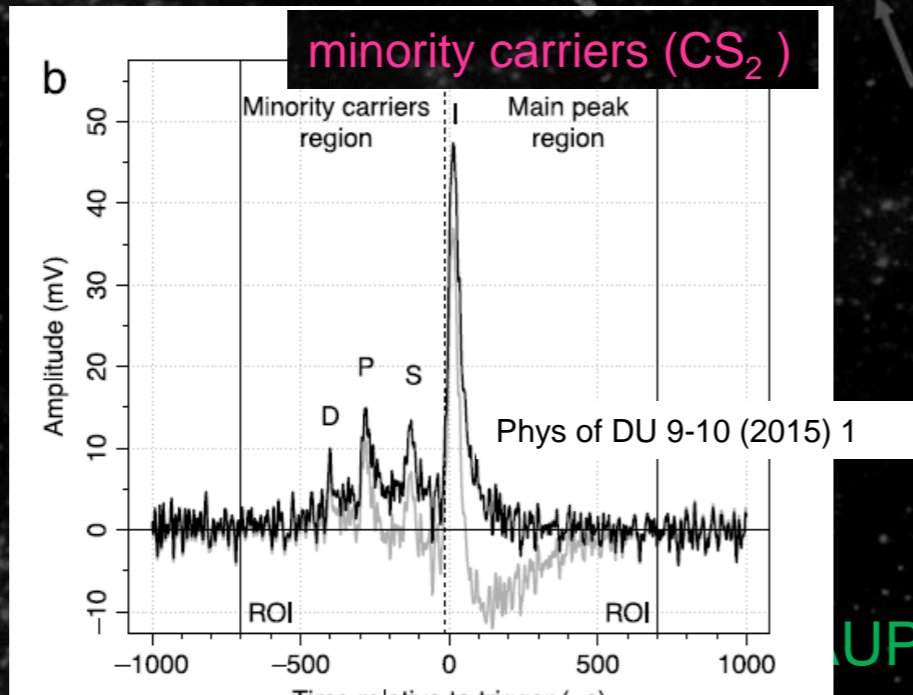


strip readout with various threshold



Highlight 2: Negative ion TPC Study

- Pioneered by DRIFT group small diffusion
- Minority carrier discovery ($\text{CS}_2 + \text{O}_2$, Occidental group)
 - use several ion species with different drift velocities
 - ⇒ z fiducialization possible ⇒ LOW BG !
- SF_6 discovery (2015, UNM group).
 - z-fiducialization 7.3mm FWHM

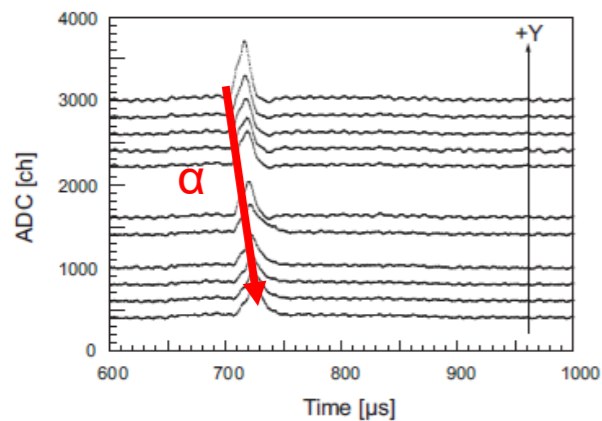
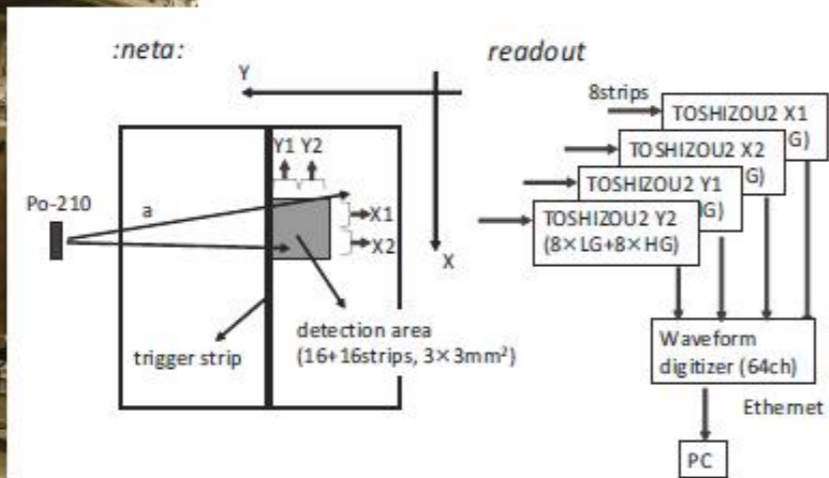
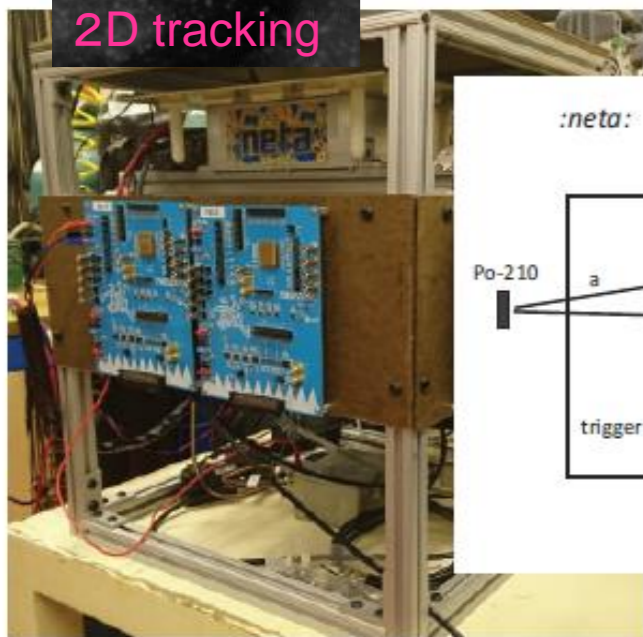


• to be CYGNUS: Trackings

- strip readout + ASICs

LTARS2016 + Wellesley's micromegas resistive-strip readout

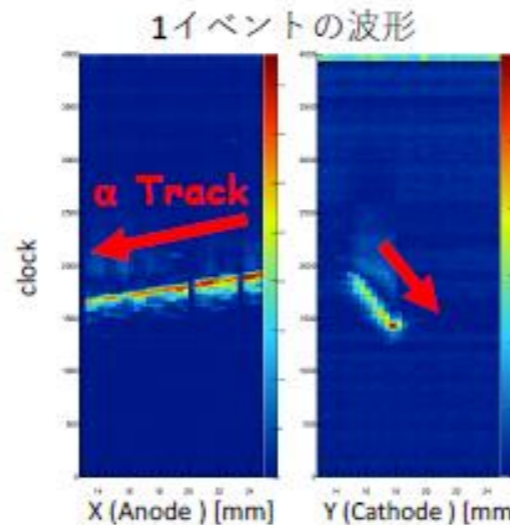
2D tracking



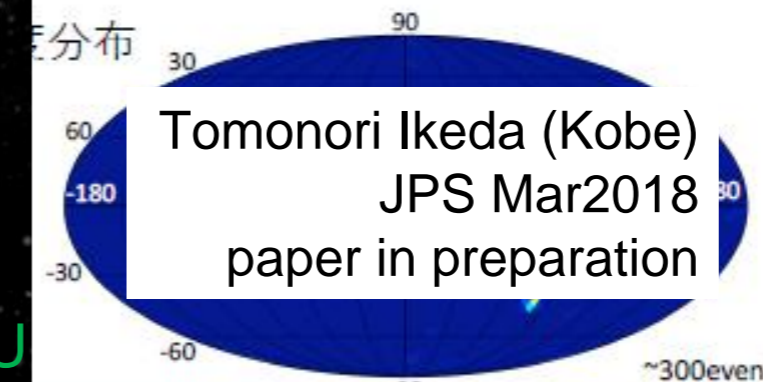
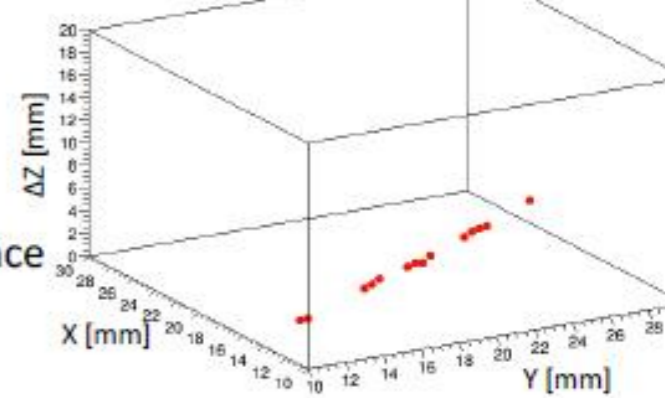
2019 J. Inst. 14 T01008

for optical readout: See E.Barracchini's Talk

3D tracking+ fiducialisation



coincidence

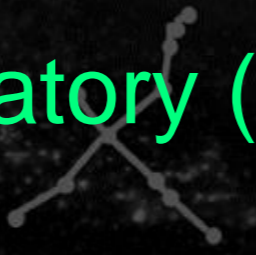


Ke

TAU

Summary



- CYGNUS: direction sensitive DM direct search
 - community, collaboration
 - multi-site observatory ($1\text{m}^3 \Rightarrow$ larger scale detectors)
 - New comers (physics, detectors...) are welcome!
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